

THE EVOLUTION OF MAUNA KEA VOLCANO, HAWAII:  
PETROGENESIS OF THOLEIITIC AND ALKALIC BASALTS

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**Abstract.** Mauna Kea Volcano has three exposed rock units. Submarine shield-building tholeiites form the oldest unit. Subaerial, interbedded tholeiitic and alkalic basalts form the intermediate age unit (70-240 Ka), and they are partially covered by evolved alkalic lavas, hawaiites and mugearites (4-66 Ka). In contrast to other Hawaiian volcanoes, such as Haleakala and Kauai, lavas from Mauna Kea do not define systematic temporal variations in Pb, Sr or Nd isotopic ratios. However with decreasing age the tholeiitic basalts are increasingly enriched in incompatible elements; therefore the shield and postshield tholeiites were derived from compositionally distinct parental magmas. Submarine shield lavas from the east rift contain forsterite-rich olivine (up to Fo<sub>90.5</sub>) providing evidence for MgO-rich (14.4 to 17%) magmas. Postshield tholeiitic and alkalic basalts with similar isotopic ratios may have been derived from the same source composition by different degrees of partial melting. If a compositionally and isotopically homogeneous source and a batch melting model are assumed, inversion of incompatible element abundance data for the postshield basalts requires low degrees (<2%) of melting of a garnet lherzolite source which had near-chondritic abundances of heavy rare-earth elements (REE) but less than chondritic abundances of highly incompatible elements such as Ba, Nb and light REE. As the volcano migrated away from the hotspot, eruption rates decreased enabling high Fe-Ti basalts to form by fractional crystallization in shallow crustal magma chambers. The associated phenocryst-rich, high-MgO postshield lavas (picrites and ankaramites) are products of phenocryst accumulation. Eventually basaltic eruptions ceased, and the youngest Mauna Kea lavas are exclusively hawaiites and mugearites which formed from alkalic basalt parental magmas by clinopyroxene-dominated fractionation at lower crustal pressures.

### 1. Introduction

Hawaiian volcanoes evolve through well-defined stages: preshield, shield, postshield, and posterosional [e.g., Clague and Dalrymple, 1987]. The shield stage forms the bulk (95-98 vol. %) of each volcano, and it is dominantly composed of tholeiitic basalt. However, eruption of alkalic basalt precedes and postdates formation of the shield [e.g., Clague, 1987]. This compositional variation correlates with magma supply rates [e.g., Feigenson and Spera, 1981] which apparently are initially low but increase to high rates (100 x 10<sup>6</sup> m<sup>3</sup>/yr

[Dzurisin et al., 1984]) during shield formation and then decrease sharply during the postshield stage (10 x 10<sup>6</sup> to 4 x 10<sup>4</sup> m<sup>3</sup>/yr [Moore et al., 1987; Spengler and Garcia, 1988; Frey et al., 1990]). Magma supply rate is presumed to be controlled by the location of the volcano relative to the focus of melting, i.e., "the hotspot". In order to understand the processes controlling the compositional variations of basalts during the transition from shield to postshield volcanism, we selected Mauna Kea Volcano for a detailed mapping and geochemical study. Mauna Kea Volcano is the youngest of the Hawaiian volcanoes which has essentially completed its postshield volcanism.

Previously, we discussed the evolution and origin of the evolved lavas (hawaiites and mugearites) which form the postshield hawaiitic substage at Mauna Kea volcano [West et al., 1988; Frey et al., 1990]. In this paper, we focus on the oldest subaerial lavas exposed on Mauna Kea (Figure 1). These lavas include tholeiitic and alkalic basalts, ankaramites and picrites, and they form the postshield basaltic substage [Frey et al., 1990]. This substage formed during the transition from tholeiitic to alkalic volcanism. Intercalation of tholeiitic and alkalic basalt is common during the transition from shield to postshield volcanism at Hawaiian volcanoes [e.g., Macdonald and Katsura, 1964; Feigenson et al., 1983; Chen and Frey, 1985; Lanphere and Frey, 1987; Clague, 1987]. However, no previous study has documented and explained the origin of the diverse basalt types found in the postshield lava sequences. In order to better define temporal geochemical variations, we also studied submarine shield lavas dredged from the east rift of Mauna Kea (Figure 1). These lavas are olivine-phyric, MgO-rich tholeiites. Our results provide the first Sr isotopic and trace element abundance data for shield basalts from Mauna Kea Volcano.

The objectives of this paper are to (1) document the temporal geochemical trends defined by basaltic lavas erupted during the late shield and postshield stages of Mauna Kea Volcano; and (2) understand the processes that created the diversity of basalt compositions.

### 2. Description Of Sampling Locations

Mauna Kea, 4205 m above sea level, is the tallest of the five shield volcanoes comprising the island of Hawaii. Most of the volcano is submarine (>90%) and its volume is estimated to be 40 x 10<sup>3</sup> km<sup>3</sup> [Wolfe et al., 1991]. A pronounced break in slope at 400-500 m below sea level on the east side of the volcano is interpreted to represent sea level at approximately the time when the frequent eruptions of the shield stage ended [Moore and Campbell, 1987]. Shield-building flows were erupted from the summit and along two rift zones. The east rift zone forms a prominent submarine ridge (Figure 1); the west rift is more subdued. Postshield lavas, erupted from vents scattered over the surface of the volcano, mantle the entire subaerial portion of the volcano (Figure 1). Stratigraphically, these lavas have been divided into two units: a lower basaltic sequence designated as the Hamakua Volcanics and an upper hawaiite to mugearite sequence called the Laupahoehoe Volcanics [Stearns and Macdonald, 1946; Langenheimer and Clague, 1987]. The Hamakua lavas range in age from 240 to 70 kyr; the Laupahoehoe lavas from 66 to 4 kyr [Porter, 1979; Wolfe et al., 1991].

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Paper number 91JB00940.  
0148-0227/91/91JB-00940\$05.00

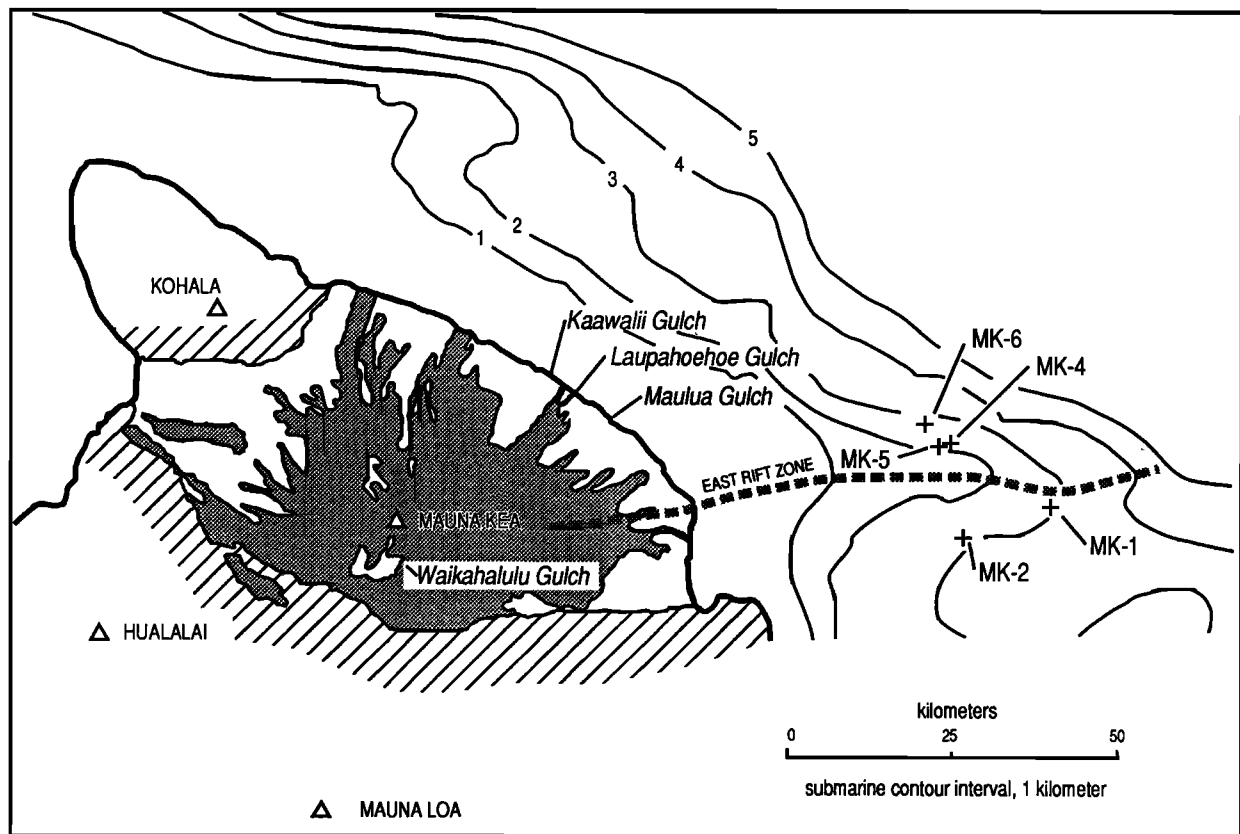


Fig. 1. Geologic and bathymetric map of Mauna Kea volcano, Hawaii. Postshield lavas cover the entire subaerial surface of the volcano; the unshaded area consists of basalts, while the shaded area is mostly hawaiite with some mugearite. The locations of four canyons (gulches) where the stratigraphic sections of the postshield basaltic lavas were collected are labeled. Also shown are locations of dredged samples from the east rift zone of the shield stage.

Stream erosion, probably accompanying glaciation at the summit of the volcano, produced deep gulches on the windward coast (east flank) of Mauna Kea. In three of these gulches, stratigraphic sections 90 to 165 m thick were collected (Figure 2). Maulua Gulch provides the deepest exposure into the volcano. Approximately 200 m of section is present but only a 165 m section could be sampled. The Laupahoehoe area was sampled by Macdonald and Katsura [1964] on the south side of the gulch (~50 m section), but the road used then is now poorly exposed. A new composite section of 100 m was collected beginning at sea level and continuing along a new road on the north side of the gulch (Figure 2). Samples from the new and old Laupahoehoe Gulch sections correspond closely in rock type and lava compositions. Sample LP-7 has been K-Ar dated at  $240 \pm 30$  Ka [Wolfe et al., 1991]. Kaawalii Gulch was collected along the coastal road, along the sea cliffs and in the gulch above the road to yield a section 90 m thick. Each section is covered with 1.5 to 2 m of ash which was mostly derived from Laupahoehoe eruptions [Macdonald, 1949]. All of the lavas in these sections are from the Hamakua Volcanics. Basalts (<5 vol. % phenocrysts) are the most common and voluminous members of the sections. In order of decreasing volume there are also ankaramites (>20 vol. % total phenocrysts with subequal amounts of clinopyroxene and olivine phenocrysts), olivine basalts (5-15 vol. % olivine phenocrysts), high Fe-Ti basalts (nearly aphyric with >4 wt. %  $\text{TiO}_2$ ), picritic basalts (>15 vol. % olivine phenocrysts) and plagioclase basalts (>5 vol. % plagioclase).

The dredged lavas are pillow basalts recovered from 1.9 to 3.0 km water depths on the east rift. They are inferred to be

shield lavas because they were collected at depths well below the ~500 m break in slope that marks the end of the shield-building stage (Figure 1). Major element analyses, volatile contents and oxygen isotopic data for pillow rim glasses from these lavas were discussed by Garcia et al. [1989]. Based on the compositional diversity of these glasses, eight whole-rock samples were selected for geochemical and petrologic study. All are olivine phryic, 5-22 vol.%, and some have rare, <1 vol.%, plagioclase and clinopyroxene phenocrysts.

### 3. Results

#### 3.1. Whole Rock Major Element Compositions

The dredged shield lavas are MgO-rich tholeiites (9.4 to 16.8%, Table 1). In a  $(\text{Na}_2\text{O}+\text{K}_2\text{O})$  versus  $\text{SiO}_2$  plot they are well within the tholeiitic field overlapping with recent tholeiitic basalts erupted from Kilauea Volcano (Figure 3a). Previously, Moore [1966] studied three lavas dredged from the east rift of Mauna Kea. These samples are also MgO-rich (~13 to 20% MgO), but two are transitional between tholeiitic and alkalic compositions. The postshield Hamakua lavas exposed in three east flank gulches have a wide range in composition (Table 2) varying from tholeiitic to alkalic without a compositional break between these basalt types (Figure 3a). The oldest Hamakua flows occur in Maulua Gulch, and they are tholeiitic (Table 2b, MU-13 through MU-4 in Figure 2). However, except for sample MU-8, they lie closer to the alkalic-tholeiitic boundary than the dredged shield lavas (Figure 3a). A stratigraphic suite of Hamakua lavas which directly underlie the Laupahoehoe Volcanics in Waikahalulu

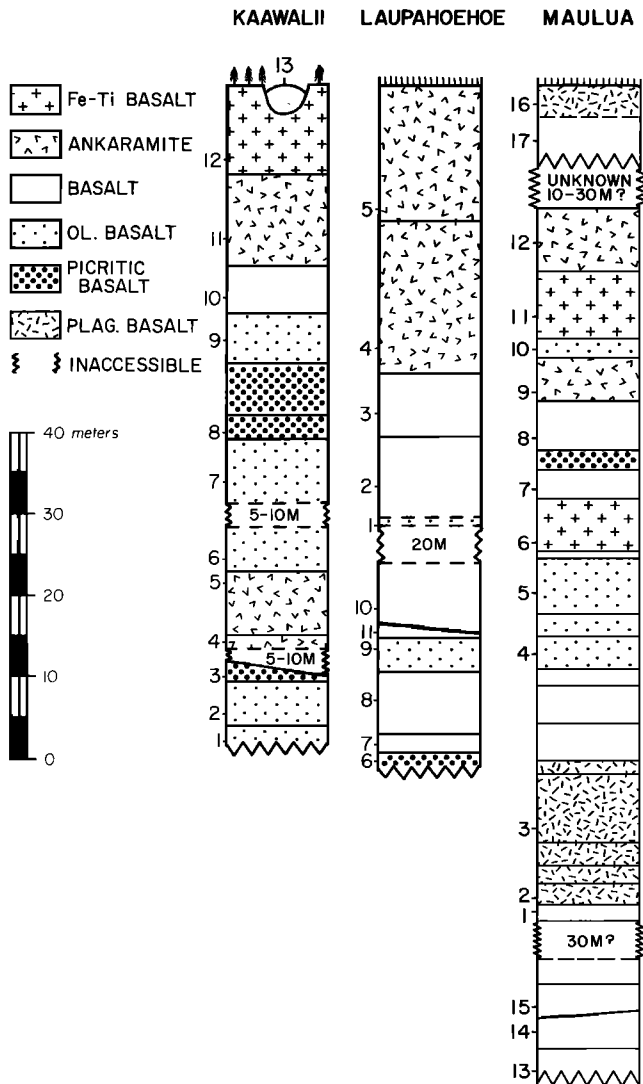


Fig. 2. Composite stratigraphic columns for Hamakua basalts collected in three gulches along the northeast flank (Hamakua Coast) of Mauna Kea volcano (Figure 1). Numbers indicate sample locations. Inaccessible sections are not shown to scale, but their probable thicknesses are indicated. The top of the sections are at approximately the same stratigraphic level; thus the Maulua section contains the oldest lavas. Rock types are defined in the text.

Gulch on the upper south flank (see Figure 4 of Frey et al. [1990]) also includes tholeiitic and alkalic lavas (Figure 3a). In this gulch the six stratigraphically lowest samples plot within the tholeiitic field. They contain ~48 to 49% SiO<sub>2</sub>, 5-10.6% MgO and less than 3% (Na<sub>2</sub>O+K<sub>2</sub>O) (Figure 3a, this paper; Table 1b in Frey et al. [1990]).

Many of the Hamakua lavas are phenocryst-rich (>15 vol.%); i.e., picrites and ankaramites (Table 3). These samples plot in the tholeiitic field at relatively low SiO<sub>2</sub> contents and low Al<sub>2</sub>O<sub>3</sub>/CaO ratios (Figure 3). Most of the 16 tholeiitic Hamakua lavas with <15% phenocrysts have relatively low MgO content (5.9 to 6.8%), and phenocrysts of plagioclase, olivine and clinopyroxene (Tables 2 and 3). MgO variation diagrams for all the Hamakua data in Table 2 are in Figure 8 of Frey et al. [1990]). Sample MU-8 is nearly aphyric (Table 3), and it has high MgO, and Ni contents (10.9% and 239 ppm, respectively). Only phenocryst-rich samples have higher MgO; thus, MU-8 may be the best

example of a parental tholeiitic melt. Among the alkalic Hamakua basalts with less than 15% phenocrysts, the most primitive lavas (9.6 to 10.2% MgO, ~215 ppm Ni) are KI-1, KI-2 and LP-9 which are from the lower parts of these gulches (Figure 2, Table 2). The most evolved Hamakua lavas are aphyric, high Fe-Ti basalts (e.g., MU-6, MU-11 and KI-12 in Table 2) which have high Al<sub>2</sub>O<sub>3</sub>/CaO (1.6 to 1.7) coupled with high (Na<sub>2</sub>O+K<sub>2</sub>O) and relatively low SiO<sub>2</sub> contents (Figure 3).

The tholeiitic basalts dredged from the shield have K<sub>2</sub>O/P<sub>2</sub>O<sub>5</sub> (1.56 to 1.85) typical of unaltered Hawaiian tholeiitic lavas. They contain fresh glass, unaltered olivine and the vesicles are free of alteration minerals. In contrast, the degree of postmagmatic alteration in the Hamakua lavas is highly variable and ranges from unaltered (i.e., secondary minerals are absent and H<sub>2</sub>O<sup>+</sup> <0.4%) to moderately altered (e.g., alteration rims on olivine, altered groundmass minerals, patches of altered glass and H<sub>2</sub>O<sup>+</sup> of 0.5 to 1.7%, Table 2). Because K<sub>2</sub>O/P<sub>2</sub>O<sub>5</sub> ratios are greater than 1.5 in unaltered Hawaiian lavas (e.g., see Wright [1971] for tholeiitic basalts; Clague et al. [1980] and Clague [1987] for alkalic basalts), the low K<sub>2</sub>O/P<sub>2</sub>O<sub>5</sub> (<1) in many of these Hamakua lavas reflects loss of K<sub>2</sub>O during subaerial weathering [e.g., Feigenson et al., 1983; Chen and Frey, 1985; Frey et al., 1990]. Specifically, ~35% of the Hamakua lavas from the high rainfall east flank of Mauna Kea have K<sub>2</sub>O/P<sub>2</sub>O<sub>5</sub> <1 (Table 2). In contrast, all samples of similar to slightly younger age from the relatively low rainfall west flank have K<sub>2</sub>O/P<sub>2</sub>O<sub>5</sub> >1 and H<sub>2</sub>O<sup>+</sup> contents generally less than 0.4 wt.% [Frey et al., 1990]. Clearly, a classification diagram based on abundances of alkali metals (Figure 3a) must be used with caution. For example, sample MU-10 lies in the tholeiitic field, but if the K<sub>2</sub>O/P<sub>2</sub>O<sub>5</sub> ratios is raised from the observed 0.7 to 2 the sample lies near the boundary line (Figure 3a). However, as discussed below, the abundances of relatively immobile incompatible elements in these lavas confirms that most (15 of 17) of the nonporphyritic samples plotting in the tholeiitic field have incompatible element abundance ratios distinct from lavas plotting in the alkalic field.

### 3.2. Whole Rock Trace Element Abundances

At comparable MgO contents the Mauna Kea shield lavas (except for sample 1-8) have lower abundances of highly incompatible elements than the postshield Hamakua lavas (Figure 4). Abundance ratios of incompatible elements clearly show the geochemical differences between the shield and postshield tholeiitic lavas; e.g., the shield lavas are less enriched in highly incompatible elements so they have lower La/Ce and La/Yb (Figures 5 and 6). Moreover, the Mauna Kea shield lavas have lower La/Yb and higher Zr/Nb than historic Kilauea lavas (Figure 6). The chondrite-normalized REE patterns for the shield samples are typical of Hawaiian shield tholeiites [e.g., Hofmann et al., 1984; Budahn and Schmitt, 1985]; i.e., they have (La/Ce)<sub>N</sub> ≤ 1, relatively flat patterns from La to Eu but steeply sloped patterns from Eu to Lu (Figure 5).

There are important geochemical differences among postshield tholeiitic basalts. Specifically, tholeiitic lavas from Waikahalulu Gulch on the south flank have relatively higher La/Yb and lower Zr/Nb ratios than tholeiitic lavas from the east flank (Figure 6). In addition seven of the 36 samples analyzed from the west flank of Mauna Kea are tholeiitic (see Figure 7 of Frey et al. [1990]), but their La/Yb and Zr/Nb ratios do not overlap with the tholeiitic samples from the east flank. For example, Zr/Nb in samples 1, 14, 15, 20, 21, 22 and 24 (Table 1 of Frey et al., 1990) ranges from 7.9 to 9.5, whereas tholeiitic lavas from the shield and east flank have Zr/Nb >10.5. Based on their location (Figure 1), we presume that the east rift submarine lavas are older than the subaerial lavas from the east flank. In turn, the subaerial tholeiitic lavas in

TABLE 1. Major (wt.%) and Trace Element (ppm) Contents of Tholeiitic Lavas Dredged From the Mauna Kea Shield

	1-3*	1-10*	1-8†	2-1	5-3‡	5-13	6-6‡	6-18	
SiO <sub>2</sub>	48.76	48.99	50.76	48.68	49.18	50.05	49.33	47.98	49.70
TiO <sub>2</sub>	2.11	2.14	2.70	2.01	1.99	2.14	1.97	1.76	2.06
Al <sub>2</sub> O <sub>3</sub>	10.10	10.27	11.34	10.86	11.96	12.97	11.67	10.28	12.28
Fe <sub>2</sub> O <sub>3</sub>	12.42	12.44	12.86	12.18	11.90	11.74	11.94	12.29	11.83
MnO	0.17	0.17	0.17	0.17	0.17	0.17	0.16	0.17	0.17
MgO	15.76	15.43	10.93	14.61	12.16	9.42	13.23	16.85	11.33
CaO	7.68	7.77	7.95	9.12	9.55	10.30	9.32	8.11	9.76
Na <sub>2</sub> O	1.86	1.90	2.29	1.69	2.19	2.19	1.92	1.69	2.04
K <sub>2</sub> O	0.42	0.43	0.63	0.32	0.32	0.35	0.31	0.28	0.33
P <sub>2</sub> O <sub>5</sub>	0.24	0.24	0.34	0.18	0.20	0.21	0.19	0.18	0.20
Total	99.52	99.78	99.97	99.82	99.62	99.54	100.04	99.59	99.70
Mg#(§)	73.6	73.2	65.2	72.5	69.2	63.8	70.9	75.1	67.8
Rb	7.0	6.8	11.1	5.2	4.8	5.6	4.9	4.6	4.9
Sr	241	249	273	249	257	276	260	225	274
Ba	90	87	133	76	70	67	63	61	63
Sc	24.0	24.3	25.4	27.4	27.4	29.3	27.6	24.7	28.9
V	232	236	276	231	246	264	240	217	246
Cr	927	939	593	846	656	598	705	1015	583
Ni	729	716	441	619	468	253	533	804	415
Zn	122	121	145	113	108	108	111	114	110
Ga	16.6	16.9	19.4	17.0	17.3	19.9	17.7	15.7	17.9
Y	25.4	25.7	34.6	21.4	22.1	24.0	21.8	19.8	22.8
Zr	159	160	234	126	130	139	122	110	129
Nb	13.2	13.4	18.8	11.4	9.6	10.5	9.6	9.7	10.2
Hf	3.8	3.7	5.6	3.0	3.0	3.2	3.1	2.8	3.2
La	10.8	10.9	16.2	8.48	8.22	8.71	8.12	7.47	8.66
Ce	29.6	30.0	41.1	23.3	-	22.5	22.2	19.8	23.2
Nd	19.0	19.6	27.5	16.1	15.6	16.8	15.1	13.4	16.3
Sm	5.02	4.99	7.22	4.17	4.34	4.70	4.20	3.98	4.42
Eu	1.71	1.72	2.44	1.50	1.48	1.60	1.48	1.36	1.57
Tb	0.78	0.84	0.98	0.78	0.66	0.70	0.78	0.72	0.79
Yb	1.99	2.83	2.83	1.72	1.71	1.86	1.75	1.73	1.92
Lu	0.26	0.27	0.40	0.25	0.24	0.27	0.25	0.22	0.27

See Table 2 for discussion of analytical techniques, accuracy and precision. Major element data for the glasses associated with several of these whole rocks are in Table 1 of Garcia et al. [1989].

\*Samples 1-3 and 1-10 are petrographically different, but they have nearly identical major and trace element compositions. They are probably from the same flow.

†<sup>87</sup>Sr/<sup>86</sup>Sr for MK1-8 is 0.70365 and for MK6-6b is 0.70350. Two sigma is <0.005% and ratios are relative to E and A SrCO<sub>3</sub> = 0.70800 (P. Gurriet, analyst).

‡Olivine in this sample is heterogeneously distributed. The MgO-rich sample contains ~8% more olivine.

Only this shield sample was ground in agate; Ta values are 0.57 and 0.65, respectively.

§Mg# = 100 MgO/(MgO+FeO) with FeO = 0.9 total Fe.

east flank gulches are at a deeper stratigraphic level than the tholeiitic lavas from the west flank and Waikahalulu Gulch. Consequently, among the tholeiitic lavas there is a general trend of increasing relative enrichment in highly incompatible elements with decreasing age. However, some of the postshield tholeiitic lavas from the Hamakua coast have distinctive compositions. In particular, sample MU-2 is evolved in terms of MgO content (6.8%) and has low abundances of compatible trace elements, but relative to other postshield tholeiites it has the lowest incompatible element abundances and relatively high Zr/Nb and low La/Yb (Table 2, Figures 4 and 6). This sample is also isotopically distinct with unusually high <sup>207</sup>Pb/<sup>206</sup>Pb (15.55 [Kennedy et al., 1991]). Also east flank samples KI-10 and MU-10 are distinctive because their La/Yb and Zr/Nb ratios overlap with the field defined by tholeiites from the west and south flanks (Figure 6).

Abundances of incompatible elements are positively correlated in Hamakua lavas (see Figure 9 of Frey et

al.[1990]). The alkalic postshield lavas have higher abundances of incompatible elements than tholeiitic lavas with a similar MgO content (Figure 4). These postshield alkalic basalts also have relatively high La/Yb and low Zr/Nb (Figure 6). Therefore the basalt classification based on (Na<sub>2</sub>O+K<sub>2</sub>O) versus SiO<sub>2</sub> (Figure 3a) corresponds very well with the established fact that alkalic basalts generally have higher abundances of incompatible elements than tholeiitic basalts. Budahn and Schmitt [1985] reported trace element abundances for six Hamakua basalts collected from scattered localities on Mauna Kea Volcano. Although these samples were interpreted to be tholeiitic basalts, two of the six have high Na<sub>2</sub>O+K<sub>2</sub>O (3.37 and 3.49%) and La/Yb ratios (9.2 and 10); thus they overlap the alkalic basalts (Figures 3a and 6).

Consistent with their evolved major element contents, the high Fe-Ti basalts have very high abundances of incompatible elements (Figure 4) and relatively high LREE/HREE (LREE and HREE represent light and heavy rare earth elements, respectively) and low Zr/Nb (Figures 5 and 6). Despite their

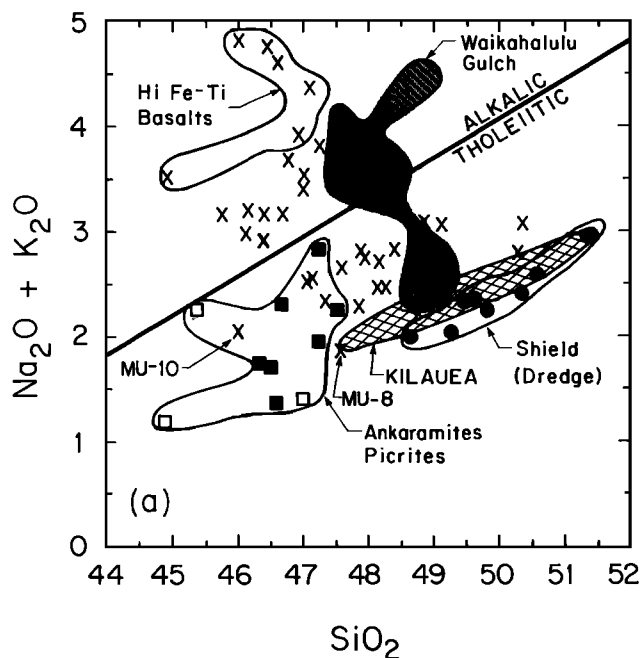


Fig. 3a. Total alkalis versus  $\text{SiO}_2$  (all in wt.%) plot showing alkalic-tholeiitic boundary line of Macdonald and Katsura [1964]. The Mauna Kea shield lavas (solid circles) are all tholeiitic and overlap with Kilauea basalts (Kilauea field based on Wright [1971]). In contrast, postshield basaltic lavas range from tholeiitic to alkalic. Lavas from the east flank of Mauna Kea are indicated by crosses (basalt), open squares (picrite), and solid squares (ankaramite). The six high Fe-Ti basalts include MU-6, MU-11 and KI-12 from Table 2 and three samples from Table 1 of Frey et al. [1990]. Also shown is a field for postshield Hamakua Volcanics from Waikahalulu Gulch on the upper south flank [Frey et al., 1990]. All data corrected to  $\text{FeO}/(\text{FeO} + 0.9 \text{Fe}_2\text{O}_3) = 0.85$ .

evolved compositions, they do not have negative Eu anomalies (Figure 5).

As expected from their high MgO and compatible trace element contents, the postshield picrites and ankaramites tend to have low abundances of incompatible elements (Table 2, Figure 4). However, ankaramite LP-5 has higher abundances of incompatible elements than the tholeiitic lavas (Figure 4). In fact, the  $\text{Zr}/\text{Nb-La}/\text{Yb}$  ratios of these highly porphyritic lavas span the range from tholeiitic to alkalic lavas (Figure 6). Ankaramite LP-5 has higher LREE abundances than the tholeiitic lavas but other picrites and ankaramites such as KI-3, -4, -5 and -8 have much lower REE contents and REE patterns similar to the tholeiitic lavas (Figure 5).

### 3.3. Radiogenic Isotope Ratios

Sr, Nd and Pb isotopic data for Hamakua lavas from the east flank gulches span narrow ranges [Kennedy et al., 1991]; e.g.,  $^{87}\text{Sr}/^{86}\text{Sr} = 0.70345$  to  $0.70361$ ;  $^{143}\text{Nd}/^{144}\text{Nd} = 0.51301$  to  $0.51308$ ;  $^{206}\text{Pb}/^{204}\text{Pb} = 18.36$  to  $18.48$ ;  $^{207}\text{Pb}/^{204}\text{Pb} = 15.44$  to  $15.49$ ;  $^{208}\text{Pb}/^{204}\text{Pb} = 37.89$  to  $38.03$ ; except for tholeiitic lava MU-2 which has  $^{207}\text{Pb}/^{204}\text{Pb} = 15.55$  and  $^{208}\text{Pb}/^{204}\text{Pb} = 38.07$ . There is complete isotopic overlap among the postshield tholeiitic and alkalic basalts. Also, Hamakua lavas from Waikahalulu Gulch and the west flank of Mauna Kea have Sr, Nd and Pb isotopic ratios which overlap with those of east flank lavas [Kennedy et al., 1991]. Moreover,  $^{87}\text{Sr}/^{86}\text{Sr}$  in the two shield tholeiites analyzed ( $0.70365$  for MK1-8 and  $0.70350$  for MK6-6, Table 1) are within the range defined by Hamakua lavas. In summary,

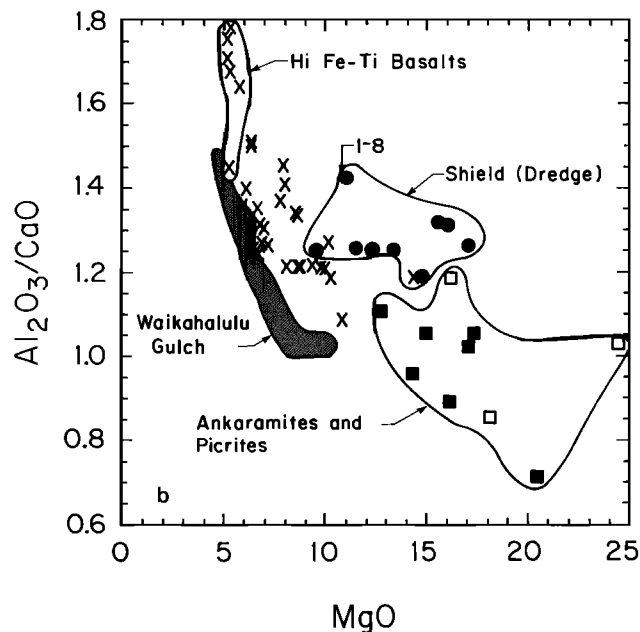


Fig. 3b.  $\text{Al}_2\text{O}_3/\text{CaO}$  versus MgO (all in wt.%). Fields and symbols as in (a). Many shield and postshield basalts have near-chondritic  $\text{Al}_2\text{O}_3/\text{CaO}$  (1.23), but the phenocryst-rich lavas range to lower ratios, and the evolved lavas (Hi Fe-Ti basalts) range to higher ratios. Lavas in the hawaiitic substage of postshield volcanism (Laupahoehoe Volcanics) have  $\text{Al}_2\text{O}_3/\text{CaO} > 2.3$  (not plotted, but see Figure 8a of Frey et al. [1990]).

Mauna Kea lavas are isotopically heterogeneous, but unlike lavas from Haleakala Volcano [Chen and Frey, 1985], there are no systematic variations in isotopic ratios with stratigraphic age.

### 3.4. Mineral Compositions

Olivine, clinopyroxene and plagioclase are the dominant phenocryst minerals in Mauna Kea lavas (Table 3). Phenocryst and microphenocryst compositions were determined for each distinctive lava type (Table 4) in order to evaluate which samples may be representative of crystallized melts and to identify samples that reflect magma mixing or crystal accumulation.

**Submarine Shield Tholeiites.** The cores of olivine grains in the dredged lavas range widely in forsterite (Fo) content (78-90; Table 4a). Olivine cores in four of the samples show small ranges in Fo content (1-4% Fo), but MK1-3 and 5-13 have larger ranges (7 to 12% Fo) and the lower Fo grains in samples 1-3 and 5-13 are reversely zoned. All of the higher Fo grains are normally zoned. Most of the samples contain Fo-rich (89-90) olivine and in several samples the olivine core compositions would be in equilibrium with melts having the bulk-rock composition (Figure 7). Some of the olivines have resorbed margins and weak kink banding (xenocrysts in Table 4a) but others with similar composition are euhedral and undeformed. Wilkinson and Hensel [1988] noted the same feature in picrites from Mauna Loa and Kilauea. They interpreted the kinked grains as early formed cognate minerals that were deformed as the magma flowed through narrow conduits at elevated pressures. In contrast, Helz [1987] and Albarede and Tamagnan [1988] concluded that picrites containing kink banded olivine formed as relatively evolved magmas intruded and incorporated xenocrysts from ultramafic cumulates.

Clinopyroxene and plagioclase phenocrysts and

TABLE 2a. Major (wt%) and Trace Element (ppm) Contents of Hamakua Lavas From the Hamakua Coast of Mauna Kea Volcano: Laupahoehoe Gulch

	Lava Type															
	Picrite		Alkalic			Tholeiitic			Alkalic			Ankaramites		Alkalic		Tholeiitic
	6	7	8	C-77	9	10	11	1	2	3	4	C-76	5	C-75	C-74	
SiO <sub>2</sub>	44.43	47.62	46.58	46.72	45.56	47.64	49.46	45.97	47.72	49.37	47.16	46.40	46.72	45.43	46.62	
TiO <sub>2</sub>	2.53	3.17	3.17	3.22	3.20	3.06	3.02	3.26	3.17	2.93	2.31	2.14	2.83	3.06	3.12	
Al <sub>2</sub> O <sub>3</sub>	10.47	13.34	13.20	14.74	12.97	14.44	13.43	13.82	14.60	13.24	10.32	11.16	10.66	13.58	14.93	
Fe <sub>2</sub> O <sub>3</sub>	14.34	13.56	13.80	13.63	14.00	13.67	13.58	14.00	14.20	13.20	12.89	12.67	14.40	13.99	14.45	
MnO	0.19	0.19	0.19	0.18	0.19	0.18	0.19	0.20	0.19	0.17	0.18	0.17	0.20	0.18	0.19	
MgO	15.90	7.77	8.55	7.97	9.66	6.87	6.05	8.49	6.62	6.61	14.20	14.95	12.57	10.08	6.81	
CaO	8.82	9.73	9.92	10.15	10.74	10.11	10.30	10.35	10.84	10.68	10.79	10.55	9.62	10.67	11.38	
Na <sub>2</sub> O*	1.74	3.13	3.17	2.53	2.85	2.67	2.51	2.74	2.44	2.32	1.84	1.51	2.20	2.25	2.26	
Na <sub>2</sub> O(†)	1.91	3.01	2.86	2.69	2.45	2.52	2.44	2.57	2.24	2.23	1.77	1.73	2.09	2.34	2.33	
K <sub>2</sub> O	0.32	1.05	0.92	0.80	0.72	0.17	0.60	0.55	0.19	0.54	0.45	0.56	0.71	0.81	0.20	
P <sub>2</sub> O <sub>5</sub>	0.31	0.50	0.48	0.33	0.38	0.33	0.32	0.49	0.34	0.30	0.25	0.18	0.34	0.26	0.25	
Total	99.22	99.94	99.67	100.28	99.87	99.14	99.41	99.70	100.11	99.27	100.32	100.29	100.14	100.31	100.21	
Mg#	68.7	53.2	55.1	53.7	57.8	49.9	46.9	54.6	48.0	49.8	68.6	70.0	63.40	58.8	48.3	
H <sub>2</sub> O*	0.43	0.81	0.32	0.51	0.30	0.44	0.14	0.70	0.82	0.38	0.16	0.30	0.30	0.67	0.67	
Rb	1.7	16.7	8.4	10.5	10.9	0.6	10.2	5.8	0.9	9.0	7.0	6.9	12.8	9.2	0.8	
Sr	382	485	505	493	496	425	409	562	433	396	356	360	452	477	451	
Ba	179	294	307	243	226	186	177	313	155	177	163	170	259	224	119	
Sc	25.1	27.0	27.4	27.0	30.7	32.1±1	29.6	28.1±1	31.3	31.6±6	32.2	32.3	27.8	30.7	32.9	
V	205	235	227	250	261	285	278	263	301	260	253	227	240	258	313	
Cr	984	298	338	252	420	109±1	58	338±4	85	221±13	1180	1180	918	430	64	
Co	80	51	54	52	61	49.1±0.1	47	53±1	49	46.8±8	69	71	68	61	51	
Ni	615	163	190	172	216	81	67	181	75	92	409	425	344	234	78	
Zn	121	133	119	110	116	131	128	135	126	111	111	97	104	116	134	
Ga	18	23	23	24	23	24	22	24	23	23	17	18	20	22	24	
Y	23	39	37	31	29	29	30	39	30	29	21	21	25	29	31	
Zr	194	347	313	240	236	202	206	347	211	190	162	152	216	226	212	
Nb	17.4	31.9	29.0	22.5	23.0	16.2	18.0	32.0	17.8	15.9	16.4	15.3	23.5	23.0	18.0	
Hf	4.4	7.7	6.9	5.7	5.6	4.78±0.4	4.9	7.5±1	4.7	4.6±2	3.7	3.6	4.8	5.1	5.1	
Ta	0.93	1.74	1.53	1.36	1.31	1.01±0.1	0.95	1.67±0.8	1.11	0.95±0.4	0.90	0.98	1.38	1.39	1.08	
Th	0.6	1.9	1.5	1.4	1.1	0.9	0.9	2.0±1	0.6	0.8	0.6	0.6	0.9	1.4	1.0	
La	14.9	28.1	25.9	21.1	20.6	15.0±1	15.2	27.7±1	15.1	14.5±6	13.8	13.3	19.2	19.4	16.0	
Ce	36.7	66.8	63.3	52.8	48.6	38.0±0.4	37.7	66.3±1	37.1	36.3±1	35.2	34.2	45.0	49.1	40.6	
Nd	22.3	38.5	36.7	31.6	28.9	25.0±0.3	24.2	38.5±1	25.1	23.0±1	20.8	20.1	25.4	28.9	26.1	
Sm	5.76	9.50	9.22	7.94	7.47	6.55±0.20	6.64	9.71±1	6.76	6.4±3	5.09	5.10	6.25	7.21	6.96	
Eu	2.11	3.31	3.15	2.76	2.69	2.43±0.1	2.42	3.31±0.4	2.45	2.32±0.3	1.87	1.79	2.27	2.57	2.53	
Tb	0.91	1.43	1.54	1.13	0.99	1.10±0.15	0.96	1.31±0.16	1.01	0.94±0.8	0.67	0.82	0.97	1.12	1.11	
Ho	0.9	1.5	1.1	0.9	1.0	1.05±0.1	1.0	1.5±1	0.9	0.99±0.9	0.7	0.6	0.8	1.1	1.1	
Yb	1.63	3.03	2.86	2.34	2.31	2.25±0.3	2.37	2.95±0.6	2.31	2.13±0.2	1.64	1.54	1.89	2.10	2.48	
Lu	0.25	0.38	0.37	0.32	0.31	0.31±0.1	0.32	0.40±0.1	0.33	0.29±0.1	0.21	0.23	0.28	0.29	0.36	

Samples in each gulch are in stratigraphic order, oldest to youngest (Figure 2). For Tables 1 and 2, major abundances of Rb, Sr, Ba, V, Ni, Zn, Ga, Y, Zr and Nb are means of duplicate analyses determined by X ray fluorescence at the University of Massachusetts (J.M. Rhodes, supervisor). Abundances of Sc, Cr, Co, REE, Hf, Ta and Th were analyzed by instrumental neutron activation analysis at the Massachusetts Institute of Technology (P. Ila, supervisor). Means of duplicate analyses are indicated for LP-10, LP-1, LP-3, MU-6, MU-17 and KI-9. Evaluations of precision and accuracy are in Table 1d of Frey et al. [1990]. In addition, sample C-74 was analyzed by neutron activation by Budahn and Schmitz [1985]. Most of their data are within 5% of our values, and all data agree within 10%. Major element data for "C" samples in Table 2c are from Macdonald and Katsura [1964], recalculated to be anhydrous and oxidized. Mg# = 100 MgO/(MgO+FeO\*) on a molar basis with all Fe<sub>2</sub>O<sub>3</sub> converted to FeO.

\*Na<sub>2</sub>O determined by X-ray fluorescence.

†Na<sub>2</sub>O determined by neutron activation.

microphenocrysts are rare in these shield lavas (Table 3). Clinopyroxene compositions are typical for tholeiitic lavas (i.e., low abundances of TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, CaO and Na<sub>2</sub>O but high SiO<sub>2</sub>; Table 4b). Plagioclase microphenocrysts are labradorites (An<sub>58-67</sub>; Table 4c). The low An grain from sample 1-8 is partially resorbed. The CaO/Na<sub>2</sub>O partition coefficients (i.e., K<sub>D</sub> for plagioclase/glass and plagioclase/whole-rock) are low (0.71-0.80). In anhydrous experimental systems this K<sub>D</sub> is ~1 at low pressure and increases with increasing pressure [e.g. Juster et al., 1989].

**Hamakua Tholeiitic Basalts.** Sample MU-8 may be the most magnesian postshield tholeiitic lava that is unambiguously a crystallized melt. The relatively MgO-rich nature of the melt is reflected by the rare magnesian olivine (Fo<sub>85.6</sub>) and clinopyroxene phenocrysts (Tables 4a and 4b).

Olivine microphenocryst whole rock Fe/Mg ratios yield a K<sub>D</sub> ~0.30±0.03 (Figure 7) which is consistent with the hypothesis that the bulk rock composition represents a melt composition. Sample MU-2 is an evolved (6.6% MgO) but compositionally distinctive tholeiite. Phenocryst core compositions (Fo<sub>79</sub>) are too forsteritic for the bulk rock composition (Figure 7). Therefore, the bulk rock composition is not representative of a melt composition. The olivine microphenocrysts are less magnesian (Fo<sub>61-69</sub>) and apparently reflect equilibrium with highly evolved residual melt.

**Hamakua Alkalic Basalts.** Sample LP-9 is MgO-rich (9.66%) and contains <2% phenocrysts (Table 3). The cores of the normally zoned olivine phenocrysts have compositions (Fo<sub>84</sub>) which are consistent with crystallization from a melt having the Fe/Mg ratio of the bulk rock (Figure 7).

TABLE 2b Maulua Gulch

	Lava Type															
	Tholeiitic						Alkalic	Fe-Ti	Tholeiitic		Ank.	Thol.	Fe-Ti	Alkalic		
	13	14	15	2	3	4	5	6	7	8	9	10	11	17	16	
SiO <sub>2</sub>	47.12	46.85	46.57	47.68	47.83	46.65	45.93	44.28	47.47	47.02	45.93	45.48	46.18	46.06	47.53	
TiO <sub>2</sub>	3.29	3.42	3.32	2.62	3.32	3.03	3.28	4.73	3.25	2.74	2.04	2.53	4.83	3.49	3.11	
Al <sub>2</sub> O <sub>3</sub>	14.82	15.01	14.48	14.98	14.35	13.96	14.32	15.24	14.28	12.19	9.62	11.58	14.61	14.97	14.91	
Fe <sub>2</sub> O <sub>3</sub>	14.41	14.96	14.46	13.25	14.52	13.73	13.49	16.22	14.55	13.76	13.53	14.01	15.74	14.55	13.46	
MnO	0.19	0.20	0.20	0.19	0.19	0.17	0.17	0.20	0.19	0.19	0.19	0.18	0.21	0.21	0.17	
MgO	6.03	6.25	6.74	6.82	5.93	8.05	7.93	5.76	6.28	10.76	15.93	14.28	5.15	6.29	5.82	
CaO	10.62	10.01	11.42	11.91	10.79	11.51	10.18	9.29	11.13	11.24	10.80	9.75	8.57	9.94	10.97	
Na <sub>2</sub> O(*)	2.66	2.61	2.18	2.14	2.52	2.04	2.80	2.84	2.36	1.39	1.26	1.71	2.90	2.90	2.50	
Na <sub>2</sub> O(†)	2.56	2.43	2.34	2.23	2.48	2.10	2.74	2.80	2.34	1.65	1.46	1.80	2.71	2.54	2.50	
K <sub>2</sub> O	0.22	0.21	0.18	0.20	0.31	0.19	0.88	0.67	0.40	0.18	0.22	0.21	1.20	0.81	0.80	
P <sub>2</sub> O <sub>5</sub>	0.33	0.35	0.32	0.27	0.35	0.31	0.41	0.70	0.34	0.29	0.22	0.29	0.61	0.37	0.39	
Total	99.59	99.69	100.03	100.15	100.07	99.70	99.33	99.89	100.23	100.02	99.94	100.11	99.81	99.20	99.66	
Mg#	45	45	48	50.5	45	54	54	41	46	61	70	67	39	46	46	
H <sub>2</sub> O <sup>+</sup>	0.85	1.30	0.54	0.29	0.31	0.49	0.25	0.80	0.38	0.71	0.53	0.43	1.56	0.97	0.68	
Rb	0.8	0.7	1.0	1.9	1.8	0.5	16.9	3.8	5.8	0.5	1.2	2.2	22.9	9.3	12.4	
Sr	436	400	428	443	493	421	603	593	488	355	312	427	510	481	577	
Ba	167	167	158	138	207	162	303	420	217	161	162	183	422	286	270	
Sc	31.2	31.3	32.6	31.9	29.3	32.3	26.1	25.6±.4	31.3	36.4	33.5	27.3	29.9	29.3±.7	27.6	
V	298	290	324	262	304	310	234	270	269	297	243	259	338	333	289	
Cr	48	63	63	65	52	247	263	62±1	65	866	1390	745	61	72±4	115	
Co	49	50	51	49	47	53	52	49.1±.6	50	61	76	70	44	49.5±.9	46	
Ni	67	73	84	59	46	152	134	79	67	239	423	461	60	77	74	
Zn	140	136	131	111	126	125	117	150	105	116	101	111	135	114	113	
Ga	25	24	25	23	24	24	24	30	25	21	16	19	27	24	23	
Y	34	35	33	27	33	30	33	45	32	27	20	23	43	36	27	
Zr	245	258	229	176	231	207	284	416	212	182	140	178	357	221	223	
Nb	21.4	22.7	20.1	14.3	20.6	18.4	28.4	46.3	20.2	16.1	14.1	17.6	40.2	25.2	24.1	
Hf	5.6	5.7	5.3	3.9	5.3	4.7	6.3	9.0±.2	4.9	4.2	3.3	4.1	8.1	5.3±.1	5.0	
Ta	1.30	1.43	1.22	0.83	1.27	1.08	1.68	2.6±.2	1.17	1.02	0.83	1.03	2.47	1.46±.09	1.54	
Th	1.1	1.2	1.3	0.7	1.4	1.0	1.7	3.8	1.0	0.7	0.4	0.6	3.0	1.5±.1	1.5	
La	19.3	19.9	17.0	12.7	17.6	15.5	23.7	38.5±.4	17.2	14.3	12.2	15.2	34.9	26.0±.5	21.5	
Ce	45.4	47.2	42.8	31.8	44.2	40.0	61	92.1±.1	42.9	35.9	29.0	38.0	82	51±2	51	
Nd	29.1	30.8	28.0	20.5	27.4	25.0	35.3	53.1	27.1	22.6	18.3	23.2	48	31±1	29.7	
Sm	7.53	8.01	7.34	5.59	7.12	6.70	8.10	12.2±.2	6.79	5.91	4.52	5.47	11.0	7.41±.09	6.77	
Eu	2.68	2.76	2.68	2.08	2.62	2.40	2.76	4.02±.10	2.49	2.15	1.67	1.94	3.73	2.74±.06	2.47	
Tb	1.23	1.35	1.20	0.88	1.02	1.00	1.10	1.98	1.01	0.86	0.76	0.95	1.68	1.14±.17	1.10	
Ho	1.2	1.5	1.3	0.7	1.1	1.0	1.2	1.67±.01	1.1	0.9	0.64	0.73	1.5	1.2±.1	0.93	
Yb	2.48	2.43	2.48	2.10	2.40	2.27	2.42	3.31±.01	2.45	1.90	1.46	1.72	3.05	2.27±.13	1.98	
Lu	0.35	0.34	0.34	0.28	0.33	0.31	0.33	0.45±.01	0.34	0.25	0.20	0.25	0.44	0.31±.01	0.26	

Ank., ankaramite; thol., tholeiitic; Fe-Ti, high Fe-Ti basalt. MU-10 is in the tholeiitic field in Figure 3a, but it has incompatible element abundance ratios typical of alkalic lavas (Figure 6). Also, see Table 2a footnotes.

Plagioclase/whole rock CaO/Na<sub>2</sub>O ratios are ~0.93 for phenocrysts and microphenocrysts. In contrast, the MgO-rich (14.3%) sample MU-10 which is transitional between tholeiitic and alkalic compositions has a wide range of olivine phenocryst compositions, and these olivines are too poor in forsterite to be in equilibrium with a melt having the bulk rock composition (Figure 7). Evidence for two generations of olivine in this sample is that olivines with cores of ~Fo<sub>83</sub> contain spinel inclusions whereas the olivines with lower Fo (~77) lack spinel inclusions. Sample MU-5 is a slightly more evolved (7.9% MgO) alkalic basalt with ~4.5% olivine phenocrysts (Table 3). The olivine compositions (Fo<sub>80-82</sub>) yield olivine/whole rock Fe/Mg ratios slightly greater than 0.3 (Figure 7). Olivine microphenocrysts in MU-5 are much more Fe-rich, Fo<sub>66</sub> (Table 4a), and reflect equilibration with an evolved residual melt.

**Hamakua High Fe-Ti Basalts.** Although the high Fe-Ti basalts have similar FeO/MgO ratios, the cores of their olivine xenocrysts, phenocrysts and microphenocrysts range widely (Fo<sub>59.4</sub> to Fo<sub>82.9</sub>; Table 4a). The cores of olivine phenocrysts in MU-11 are too high in Fo content to be in equilibrium with a melt having the bulk rock composition, i.e.,  $K_D = 0.16$  to 0.28. Also this sample contains rare resorbed olivine xenocrysts with slight reverse zoning (Table 4a). In addition, the cores of plagioclase phenocrysts in sample MU-11 have a

wide compositional range (An<sub>66.8</sub> to An<sub>84.1</sub>, - Table 4c). The ~An<sub>84</sub> phenocrysts are outside the typical range, An<sub>58</sub>-An<sub>67</sub>, of plagioclase phenocrysts in the high Fe-Ti basalts. These typical core compositions yield (plag/whole-rock CaO/Na<sub>2</sub>O) values of 0.9 to 1.2 which are characteristic of low pressure crystallization [Juster et al., 1989]. The rare clinopyroxene phenocrysts in KI-12 are reversely zoned (i.e., the rim has lower TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub> and Na<sub>2</sub>O, but higher MgO than the core; Table 4b). Also a clinopyroxene microphenocryst has higher Mg/Fe than the phenocryst core or rim (Table 4b); apparently, the clinopyroxene phenocryst composition grew in a more evolved melt than the whole rock composition. The olivine microphenocrysts in this sample are also low in Fo content (62.6 to 59.4, Table 4a). In summary, the textural characteristics and mineral compositions in these high Fe-Ti basalts provide evidence for disequilibrium. Apparently, some of the phenocrysts were incorporated into the magma by disaggregation of wallrocks or by mixing with magmas of different composition.

**Hamakua Picrite and Ankaramite.** Cores of olivine phenocrysts in picrite LP-6 are less magnesian (~Fo<sub>85</sub>) than most of the olivine cores in the olivine-rich tholeiitic lavas dredged from the Mauna Kea shield (Table 4a and Figure 7). Olivine/whole rock Fe/Mg ratios are high (>0.4) indicating that the olivine did not grow from a melt with Fe/Mg similar to the

TABLE 2c. Kaawalii Gulch

	Lava Type												
	Alkalic		Picrite	Ank.		Alkalic		Picrite	Thol.		Ank.	Fe-Ti	Thol.
	1	2	3	4	5	6	7	8	9	10	11	12	13
SiO <sub>2</sub>	45.73	45.48	44.38	45.99	46.85	45.65	45.71	46.25	47.22	48.18	46.23	46.33	48.54
TiO <sub>2</sub>	3.09	3.11	1.55	2.17	2.08	3.28	3.41	1.81	2.76	2.98	1.46	4.74	3.17
Al <sub>2</sub> O <sub>3</sub>	13.06	12.78	7.63	9.99	9.67	13.10	13.41	8.78	13.57	14.16	7.61	13.72	13.88
Fe <sub>2</sub> O <sub>3</sub>	13.79	13.90	13.30	13.40	12.91	13.89	13.93	12.71	13.76	13.09	12.31	14.86	13.49
MnO	0.18	0.18	0.19	0.17	0.18	0.18	0.19	0.16	0.19	0.18	0.16	0.20	0.18
MgO	9.82	10.20	24.24	17.22	16.97	9.14	8.68	17.89	8.60	6.46	20.31	5.17	6.30
CaO	10.82	10.77	7.38	9.47	9.41	10.84	11.08	10.29	11.23	11.38	10.66	9.50	10.96
Na <sub>2</sub> O(*)	2.23	2.42	1.08	1.40	1.30	2.45	2.17	1.76	2.27	2.63	0.95	3.22	2.48
Na <sub>2</sub> O(†)	2.26	2.32	1.08	1.51	1.55	2.41	2.25	1.24	2.06	2.33	1.11	3.03	2.43
K <sub>2</sub> O	0.64	0.64	0.082	0.21	0.39	0.70	0.66	0.12	0.17	0.68	0.23	1.31	0.57
P <sub>2</sub> O <sub>5</sub>	0.36	0.35	0.16	0.21	0.22	0.39	0.38	0.19	0.29	0.36	0.15	0.68	0.36
Total	99.75	99.73	99.99	100.34	100.23	99.58	99.70	99.44	99.81	100.10	100.23	99.54	99.33
MgO#	59	59	78	72	72	57	55	74	55	49	77	41	48
H <sub>2</sub> O+	0.30	0.15	0.39	0.39	0.19	0.24	0.35	0.40	0.73	0.38	0.19	0.30	0.15
Rb	11.1	10.4	0.7	1.2	7.4	11.9	9.3	0.3	0.8	11.1	2.2	25	8.4
Sr	515	507	214	317	296	504	508	260	415	515	246	576	465
Ba	237	237	66	142	128	246	242	97	139	254	109	399	205
Sc	29.9	29.9	22.4	28.2	28.0	30.2	30.7	30.9	32.9±0.6	29.4	33.9	29.3	30.2
V	249	238	190	220	209	259	250	205	279	273	180	338	296
Cr	418	415	1910	1400	1245	375	373	1800	462±6	59	2070	59	62
Co	60	63	97	78	77	58	56	77	54±1	46	86	40	49
Ni	212	213	966	585	636	177	172	544	160	87	719	57	79
Zn	117	106	110	108	107	106	104	110	115	104	94	141	121
Ga	22	22	13	16	16	23	23	15	22	22	13	27	23
Y	29	27	15	21	20	30	31	18	26	27	13	43	29
Zr	225	211	105	146	144	241	235	123	182	212	92	386	225
Nb	22.0	20.1	10.0	13.0	13.3	23.0	23.5	11.4	16.7	25.0	10.5	45.7	21.2
Hf	5.1	5.0	2.4	3.4	3.4	5.6	5.5	2.6	4.44±0.06	4.6	2.2	9.0	5.1
Ta	1.3	1.3	0.52	0.81	0.75	1.4	1.4	0.6	0.96±0.01	1.4	0.59	2.7	1.2
Th	1.0	0.9	0.2	0.3	0.3	1.1	1.0	0.3	0.6	1.4	0.3	2.4	1.3
La	18.3	17.8	7.3	11.3	11.1	20.1	19.2	8.47	14.8±0.6	20.3	8.34	37.7	18.9
Ce	45	42	18.8	28.0	26.7	49	50	21.2	38±2	51	20.1	96	46
Nd	26	25	12.1	18.0	16.6	29	30	13.8	23.2±4	29	12.7	55	27.5
Sm	7.01	6.68	3.26	4.61	4.53	7.41	7.48	3.87	6.10±0.08	6.17	3.09	11.7	7.15
Eu	2.51	2.46	1.19	1.69	1.65	2.66	2.59	1.45	2.16±0.06	2.38	1.12	3.70	2.56
Tb	1.16	1.00	0.62	0.84	0.78	1.15	1.13	0.71	1.00±0.08	0.98	0.46	1.54	1.04
Ho	0.9	0.9	0.5	0.7	0.7	1.2	1.1	0.5	0.98±0.02	1.1	0.8	1.8	1.1
Yb	2.12	1.96	1.12	1.47	1.49	2.25	2.16	1.31	2.06±0.03	1.98	1.01	3.19	2.25
Lu	0.29	0.29	0.16	0.22	0.22	0.30	0.30	0.19	0.28±0.01	0.29	0.14	0.44	0.31

KI-10 is in the tholeiitic field in Figure 3a, but it has incompatible element abundance ratios typical of alkalic lavas (Figure 6). Also, see Tables 2a and 2b footnotes.

bulk rock composition (Figure 7). Like the olivine-rich shield samples, LP-6 contains unstrained and kink-banded olivine phenocrysts of similar composition (~Fog<sub>5</sub>). It also has olivine xenocrysts which are slightly lower in MgO (Fog<sub>2</sub>). Cores of olivine phenocrysts in ankaramites MU-9 and LP-5, range from Fog<sub>3</sub> to Fog<sub>7</sub> (Table 4a). Olivine-core/whole rock Fe/Mg ratios for sample MU-9 yield  $K_D > 0.4$  (Figure 7). However, olivines in sample LP-5 have compositions that would be in equilibrium with a melt having the bulk-rock composition (Figure 7).

#### 4. Discussion

##### 4.1. Origin of Postshield High Fe-Ti Basalts

A compositionally distinctive group of high Fe-Ti basalts were erupted during the basaltic substage of postshield volcanism at Mauna Kea (e.g., Figures 3-6 and Frey et al. [1990]). Compositionally similar lavas also occur in the postshield lavas of Haleakala Volcano (e.g., sample H-85-11

discussed by Chen et al. [1990]). Depending upon the classification scheme used these lavas are basalts [LeBas et al., 1986] or hawaiites [Coombs and Wilkinson, 1969]. A compilation of data for Mauna Kea lavas [Wolfe et al., 1991] includes 28 Hamakua lavas with high abundances of TiO<sub>2</sub> (4.05 to 5.17%), Fe<sub>2</sub>O<sub>3</sub> (14.3 to 18%) and P<sub>2</sub>O<sub>5</sub> (0.42 to 0.92%) but with moderate MgO contents (4.5 to 6.2%; see Figure 8). Although these high Fe-Ti lavas must have formed by a similar process during the latter part of the basaltic substage, they do not lie on a single liquid line of descent. For example, they do not form a well-defined P<sub>2</sub>O<sub>5</sub>-TiO<sub>2</sub> trend (Figure 8), and sample MU-11 has 10-20% lower incompatible elements contents than samples MU-6 and KI-12 which have similar or higher MgO contents (Table 2). Samples MU-6 and MU-11 also have distinct <sup>87</sup>Sr/<sup>86</sup>Sr (0.70345 and 0.70361, respectively) and <sup>206</sup>Pb/<sup>204</sup>Pb (18.44 and 18.36, respectively), but these values are within the range of other Hamakua lavas [Kennedy et al., 1991]. Therefore, it is likely that compositionally distinct parental magmas evolved to different high Fe-Ti basalt compositions. In addition the



TABLE 3. Modes for Lavas From the Maulua, Laupahoehoe, and Kaawalii Gulch Sections and the Submarine portion of the East Rift Zone of Mauna Kea Volcano, Hawaii

	Olivine		Clinopyroxene		Plagioclase		Opakes	Groundmass
	ph	mph	ph	mph	ph	mph	mph	
<b>Maulua</b>								
MU-16	1.1	0.4	1.3	0.6	9.1	2.5	0.1	84.9
MU-17	0.0	2.4	0.0	0.5	0.0	1.4	3.5	92.2
MU-12	1.6	6.3	0.9	10.9	0.2	7.9	0.6	71.6
MU-11	0.0	0.2	0.0	<0.1	0.4	0.6	<0.1	98.7
MU-10	8.5	12.1	0.0	0.0	0.0	0.1	<0.1	79.3
MU-9	16.6	0.2	8.2	4.1	0.1	0.1	0.2	70.5
MU-8	0.3	1.1	0.1	0.0	0.0	0.0	<0.1	98.5
MU-7	0.0	2.8	0.0	2.6	0.1	6.2	10.4	77.9
MU-6	0.1	0.0	0.0	0.0	1.2	0.0	0.1	98.6
MU-5	2.0	2.5	0.0	0.0	0.0	0.7	<0.1	94.8
MU-4	2.5	1.1	0.0	2.9	0.5	4.4	0.3	85.6
MU-3	<0.1	<0.1	<0.1	0.4	2.8	1.5	0.0	95.3
MU-2	0.4	0.5	0.6	0.9	0.4	4.0	<0.1	93.2
MU-1	1.2	0.4	0.2	0.1	3.6	1.2	<0.1	93.1
MU-15	0.3	0.1	0.1	0.6	0.8	1.5	0.0	96.6
MU-14	0.1	0.1	0.1	0.8	0.7	2.0	0.0	96.7
MU-13	<0.1	0.1	0.1	0.9	0.6	2.0	0.0	96.3
<b>Laupahoehoe</b>								
C-76	17.4	0.4	10.7	1.9	1.4	0.5	0.3	61.5
LP-5	13.3	0.0	13.6	0.0	0.0	1.8	0.2	71.1
LP-4	14.1	1.9	10.7	3.5	0.4	0.3	0.1	68.7
LP-3	0.0	0.0	0.0	0.0	0.0	0.0	0.0	100.0
LP-2	0.0	0.0	0.0	0.0	0.0	0.0	0.0	100.0
LP-1	3.1	4.8	0.0	0.0	0.0	0.0	0.0	92.1
LP-10	<0.1	0.1	<0.01	2.2	<0.01	3.1	0.0	94.6
LP-11	<0.1	0.0	0.4	0.0	2.3	0.0	0.0	97.3
LP-9	1.6	5.6	<0.1	0.0	0.0	1.9	0.0	90.9
LP-8	0.2	2.5	0.0	0.0	0.0	0.0	0.0	97.3
LP-7	2.4	1.1	0.0	0.0	0.0	0.0	0.0	96.5
LP-6	21.3	5.7	0.0	0.0	0.0	1.2	0.1	71.7
<b>Kaawalii</b>								
KI-13	0.0	0.9	0.0	6.1	0.0	6.9	0.0	86.1
KI-12	0.0	1.1	0.0	3.0	0.6	5.3	0.1	90.0
KI-11	27.7	3.8	14.8	8.3	0.6	0.0	0.1	44.7
KI-10	0.2	3.7	0.2	3.2	0.8	15.2	0.0	76.7
KI-9	5.1	2.6	5.9	3.0	1.5	3.8	0.1	78.0
KI-8	15.0	3.6	3.8	2.8	0.8	0.3	0.2	73.5
KI-7	3.0	16.4	0.0	0.0	0.4	40.6	0.5	39.1
KI-6	2.3	12.6	0.0	1.4	0.4	16.4	0.1	66.8
KI-5	13.1	1.4	9.5	2.5	1.0	1.5	0.3	70.7
KI-4	18.6	3.0	9.0	3.0	0.0	0.4	0.3	65.7
KI-3	34.3	4.0	0.0	0.0	0.0	0.1	0.4	61.2
KI-2	3.8	11.7	0.7	12.1	3.0	44.2	3.7	20.8
KI-1	4.1	9.5	0.6	9.9	3.0	26.7	5.2	41.0
<b>Submarine Tholeiites</b>								
MK1-3	13.8	8.6	0.0	0.0	0.0	0.0	0.0	77.6
MK1-8	16.4	2.1	0.6	1.6	0.8	2.4	0.0	76.1
MK1-10	16.8	11.0	0.0	0.0	0.0	0.0	0.0	72.2
MK2-1	14.0	5.4	0.0	3.4	0.2	5.2	0.0	71.8
MK5-13	14.0	3.2	0.0	0.2	0.2	6.8	0.0	75.8
MK6-6	22.0	1.8	0.2	4.2	0.4	9.6	0.0	61.8
MK6-18	5.2	4.6	0.0	0.6	0.0	2.4	0.0	87.2

Samples in the gulch sections are in stratigraphic order from top to bottom (Figure 2). Phenocrysts (ph) are >0.7 mm; microphenocrysts (mph) 0.1-0.7 mm; groundmass <0.1 mm including glass, if present. Modes based on 1000 counts per sample.

disequilibrium phenocrysts/xenocrysts in these lavas (Table 4) provide evidence for magmatic complexities that would cause deviations from a liquid line of descent.

A plausible model for the origin of the high Fe-Ti-rich lavas is crystal fractionation from a melt. For example, Chen et al. [1990] showed that a high Fe-Ti basalt ( $\text{Fe}_2\text{O}_3 = 15.04$ ,  $\text{TiO}_2 = 4.30$ ,  $\text{MgO} = 4.95$ ) could be derived from a more mafic alkalic basalt (6.28% MgO) by ~40% segregation of a clinopyroxene and plagioclase-rich assemblage. The high Fe-

Ti lavas from Mauna Kea are nearly aphyric (e.g., MU-6, MU-11 and KI-12 in Table 3); therefore, the bulk rock compositions must closely reflect melt compositions. In order to test a crystal fractionation model for major elements we used a simple mass balance model (Table 5) to evaluate the suitability of various Hamakua basalts as parents for the high Fe-Ti basalts. Some alkalic basalts, e.g., MU-5, have compositions consistent with a parental relationship but others, e.g., the evolved lavas KI-10 and KI-13, are not suitable

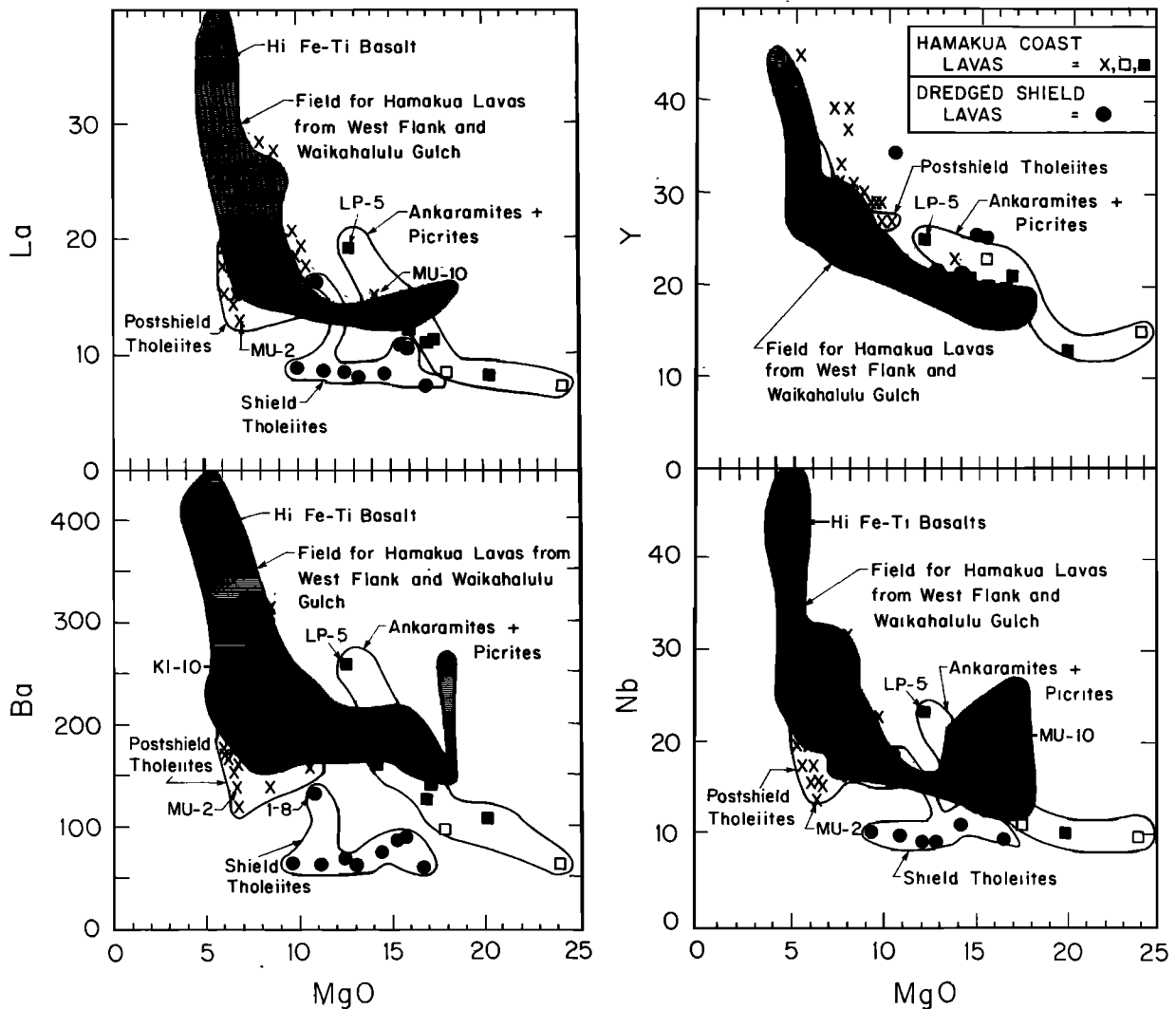


Fig. 4. Abundances of the incompatible elements, La, Ba, Y and Nb (ppm) versus MgO (wt.%) in shield lavas and the Hamakua Volcanics. Symbols as in Figure 3. The postshield tholeiitic lavas (labeled field) have lower abundances of La, Ba and Nb than alkalic postshield lavas (x's not within tholeiitic field). However, postshield alkalic and tholeiitic basalts have similar Y abundances. Shaded field for Hamakua Volcanics from west flank and Waikahalulu Gulch is from Frey et al. [1990].

parents, primarily because of their high  $\text{SiO}_2$  and low  $\text{TiO}_2$ . In addition, alkali basalt MU-5 and the immediately overlying high Fe-Ti basalt, MU-6 (Figure 2) overlap in Sr, Nd and Pb isotopic ratios [Kennedy et al., 1991]. Because of the major element compositional differences among the high Fe-Ti basalts (Table 2), these mass-balance models yield slightly different results. However, in each case ~40% segregation of plagioclase-rich assemblage is required (Table 5c). For example, the major element modelling for MU-5 to KI-12 requires ~43% segregation of a cumulate with ~50% plagioclase and subequal amounts of olivine and clinopyroxene (Table 5c). This result is in excellent agreement with the experimental results of Mahood and Baker [1986, Tables 1 and 2], who showed that evolved lavas rich in Fe and Ti were formed from alkalic basalt at 1 atm by segregation of a solid with 51% plagioclase, 20% olivine and 29% cpx (Table 5c). Moreover, olivine gabbro xenoliths in the overlying Laupahoehoe Group lavas have similar modal proportions (Table 5c). These xenoliths have been interpreted as cumulates from Hamakua alkali basalt [Fodor and Van der

Meyden, 1987], and they have Sr and Pb isotopic ratios consistent with this hypothesis [Kennedy et al., 1991].

In detail, the calculated plagioclase/clinopyroxene ratio in the fractionating assemblage is significantly lower for samples MU-6 and MU-12 than for sample KI-12 (Table 5c). Wolfe et al. [1991] recognized that  $\text{Al}_2\text{O}_3/\text{CaO}$  ratios in the high Fe-Ti basalts vary from 1.3 to 2.1 although only ~25% of the lavas with  $\text{TiO}_2 > 4\%$  have  $\text{Al}_2\text{O}_3/\text{CaO} > 1.5$  [Wolfe et al., 1991, Table 1]. They suggested that the high Fe-Ti basalts with lower  $\text{Al}_2\text{O}_3/\text{CaO}$ , such as KI-12, formed at lower pressures by segregation of an assemblage with  $\text{cpx}/\text{plag} < 1$ , whereas the less common high Fe-Ti basalts with high  $\text{Al}_2\text{O}_3/\text{CaO}$  formed at higher pressures by segregation of an assemblage with  $\text{cpx}/\text{plag} > 1$ . Our results are consistent with this interpretation, but we emphasize that plagioclase was a major fractionating phase in creating the high Fe-Ti basalts.

Qualitatively, trace element abundances are consistent with the plagioclase-rich cumulate required by the major element models for deriving the high Fe-Ti basalts. Specifically, relative to alkali basalt MU-5 the high Fe-Ti basalts are highly

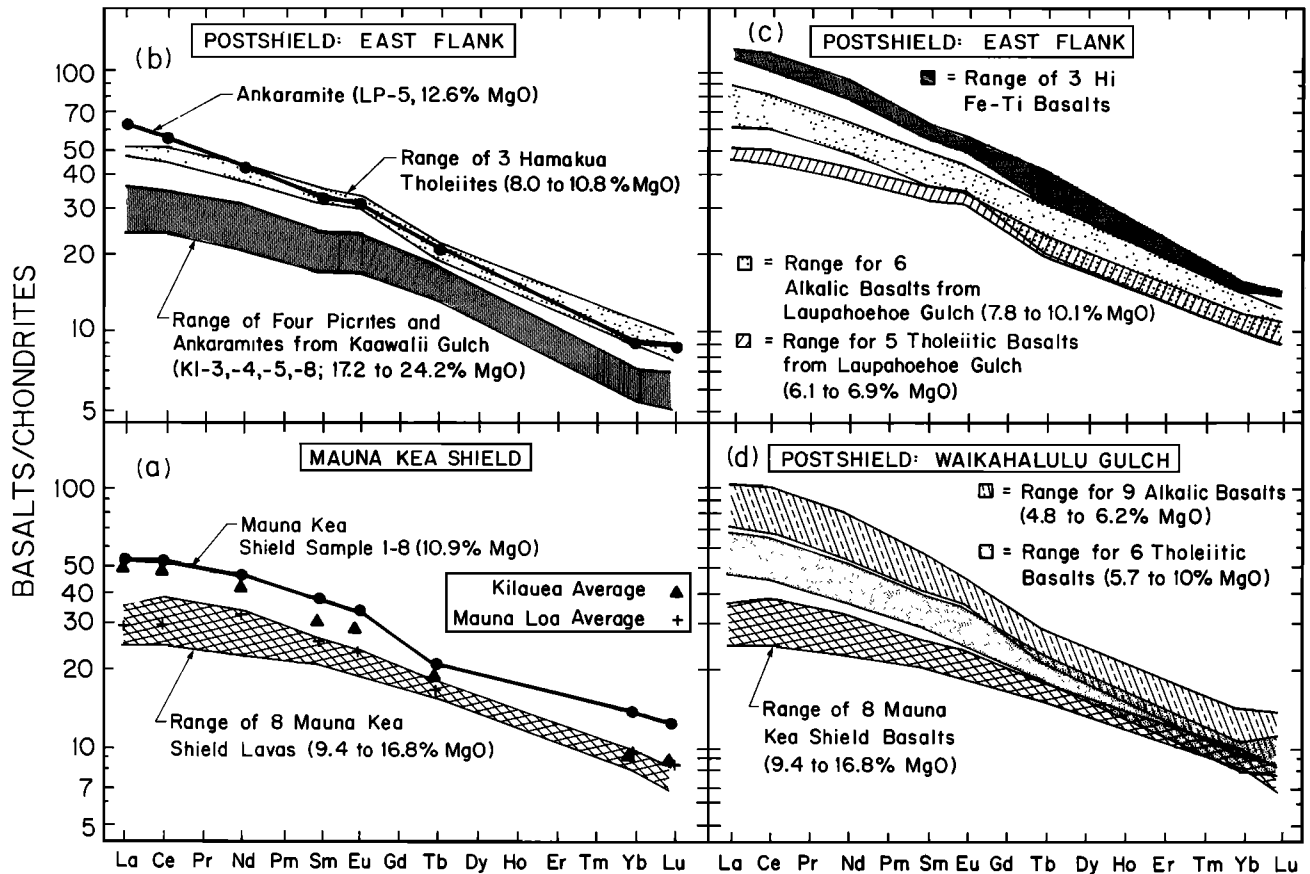


Fig. 5. Chondrite-normalized [Boynnton, 1984] REE patterns for Mauna Kea lavas. (a) Shield lavas: field for eight shield lavas encompasses an average for Mauna Loa basalt [Budahn and Schmitt, 1985]. Shield sample 1-8 is distinctive in that it is enriched in all REE (and other incompatible elements, Table 1) relative to the other Mauna Kea shield lavas. (b) Hamakua Volcanics - High MgO lavas from the east flank: Most postshield ankaramites and picrites have relatively low REE contents and several have REE patterns subparallel to the postshield Hamakua tholeiitic basalts. However, ankaramite LP-5 is highly enriched in LREE and has the trace element characteristics of an alkalic basalt. (c) Hamakua Volcanics - Laupahoehoe Gulch and Hi Fe-Ti basalts: Fields for alkalic and tholeiitic postshield basalts from Laupahoehoe Gulch are similar to the respective fields for postshield lavas from Waikahalulu Gulch (cf. Figure 5d). However, the highly evolved Hi Fe-Ti basalts have the highest REE contents. (d) Hamakua Volcanics - Waikahalulu Gulch: Fields for alkalic and tholeiitic postshield lavas from Waikahalulu Gulch (data from Frey et al. [1990]). LREE abundances increase from shield tholeiites to postshield tholeiites to postshield alkalic basalts, but these lava groups have similar HREE abundances.

enriched in P, Nb, Ta and the light REE but Sr is slightly depleted; also Ni and Cr are highly depleted, but Sc and V are slightly enriched (Table 6). The compatible behavior of Sr is consistent with major plagioclase fractionation, but the lack of a large negative Eu anomaly in these evolved basalts (Figure 5) must reflect a high  $\text{Eu}^{+3}/\text{Eu}^{+2}$  ratio in the magmas. The Sc and V enrichment in the high Fe-Ti basalts contrasts with the relative depletion in the postshield Laupahoehoe hawaiites which formed by segregation of a clinopyroxene-rich assemblage containing significant amounts (~10%) of Fe-Ti oxides [Frey et al., 1990]. In fact, an important aspect of the petrogenesis of the high Fe-Ti basalts is that Fe-Ti oxides were not important fractionating phases whereas Fe-Ti oxide segregation was important in generating the lavas forming the postshield hawaiite substage (Laupahoehoe Volcanics). Quantitatively, high Fe-Ti basalt/MU-5 enrichment factors for the most incompatible elements (P, Nb, LREE) are 1.5 to 1.6; i.e., very similar to the maximum enrichments (100/F) expected from the simple crystal fractionation model (Tables 5c and 6). We conclude that the major process forming the high Fe-Ti basalts was segregation of major amounts of

plagioclase and significant quantities of olivine and clinopyroxene from parental alkali basalts. However, the complexity of the phenocryst and xenocryst compositions in the high Fe-Ti lavas from Mauna Kea (Table 4) indicate that mixing processes also occurred.

In general, the high Fe-Ti Hamakua basalts have higher MgO contents than the younger Laupahoehoe hawaiites. In several compositional plots such as  $\text{K}_2\text{O}-\text{MgO}$ ,  $\text{Al}_2\text{O}_3-\text{MgO}$  and  $\text{Al}_2\text{O}_3/\text{CaO}-\text{MgO}$  [see Frey et al., 1990, Figure 8] the Hamakua alkalic and high Fe-Ti basalts and Laupahoehoe hawaiites define trends that could be interpreted as a single liquid line of descent. This interpretation is unlikely because (1) there is a compositional gap between the Laupahoehoe hawaiites and all of the Hamakua basalts [Frey et al., 1990], and (2) in some plots it is evident that the hawaiites and high Fe-Ti basalts define different fractionation paths rather than a single continuous trend (Figure 8).

Therefore the high Fe-Ti basalts in the Hamakua Volcanics (basaltic substage) provide evidence for important changes in the evolution of Mauna Kea volcano during the postshield stage. During the early part of postshield volcanism, alkalic

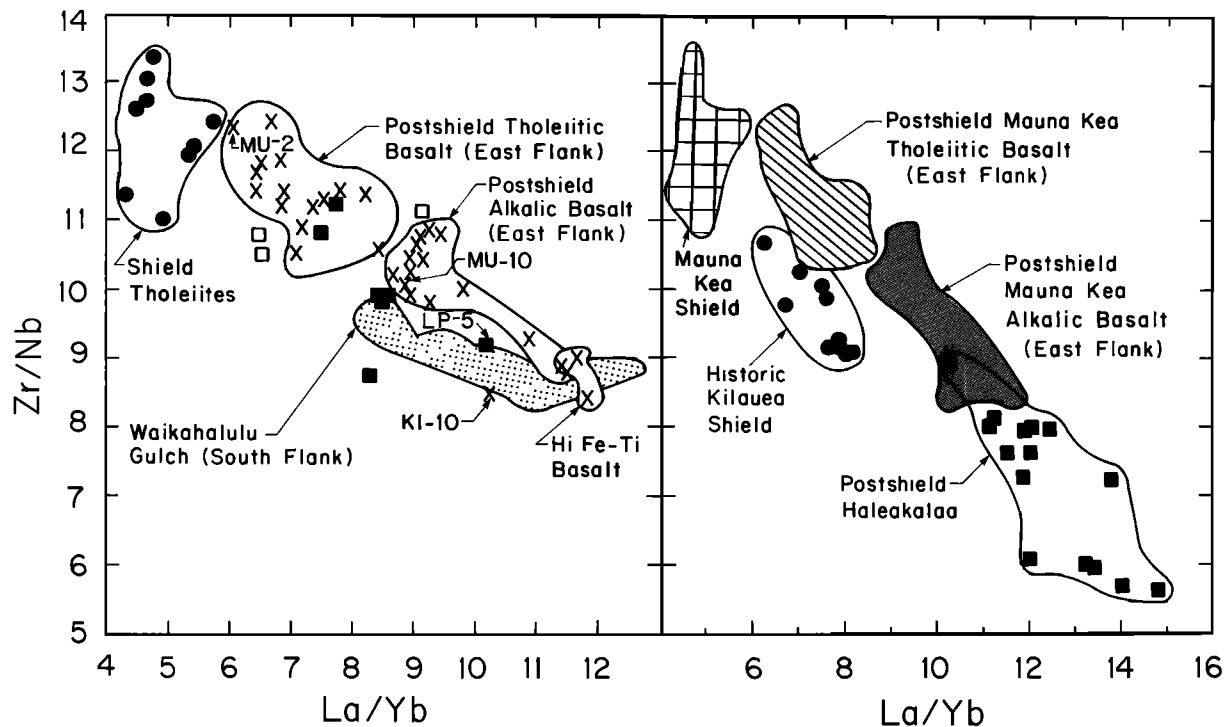


Fig. 6. Zr/Nb versus La/Yb for Mauna Kea shield lavas and the Hamakua Volcanics (note different La/Yb scales in the two panels). Because Nb and La are more incompatible than Zr and Yb, respectively, most basalt suites define inverse trends in this plot. (Left) Shield tholeiites have the lowest La/Yb and form a distinct field. Compared to the shield lavas, postshield tholeiites from the east flank have higher La/Yb and similar Zr/Nb. Postshield alkalic basalts from the east flank have lower Zr/Nb and higher La/Yb than shield or postshield tholeiites from the east flank. The evolved high Fe-Ti basalts have the most extreme values. Postshield basalts from Waikahalulu Gulch on the upper south flank also range from tholeiitic to alkalic (Figure 3a and Frey et al. [1990]; however, they are, in general, younger than postshield lavas from the east flank and they define a La/Yb-Zr/Nb field most similar to the alkalic postshield basalts. Labeled samples MU-10 and KI-10 (discussed in the text) are exceptions in that they are in the tholeiitic field in Figure 3a but they have higher La/Yb and lower Zr/Nb than other tholeiites from the east flank. The phenocryst-rich lavas (open and solid squares designating picrites and ankaramites, respectively) range widely in La/Yb and Zr/Nb. (Right) Comparison of fields for Mauna Kea lavas (shield tholeiites, postshield tholeiitic and alkalic basalts from the east flank) with fields for historic Kilauea tholeiites (data from Rhodes et al. [1989]) and postshield alkalic basalts (Kula Volcanics) from Haleakala Volcano (data from Chen and Frey [1985]). Important features are (1) The Mauna Kea shield lavas have higher Zr/Nb and lower La/Yb than Kilauea shield tholeiites; and (2) the postshield alkalic basalts from Mauna Kea generally have higher Zr/Nb and lower La/Yb than postshield alkalic basalts in the Kula Volcanics of Haleakala Volcano.

basalt magmas stagnated and fractionated at low pressures. Plagioclase fractionation was important, but segregation of Fe-Ti oxides was not a major process and residual melts were Fe-Ti rich basalts with  $Al_2O_3/CaO < 2.1$ . However, later alkalic basalt magmas stagnated and fractionated at moderate pressures, and segregation of a clinopyroxene-rich assemblage containing Fe-Ti oxides was a major process in creating the hawaiites ( $Al_2O_3/CaO > 2.3$ ) of the Laupahoe Volcanics [Frey et al., 1990].

What caused this difference in the petrogenetic role of Fe-Ti oxides? Although it is difficult to infer oxygen fugacities during the melt-solid segregation processes that led to the high Fe-Ti basalts and hawaiites, there is no evidence that these distinctive lava types formed at different oxygen fugacity. For example, Frey et al. [1990] found that coexisting oxides in both lava types reflect equilibration close to the QFM buffer; whole-rocks have similar  $Fe_2O_3/FeO$  ratios (five of the six high Fe-Ti basalts studied have  $Fe_2O_3/FeO$  of 0.33 to 0.46 and several Laupahoe hawaiites have similar ratios (Table 1 of Wolfe et al. [1991]), and none of these lavas have significant Eu anomalies (Figure 5 and Figure 7 of West et al. [1988]).

An alternative explanation for the different roles of Fe-Ti oxides in the two lava suites is that the Hamakua high Fe-Ti basaltic and Laupahoe hawaiitic magmas contained different concentrations of  $H_2O$ . In experimental studies of the liquid lines of descent of tholeiitic lavas, Spulber and Rutherford [1983] and Juster et al. [1989] found that at 1 atm under anhydrous conditions and a range of oxygen fugacities (QFM to  $-1W$ ), Fe-Ti oxide crystallization does not occur until  $TiO_2$  contents reach  $\sim 5\%$ ; i.e., similar to the maximum values in the high Fe-Ti basalts (Figure 8). Similar experimental results were obtained by Mahood and Baker [1986], who studied a more alkalic basalt. However, with a tholeiitic composition at low oxygen fugacities (graphite-methane) but under hydrous conditions ( $f_{H_2O} \sim 0.67 P_{fluid}$  1-2 kbar total pressure) Spulber and Rutherford [1983] found that Fe-Ti oxides were saturated at much lower,  $\sim 2\%$ ,  $TiO_2$  contents. Also, Baker and Egger [1983, Figure 3] showed that Fe-Ti oxides are near liquidus phases in hydrous (4-6 wt %  $H_2O$ ) basaltic melts at 2 kbar total pressure. Therefore a plausible explanation for the differing role of Fe-Ti oxides in the basaltic and hawaiitic substages is that in the moderate pressure magma chambers present during the hawaiite substage (see Figure 15

TABLE 4a. Olivine Compositions (wt.%) in Mauna Kea Lavas

Rock:	Submarine Shield Tholeiites														
	MK1-3			MK1-8						MK1-10					
	Ph	Ph		Ph	Ph		Ph	Ph	Xeno	Ph	Ph		Ph	Ph	Xeno
Type:	2.5	1.0		1.6	0.6		3.0	0.8	0.5	2.0					
Size: (mm):															
Part:	Core	Core	Rim	Core	Rim	Core	Rim	Core	Rim	Core	Rim	Core	Rim	Core	Rim
SiO <sub>2</sub>	40.4	38.5	39.2	40.1	39.2	39.6	38.9	40.4	39.4	39.8	39.9	39.3	39.7	40.5	39.8
FeO	9.35	19.8	16.8	12.7	17.1	14.9	18.0	12.3	16.7	13.9	14.9	13.4	15.2	10.9	15.2
MgO	49.3	40.7	43.4	46.2	42.9	45.0	42.3	46.7	43.4	46.2	44.3	47.2	45.0	48.1	44.9
Total	99.05	99.0	99.4	99.0	99.2	99.5	99.2	99.4	99.5	99.9	99.1	99.9	99.9	99.5	99.8
Fo%	90.4	78.5	82.2	86.6	81.7	84.4	80.8	87.2	82.2	85.5	84.1	86.3	84.1	88.7	84.0

Rock:	Submarine Shield Tholeiites										Postshield Tholeiitic Basalts					
	MK2-1			MK5-13			MK6-18				MU-2					
	Ph	Ph		Ph	Ph		Ph	Ph	Xeno	Ph	Ph	Mph				
Type:	2.0	3.0		2.5	1.5		1.2	0.8	2.0	0.7	0.2					
Size: (mm):																
Part:	Core	Core	Rim	Core	Core	Rim	Core	Rim	Core	Rim	Core	Rim	Core	Rim		
SiO <sub>2</sub>	40.5	40.1	39.2	40.3	39.2	39.5	40.6	40.3	40.6	40.4	41.0	40.1	39.3	36.1	37.6	36.8
FeO	9.7	13.3	16.5	10.25	11.45	15.0	10.0	11.6	10.3	10.8	9.3	13.6	18.8	37.5	27.7	32.5
MgO	48.8	45.6	43.4	48.5	43.7	44.8	48.6	47.6	48.0	48.3	49.5	46.2	41.2	25.4	34.6	31.0
Total	99.0	99.0	99.75	99.05	99.35	99.3	99.2	99.5	98.9	99.5	99.8	99.9	99.3	99.0	99.9	100.3
Fo%	90.0	86.0	82.4	89.4	82.6	84.2	89.7	88.0	89.3	88.9	90.5	85.8	79.6	54.7	69.0	62.9

Rock:	Postshield Tholeiitic Basalts				Postshield Alkalic Basalt						Postshield Picrite			
	MU-8				LP-9			MU-5		MU-10		LP-6		
	Xeno	Mph			Xeno	Ph	Mph	Ph	Mph	Xeno	Ph	Xeno	Ph	Ph
Type:	0.4	0.2			2.0	0.7	0.3	0.8	0.1	1.2	0.6	3.0	1.2	0.7
Size: (mm):														
Part:	Core	Rim	Core	Core	Core	Rim	Core	Core	Core	Core	Rim	Core	Core	Core
SiO <sub>2</sub>	39.7	39.9	40.4	39.5	39.8	39.2	39.3	39.3	37.7	38.7	39.5	39.5	39.5	40.0
FeO	14.6	14.7	13.7	17.1	15.1	18.2	17.9	16.8	29.4	21.3	15.7	16.9	14.6	14.5
MgO	44.8	44.6	45.6	42.8	44.5	42.2	42.5	43.3	34.5	40.1	44.2	43.6	45.8	45.0
Total	99.1	99.2	99.7	99.4	99.4	99.6	99.7	99.4	99.6	100.1	99.4	100.0	99.9	99.5
Fo%	84.5	84.5	85.6	81.7	84.0	80.5	80.9	82.1	66.3	77.0	83.4	82.1	84.9	84.7

Rock:	Postshield Ankaramites										
	MU-9					LP-5					
	Ph	Ph	Mph	Mph		Ph	Ph	Mph			
Type:	2.0	1.2	0.3	0.2		2.8	2.1	0.3			
Size:(mm):											
Part:	Core	Rim	Core	Core	Core	Core	Rim	Core	Rim	Core	Rim
SiO <sub>2</sub>	40.3	39.9	40.1	40.0	39.9	41.2	39.3	40.5	39.5	38.7	36.1
FeO	13.2	15.5	14.1	14.1	15.7	12.3	22.5	14.3	19.4	24.3	29.0
MgO	45.8	44.1	45.2	45.3	44.0	46.5	38.5	45.2	40.9	36.6	25.2
Total	99.3	99.5	99.4	99.4	99.6	100.0	100.2	100.0	99.8	99.6	100.3
Fo%	86.1	83.5	85.1	85.2	83.3	87.1	75.3	84.9	79.0	72.9	53.5

Rock:	Postshield High Fe-Ti Basalt											
	MU-6		MU-11					KI-12				
	Gm	Gm	Ph	Ph	Mph	Xeno	Mph	Mph	Xeno			
Type:	-	-	2.0	1.8	0.4	1.4	0.3	0.2	0.8			
Size:(mm):												
Part:	Core	Core	Core	Rim	Core	Rim	Core	Core	Rim	Core	Core	Core
SiO <sub>2</sub>	37.2	36.7	39.7	39.4	38.7	38.5	38.4	38.1	38.2	36.2	34.9	37.7
FeO	28.3	30.3	16.1	17.7	21.5	22.0	22.1	24.3	23.4	33.0	35.3	23.9
MgO	34.4	32.7	43.8	42.6	39.7	39.8	39.4	37.6	38.3	30.9	29.0	38.7
Total	99.9	99.7	99.6	99.7	99.9	100.3	99.9	100.0	99.9	100.1	99.2	100.3
Fo%	68.4	65.8	82.9	82.1	76.7	76.3	76.1	73.4	74.5	62.6	59.4	74.3

All data obtained by electron microprobe at the University of Hawaii.

Ph, phenocryst; Mph, microphenocryst; Xeno, xenocryst (grains, >0.5 mm, with strongly resorbed margins).

TABLE 4b. Clinopyroxene Compositions in (wt.%) Mauna Kea Lavas

Rock: Type: Size (mm): Part:	Submarine Shield Tholeiites					Postshield Tholeiitic Basalts			
	MK1-8		MK6-18		MU-2		MU-8		
	Ph 0.7	Mph 0.2	Mph 0.2	Mph 0.2	Mph 0.35	Mph 0.15	Core	Rim	
	Core	Rim	Core	Rim	Core	Core	Rim	Core	Rim
SiO <sub>2</sub>	51.2	51.4	51.7	51.5	51.2	52.25	52.71	51.74	49.88
TiO <sub>2</sub>	1.1	1.0	0.9	1.0	1.0	0.82	0.83	0.92	1.61
Al <sub>2</sub> O <sub>3</sub>	3.0	2.5	2.6	2.9	4.2	2.45	1.59	2.91	4.55
Cr <sub>2</sub> O <sub>3</sub>	0.5	0.4	1.0	0.6	0.7	0.13	0.27	0.52	0.74
FeO	8.5	9.3	6.9	6.9	6.9	6.83	8.81	6.76	7.38
MgO	16.5	16.5	17.3	17.0	16.7	17.08	17.48	17.72	15.83
CaO	18.6	18.0	19.0	18.6	18.8	19.98	17.78	18.48	19.44
Na <sub>2</sub> O	0.3	0.2	0.2	0.2	0.2	0.22	0.14	0.22	0.27
Total	99.7	99.3	99.6	98.7	99.7	99.76	99.61	99.27	99.70
Wo	38.5	37.3	39.2	39.1	39.6	40.7	36.3	38.2	41.2
En	47.7	47.7	49.7	49.6	49.0	48.4	49.7	50.9	46.6
Fs	13.8	15.0	11.1	11.3	11.4	10.9	14.0	10.9	12.2

Rock: Type: Size (mm): Part:	Postshield Ankaramites					Postshield High Fe-Ti Basalt				
	MU-9		LP-5			KI-12				
	Ph 0.75	Mph 0.35	Ph 1.2	Ph 0.8	Ph 2.0	Mph 0.3	Core	Rim	Core	
	Core	Rim	Core	Core	Rim	Core	Rim	Core	Rim	
SiO <sub>2</sub>	52.00	52.23	52.09	52.04	52.14	50.90	48.89	45.4	46.8	49.4
TiO <sub>2</sub>	0.67	0.69	0.73	0.65	0.82	1.18	1.92	3.2	2.8	1.6
Al <sub>2</sub> O <sub>3</sub>	2.39	2.63	2.52	2.50	2.78	2.68	4.50	7.3	6.0	3.3
Cr <sub>2</sub> O <sub>3</sub>	1.07	1.17	0.93	1.10	1.44	0.28	0.45	0.3	0.1	0.2
FeO	4.77	4.65	5.01	4.83	4.90	8.29	8.82	9.4	9.0	9.0
MgO	17.13	17.02	17.05	17.04	16.89	15.70	14.72	12.6	13.2	15.0
CaO	21.35	21.72	21.24	20.73	20.91	20.15	20.17	21.4	21.1	20.3
Na <sub>2</sub> O	0.24	0.21	0.24	0.25	0.29	0.30	0.27	0.6	0.5	0.4
Total	99.62	100.32	99.81	99.14	100.17	99.48	99.74	100.2	99.5	99.2
Wo	43.7	44.3	43.5	43.0	43.4	41.6	42.4	46.2	45.4	42.1
En	48.7	48.3	48.5	49.2	48.7	45.1	43.1	37.9	39.5	43.3
Fs	7.6	7.4	8.0	7.8	7.9	13.3	14.5	15.9	15.1	14.6

#See Table 4a footnotes.

of Frey et al. [1990]) the magmas did not lose H<sub>2</sub>O, and Fe-Ti oxides crystallized early in the liquid line of descent; in contrast, during the low pressure fractionation prevalent during the basaltic substage, magmas were devolatilized and at the resulting low H<sub>2</sub>O contents Fe-Ti oxide saturation did not occur until residual melts contained ~5% TiO<sub>2</sub>. Supporting evidence for evolution of the hawaiites and Fe-Ti basalts at different P<sub>H<sub>2</sub>O</sub> are the high K<sub>D</sub> ((Ca/Na)<sub>plag</sub>/(Ca/Na)<sub>WR</sub>) ratios (>1.5) in the hawaiites [West et al., 1988] and the much lower ratios, 0.9 to 1.2, for the alkalic and high Fe-Ti basalts (Tables 2 and 4). Juster et al. [1989, p. 9266] noted that lower K<sub>D</sub> ratios, ~1, are typical of 1 atm anhydrous experiments, but higher K<sub>D</sub> ratios are typical of moderate pressure (1-2 kbar) experiments with H<sub>2</sub>O-bearing melts.

#### 4.2. Origin of Postshield MgO-rich Lavas

Lavas with high MgO contents (>12%) are important because they may be near-primary magmas. In particular, picritic lavas with 12 to 25% MgO have been proposed as the primary magmas for Hawaiian volcanoes [e.g., Wright, 1984; Wilkinson and Hensel, 1988; Nicholls and Stout, 1988]. At other volcanoes, ankaramites have also been proposed as representative of primary magmas [e.g., Thompson and

Flower, 1971; Maaloe et al., 1986]. In the lava suite from the east flank of Mauna Kea there are 11 samples with high MgO contents (12.6% to 24%, Table 2). They are all porphyritic, ranging from the olivine-rich basalt MU-10 (8.5% olivine phenocrysts) to the picrites LP-6 and KI-3 (21% and 34% olivine phenocrysts, respectively), to ankaramites, such as MU-9, LP-4, -5, C-76, KI-4, -5 and -11 which contain abundant olivine and clinopyroxene phenocrysts (Table 3). In order to understand the origin of these phenocryst-rich lavas, it is necessary to determine if they represent crystallized high MgO melts or if they are mixtures of low MgO melts and accumulative or xenocrystic minerals.

An approach for identifying lavas that are mixtures of melts and accumulated minerals is to utilize the phenocryst compositions (Table 4). Specifically, the phenocryst core compositions can be compared with the mineral compositions that would form in equilibrium with a melt having the bulk rock composition. In samples MU-9, MU-10 and LP-6 the Fo contents of the olivines (Table 4) are too low (i.e., K<sub>D</sub> >0.4) to be in equilibrium with melts having the bulk rock composition (Figure 7). However, this observation does not preclude the hypothesis that these rocks represent a high MgO melt because the relatively low Fo contents may reflect equilibration with the evolved melts that formed the

TABLE 4c. Plagioclase Compositions in Mauna Kea Lavas

Rock:	Submarine Shield Tholeiites			Postshield Tholeiitic		Postshield Alkalic	
	MK1-8	MK6-18		MU-2		LP-9	
Type:	Mph	Mph	Mph	Mph	Mph	Ph	Mph
Size (mm):	0.2	0.2	0.1	0.35	0.20	1.0	0.3
Part:	Core	Core	Core	Core	Core	Core	Core
SiO <sub>2</sub>	53.4	51.9	52.4	47.00	47.00	50.45	50.75
Al <sub>2</sub> O <sub>3</sub>	28.7	29.5	29.3	33.20	33.10	31.05	33.75
FeO	0.8	0.9	0.8	0.70	0.65	0.50	0.80
CaO	11.9	13.3	12.9	16.65	16.50	13.85	13.65
Na <sub>2</sub> O	4.6	3.6	3.8	1.90	2.15	3.40	3.45
K <sub>2</sub> O	0.2	0.1	0.1	0.05	0.05	0.15	0.25
Total	99.6	99.3	99.3	99.50	99.45	99.40	99.65
An%	58.2	67.0	64.9	82.3	80.7	68.6	67.6

Rock:	Postshield Ankaramites				Postshield High Fe-Ti Basalts				
	LP-5		MU-9		MU-6				
Type:	Ph	Mph	Xeno	Mph	Ph		Mph		
Size (mm):	1.25	0.50	0.40	0.35	2		0.3		
Part:	Core	Rim	Core	Core	Core	Rim	Core	Rim	
SiO <sub>2</sub>	47.75	53.4	50.90	47.85	50.30	51.8	53.2	53.7	53.3
Al <sub>2</sub> O <sub>3</sub>	33.25	28.1	30.60	33.25	31.30	29.5	28.9	29.0	29.0
FeO	0.55	1.3	0.70	0.60	0.75	0.6	0.6	0.7	0.6
CaO	16.50	11.8	13.75	16.35	14.40	13.1	12.2	11.9	11.9
Na <sub>2</sub> O	1.90	4.3	3.10	2.00	3.10	3.8	4.1	4.4	4.4
K <sub>2</sub> O	0.07	0.3	0.15	0.09	0.15	0.24	0.29	0.34	0.33
Total	100.02	99.2	99.20	100.14	100.00	99.0	99.3	100.0	99.5
An%	82.4	59.2	70.4	81.5	71.3	64.9	61.2	58.8	58.9

Rock:	Postshield High Fe-Ti Basalts						
	MU-11			KI-12			
Type:	Ph		Ph		Mph		
Size (mm):	1.6		2.2		0.3		
Part:	Core	Rim	Core	Rim	Core	Core	
SiO <sub>2</sub>	47.4	47.4	51.4	50.6	51.5	52.6	53.1
Al <sub>2</sub> O <sub>3</sub>	33.0	32.8	30.1	30.5	29.8	29.4	29.4
FeO	0.6	0.6	0.7	0.6	0.8	0.7	0.9
CaO	17.2	16.5	13.4	14.2	13.6	12.7	12.0
Na <sub>2</sub> O	1.8	2.0	3.6	3.2	3.6	4.0	4.0
K <sub>2</sub> O	0.14	0.1	0.21	0.16	0.19	0.27	0.37
Total	99.9	99.4	99.4	99.3	99.5	99.7	99.8
An%	84.1	81.6	66.8	70.7	66.8	62.7	60.8

See Table 4a footnotes.

groundmass [e.g., Helz, 1987]. The ankaramite LP-5 contains normally zoned phenocrysts of olivine, clinopyroxene and plagioclase. Although the core compositions of olivine, plagioclase and the low Fe/Mg cpx in sample LP-5 are consistent with crystallization from a melt having the whole-rock composition (e.g., Figure 7), the diverse compositions of clinopyroxene cores reflect magmatic complexity (Table 4b).

Bulk rock compositions can also be used to identify cumulative lavas. Unlike crystallized melts, lavas with cumulative minerals have geochemical characteristics which reflect their mineralogy. For example, clinopyroxene phenocrysts in Mauna Kea lavas have much lower Al<sub>2</sub>O<sub>3</sub>/CaO (~0.1, Table 4b) ratios than Mauna Kea lavas (Figure 3b). Therefore, lavas with accumulative clinopyroxene should have distinctively lower Al<sub>2</sub>O<sub>3</sub>/CaO than most Mauna Kea basalts which have Al<sub>2</sub>O<sub>3</sub>/CaO ≥ 1.2, except for MU-8 a nearly aphyric tholeiitic basalt with Al<sub>2</sub>O<sub>3</sub>/CaO = 1.08 (Figure 3b). The seven ankaramites range widely, in Al<sub>2</sub>O<sub>3</sub>/CaO (0.71 to 1.11). All have Al<sub>2</sub>O<sub>3</sub>/CaO less than most of the basalts, but four ankaramites (C-76, LP-5, KI-4, KI-5) have ratios, 1.03 to 1.11, close to the lowest basalt value, 1.08 (Figure 3b). Although Sc is usually concentrated in clinopyroxene relative to coexisting melt and minerals [e.g., Hofmann et al., 1987], the Sc contents of these MgO-rich lavas (22.4 to 33.9 ppm)

overlaps the Sc range (27.6 to 36.4 ppm) in the Hamakua basalts (Figure 9). However, among the MgO-rich lavas the highest Sc contents, 32.2 to 33.9 ppm, are in ankaramites (LP-4, C-76, MU-9 and KI-11) with 10-15% clinopyroxene phenocrysts; three of these samples (LP-4, MU-9, and KI-11, Figure 9) also have relatively low Al<sub>2</sub>O<sub>3</sub>/CaO (<0.96, Figure 3b). Therefore, clinopyroxene appears to be an accumulative phase in some of these MgO-rich lavas.

In summary, we cannot convincingly demonstrate that any of the MgO-rich, postshield picrites and ankaramites are representative of high MgO melts. Based on petrography, phenocryst compositions and whole rock geochemistry these samples are best explained as mixtures of relatively evolved melts and accumulated phenocrysts. Comparisons of incompatible element ratios in the phenocryst-rich lavas with those in tholeiitic and alkalic lavas are useful in evaluating the parental magma for the picrites and ankaramites. For example, samples KI-11, LP-4, -5, C-76, MU-9 and MU-10 have low Zr/Nb and high La/Yb which clearly implies an alkalic heritage; in particular, sample LP-5 has the highest La/Yb and lowest Zr/Nb in the Laupahoehoe Gulch section, and it is clearly related to the alkalic lavas (Figure 6). In contrast, samples KI-3, -4, -5, and -8 have Zr/Nb and La/Yb ratios which overlap with the tholeiitic basalts (Figure 6).

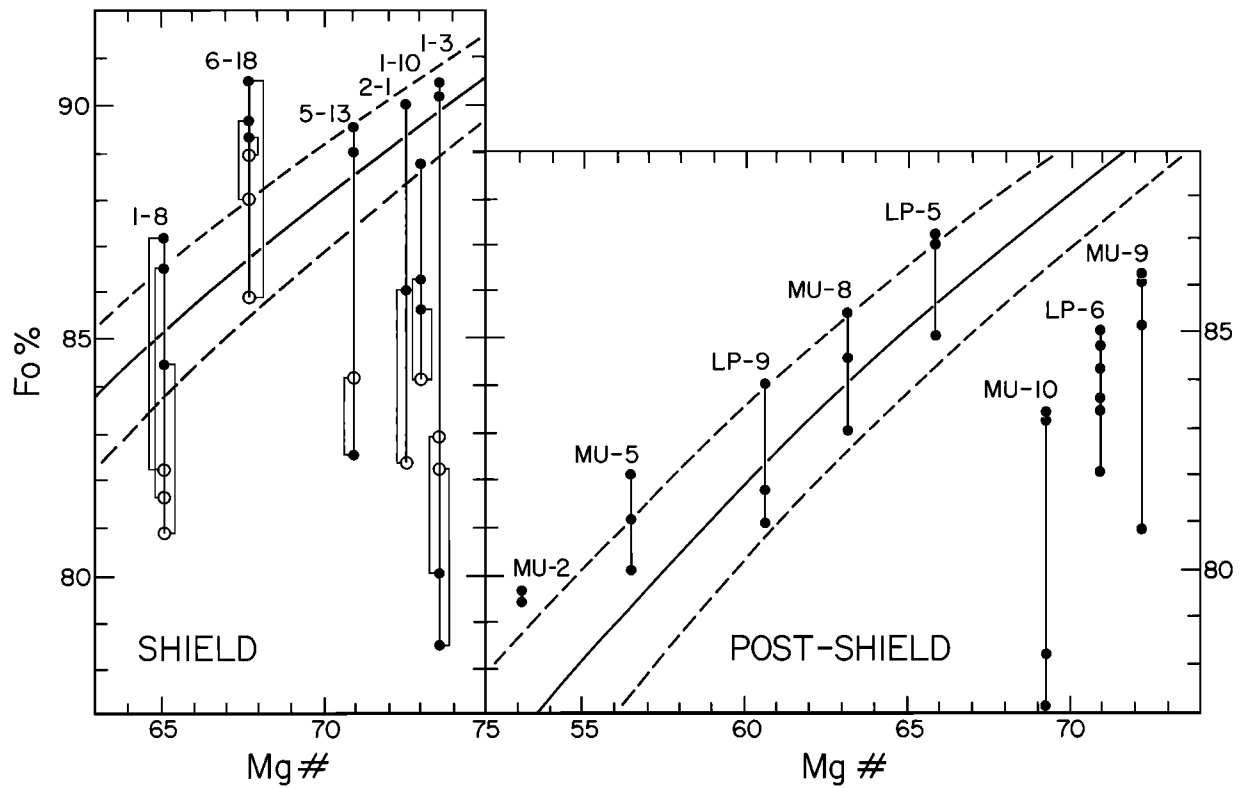


Fig. 7. Forsterite (Fo) content of olivine phenocrysts versus whole rock Mg# [ $100\text{MgO}/(\text{MgO} + \text{FeO})$ ] with  $\text{FeO}/(\text{FeO} + 0.9\text{Fe}_2\text{O}_3) = 0.9$  for Mauna Kea shield and postshield lavas. Solid circles indicate phenocryst cores; open circles are phenocryst rims. Thin tie lines connect the core and rim from the same phenocryst for the shield lavas. The heavy solid and dashed lines indicate the range for equilibrium olivine; i.e.,  $K_D = 0.30 \pm 0.03$ .

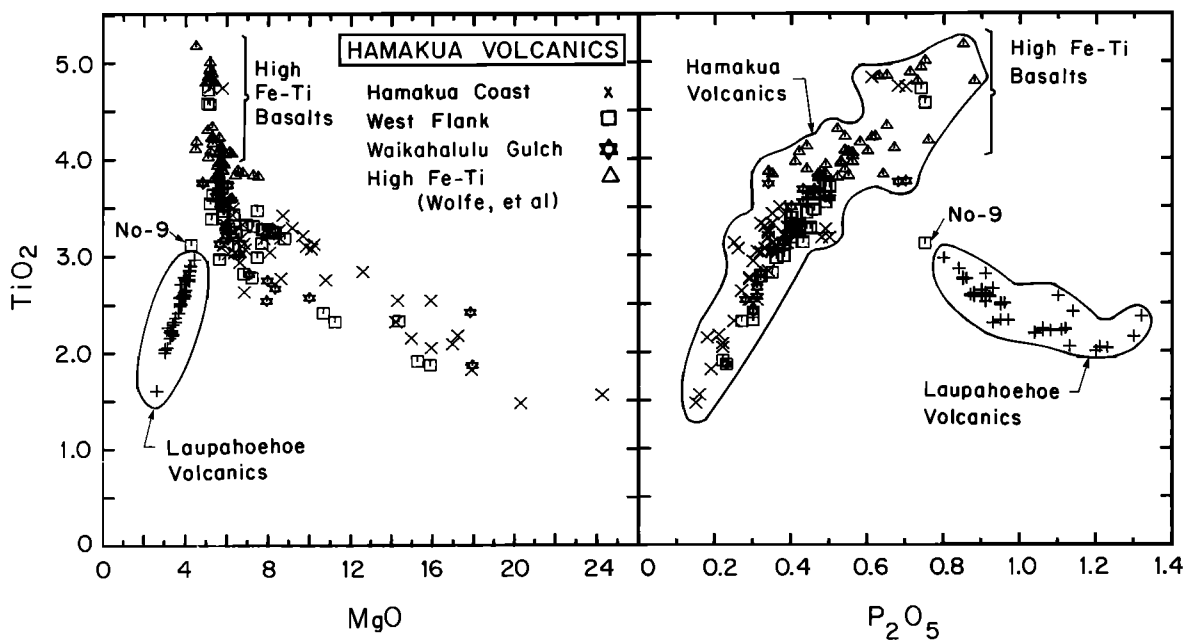


Fig. 8.  $\text{TiO}_2$  versus  $\text{MgO}$  and  $\text{TiO}_2$  versus  $\text{P}_2\text{O}_5$  (all in wt.%) for all analyzed postshield Mauna Kea lavas in this paper [West et al., 1988; Frey et al., 1990]. In addition, 39 other analyses of Mauna Kea lavas with more than 3.8%  $\text{TiO}_2$  are plotted (open triangles, data from compilation in Table 1 of Wolfe et al. [1991]). These high Fe-Ti basalts form the evolved end of the postshield Hamakua Volcanics and define a differentiation trend that is distinctly different from the trend defined by the Laupahoehoe Volcanics (hawaiites and mugearites).



TABLE 5a. Mineral Compositions Used in Deriving High Fe-Ti Basalts, MU-6, MU-11, and KI-12, From Alkalic Basalt MU-5

	Olivine		Plagioclase			Clinopyroxene	Ti-Magnetite
	1	2	3	4	5	6	7
SiO <sub>2</sub>	37.61	38.10	52.84	51.33	54.08	49.80	-
TiO <sub>2</sub>	-	-	-	-	-	1.67	25.2
Al <sub>2</sub> O <sub>3</sub>	-	-	29.64	30.80	28.83	3.26	2.1
FeO	23.81	24.30	1.05	0.55	0.87	9.09	69.8
MgO	38.60	37.60	-	-	-	15.15	2.2
CaO	-	-	12.01	13.61	11.79	20.51	-
Na <sub>2</sub> O	-	-	4.09	3.48	4.43	0.38	-
K <sub>2</sub> O	-	-	0.37	0.23	-	-	-

Column 1, xenocrystic olivine (Fo 74.3) in KI-12. The  $K_D$  for olivine/KI-12 is 0.28; 2, core of resorbed olivine phenocryst (Fo 73.4) in MU-11.  $K_D$  for olivine/MU-6 is 0.30; 3, microphenocryst in KI-12 (An 60.5),  $K_D$  (Ca/Na) for plag/KI-12 is 0.95; 4, phenocryst core in MU-6 (An 67.4),  $K_D$  (Ca/Na) for plag/MU-6 is 1.18 and for plag/MU-11 is 1.24; 5, phenocryst core in MU-6 (An 58.3),  $K_D$  (Ca/Na) = 0.80; 6, microphenocryst in KI-12,  $K_D$  for cpx/KI-12 and cpx/MU-6 are 0.27 and 0.28, respectively; and 7, magnetite in high Fe-Ti basalt from Table 2 of Frey et al., [1990].

TABLE 5b. Results of Least Squares Crystallization Models for Deriving High Fe-Ti Basalts From Alkalic Basalt MU-5

	MU-5 Observed	MU-6 Model	MU-11 Model	KI-12 Model
SiO <sub>2</sub>	47.12	47.05	47.46	47.17
TiO <sub>2</sub>	3.37	3.40	3.42	3.42
Al <sub>2</sub> O <sub>3</sub>	14.70	15.13	14.74	14.71
FeO	12.45	12.64	12.49	12.44
MgO	8.14	8.27	8.19	8.14
CaO	10.44	10.75	10.48	10.42
Na <sub>2</sub> O	2.88	2.62	2.34	2.80
K <sub>2</sub> O	0.91	-	0.81	0.84
$\Sigma r^2$		0.47	0.42	0.02

MU-6 model used olivine (2), plagioclase (5) and cpx (6) in Table 5a. (K<sub>2</sub>O was not used because K<sub>2</sub>O abundance in MU-6 was decreased by alteration). MU-11 model used olivine (2), plagioclase (4) cpx (6) and Ti-magnetite (7) in Table 5a. KI-12 model used olivine (1), plagioclase (3) cpx (6) and Ti-magnetite (7) in Table 5a.  $\Sigma r^2$  = sum of squares of residuals; i.e., model-observed.

TABLE 5c. Phase Proportions (wt.%) for Models

	MU-6 Model	MU-11 Model	KI-12 Model	Mahood-Baker*	Ol-Gabbro Xenoliths†
olivine	6.4	7.4	8.7(20.3)‡	7(20%)	5-32%
plagioclase	15.5	15.9	21.1(49.2)	18(51%)	20-51%
clinopyroxene	13.1	13.8	11.3(26.3)	10(29%)	30-56%
ti-magnetite	-	0.3	1.8(4.2)		
residual melt (F)	65.0	62.6	57.1	66	
100/F§	1.54	1.60	1.75		

\*Experimental results at 1 atm [Mahood and Baker, 1986] for deriving an Fe-Ti rich hawaiite (SI-1) from an alkali basalt (PSU-17). Mineral proportions in wt.%. These compositions are quite similar to KI-12 and MU-5, respectively (cf. Table 2, this paper with Table 2 of Mahood and Baker).

†Phase proportions (vol.%) for 12 olivine gabbro xenoliths interpreted as cumulates [Fodor and Van der Meyden, 1988].

‡Values within parentheses wt.% of phase in solid aggregate.

§This is the enrichment factor for a perfectly incompatible element; cf. with values in Table 6.

Apparently, these MgO-rich postshield lavas containing accumulative phases formed from diverse melt compositions ranging from alkalic to tholeiitic.

#### 4.3. Primitive Mantle Normalized Abundance of Incompatible Elements: Implications for Petrogenesis and Source Composition

Relative to an estimated primitive mantle composition, the incompatible element abundances in Mauna Kea lavas are

typical of oceanic island basalts (Figure 10); Y, Yb and Lu have lower abundances than MORB, and from Ti to Ta the most incompatible elements are increasingly enriched. (The deviations from a smooth trend at P, Sr and Pb are discussed separately). Qualitatively these differences between MORB and Mauna Kea basalts probably reflect derivation of Hawaiian basalts from a source that is enriched in incompatible elements relative to the MORB source and the equilibration of Hawaiian basalts with residual garnet [e.g., Hofmann et al., 1984; Chen and Frey, 1985; Budahn and Schmitt, 1985]. Also in the

TABLE 6. Trace Element Enrichment Ratios (Derived Melt/Parental Melt)

	MU-6/MU-5	MU-11/MU-5	KI-12/MU5
K <sub>2</sub> O	0.76	1.36	1.49
Rb	0.22	1.36	1.48
P <sub>2</sub> O <sub>5</sub>	1.32	1.49	1.66
Sr	0.98	0.84	0.96
Ba	1.39	1.39	1.32
Y	1.36	1.30	1.30
Zr	1.46	1.26	1.36
Hf	1.43	1.28	1.43
Nb	1.63	1.48	1.61
Ta	1.55	1.47	1.61
Th	1.73	1.31	1.41
La	1.62	1.47	1.59
Ce	1.51	1.34	1.57
Nd	1.50	1.36	1.56
Sm	1.51	1.36	1.44
Eu	1.46	1.35	1.34
Tb	1.80	1.53	1.40
Yb	1.37	1.26	1.32
Lu	1.36	1.33	1.33
Sc	0.98	1.15	1.12
V	1.15	1.44	1.44
Ni	0.59	0.44	0.42
Cr	0.24	0.23	0.22

K<sub>2</sub>O and Rb are anomalously low in sample MU-5 because of alkali loss during alteration. Th data for MU-11, MU-5 and MU-6 from Kennedy et al. [1991]. The lack of enrichment in Sr, Ni and Cr are consistent with segregation of plagioclase, olivine and clinopyroxene. The anomalously high enrichments of Th and Tb in MU-6 are not understood but they may reflect analytical uncertainties.

HREE region of Figure 10 there is complete overlap between tholeiitic and alkalic basalts from Mauna Kea (see also Figure 5); however, the alkalic basalts have higher concentrations of the more incompatible elements, Ti through Rb (Figure 10). Qualitatively, these differences are consistent with derivation of these basalts from a garnet-bearing source with the alkalic lavas resulting from a smaller degree of melting.

Abundance ratios of highly incompatible elements will not be affected by segregation or accumulation of olivine and pyroxene. In Figure 11 abundance ratios for elements with very similar incompatibility (normalized to the primitive mantle ratios of Sun and McDonough [1989]) are shown as a function of incompatible element abundance. Variations in these ratios and comparisons with primitive mantle estimates have important implications for the petrogenesis of Mauna Kea lavas.

*Sr/Nd and Ti/Eu.* The elements Sr and Ti are compatible in plagioclase and Fe-Ti oxides, respectively. Because these are important crystallizing phases during the crustal evolution of Mauna Kea lavas, these elements can be used to identify samples that have experienced significant amounts of crystal fractionation within the crust. For example, Sr/Nd ratios are slightly higher than the chondritic ratio in most Mauna Kea basalts, but with increasing differentiation, i.e., increasing P<sub>2</sub>O<sub>5</sub> content, Sr/Nd decreases with relatively low Sr/Nd in the high Fe-Ti basalts. This trend reflects segregation of plagioclase. Surprisingly, the MgO-rich dredged tholeiites, also define a trend to low Sr/Nd with increasing P<sub>2</sub>O<sub>5</sub> content (sample 1-8 contains 10.9% MgO, but it has the lowest Sr/Nd ratio).

Sun and McDonough [1989] found that Ti/Eu ratios in most basalts are ~7740, and this ratio is close to the mean value for

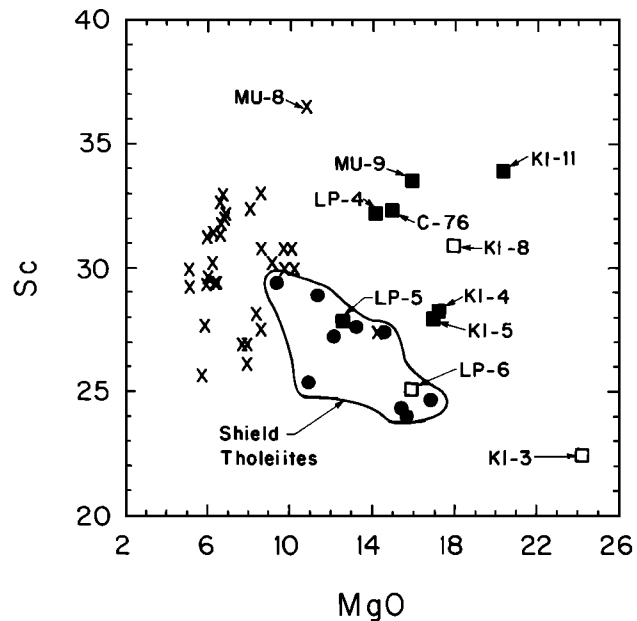


Fig. 9. Sc (ppm) versus MgO (wt.%) for shield lavas and Hamakua Volcanics from the east flank (same symbols as in Figure 3). The cpx-rich lavas (ankaramites are solid squares) have a Sc abundance range similar to the basaltic lavas.

Mauna Kea lavas. However, seven samples have anomalously low TiO<sub>2</sub>/Eu, a dredged shield tholeiite, 1-8; three alkalic Hamakua basalts from Laupahoehoe Gulch (LP-1, -7 and -8) and three evolved, alkalic Hamakua lavas (H17 and 17B from the upper part of Waikahalulu Gulch and NO-9 from the west flank [Frey et al., 1990]). The simplest interpretation is that Fe-Ti oxides segregated from these samples. This is a surprising result for the MgO-rich shield tholeiite 1-8 and the three alkalic basalts from Laupahoehoe Gulch (7.8 to 8.6% MgO). However, the anomalously low Ti/Eu are accompanied by relatively low Sr/Nd. Therefore, we conclude that these relatively high MgO samples have had a complex crustal history, perhaps mixing of relatively primitive and highly evolved lavas. The resorbed sodic plagioclase in shield sample 1-8 (Tables 3 and 4c) is consistent with a mixing process.

*Nb/La, Nb/U, Nb/Th and Ce/Pb* (Data for U, Th and Pb determined by isotopic dilution are in Kennedy et al. [1991]). All of the shield and postshield basalts have Nb/La higher than the bulk earth/chondritic ratio of 1.04; an exception is evolved Hamakua lava MU-17 which is anomalously enriched in La (La/Ce = 0.51). In general, relatively high abundances of Nb and Ta but lower abundances of Th and U are typical of Mauna Kea basalts (Figure 10) and OIB in general [e.g., Hofmann et al., 1986]. Ratios of Nb/U are greater than chondritic in all Mauna Kea lavas, but they range widely, 35 to 110, because of U loss during low temperature alteration [Kennedy et al., 1991]. Ratios of Nb/Th which are less affected by alteration are more uniform, 13.0 to 17.7, in the 10 samples analyzed for Th by isotope dilution [Kennedy et al., 1991] and overlap the average (15.2) for Hawaiian lavas reported by Hofmann [1986]. These values are much higher than the chondritic value of 8.6. In addition, the marked relative depletion of Pb in postshield Mauna Kea lavas (Figure 10) reflects the Pb depletion that is characteristic of OIB and MORB [Hofmann et al., 1986]. In fact, Ce/Pb ratios in most Mauna Kea lavas are even higher, ~40, [Kennedy et al., 1991], than in most OIB, 25±5 and much higher than estimates, ~9, for primitive mantle [Hofmann et al., 1986].

*Zr/Hf, Sm/Hf and Sm/Zr.* All of the Mauna Kea lavas have

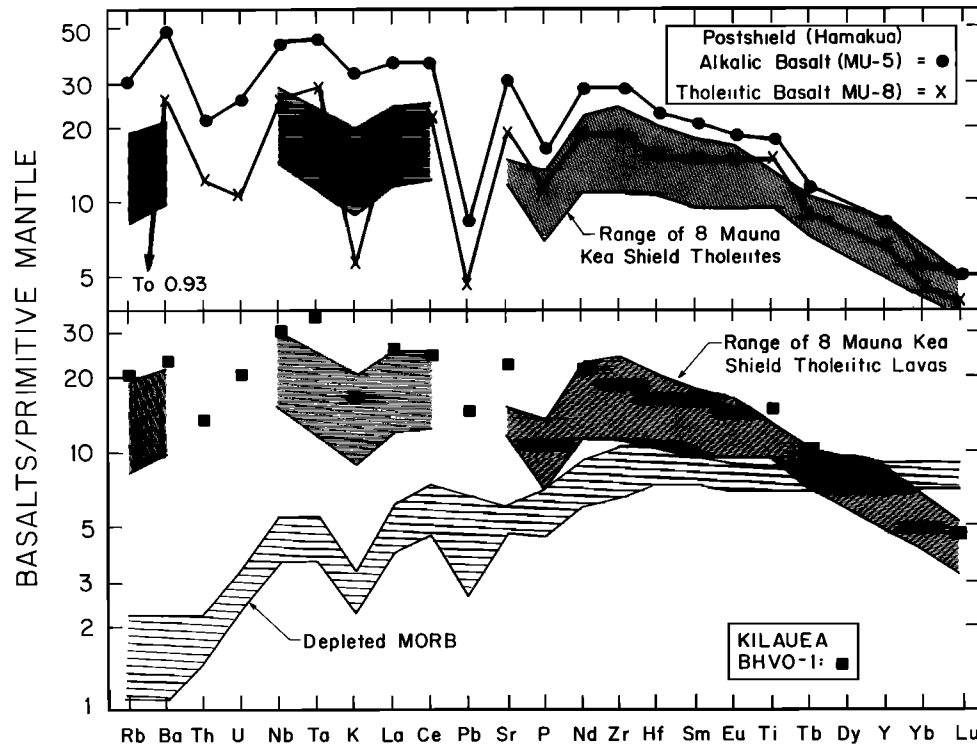


Fig. 10. Incompatible element abundances in Mauna Kea lavas normalized to primitive mantle estimates [Hofmann, 1988]. Lower panel shows field for shield tholeiites compared with a MORB average from Hofmann [1988]. The shield lavas define a generally convex-upwards pattern with valleys at K and P. Shown for comparison are data for BHVO-1 from Kilauea Volcano. Upper panel shows patterns for typical Hamakua Volcanics, alkali basalt MU-5, and tholeiitic basalt MU-8. The K and Rb depletions in MU-8 reflect alkali loss during postmagmatic alteration. Data for Th, U, and Pb from Kennedy et al. [1991].

higher Zr/Hf than the primitive mantle/chondritic ratios of ~36; specifically, the seven shield samples range in Zr/Hf from ~39 to 43 and Hamakua lavas range from ~41 to 47. Although a systematic analytical error is possible, Budahn and Schmitt [1985] report a similar range (average 45.1) for Mauna Kea lavas. Ratios of Sm/Hf in the shield and postshield lavas are very similar and their average is close to the primitive mantle/chondritic ratio of 1.44. In contrast, all Mauna Kea shield and postshield lavas have Sm/Zr ratios ( $0.033 \pm 0.002$  in the shield samples) less than the primitive mantle/chondritic ratio of 0.040. Therefore, the relatively high Zr/Hf ratio in Mauna Kea lavas is caused by relatively high Zr contents. Either the source of Mauna Kea lavas has relatively high Zr/Hf or Hf was preferentially retained in the residue during partial melting. The Kilauea sample BHVO-1 used as a geochemical standard also has  $Sm/Zr \leq 0.035$  (e.g., Table 1d of Frey et al. [1990]).

**P/Nd.** The bulk Earth abundance of P is difficult to estimate. However, Sun and McDonough [1989], report that  $P/Nd = 74 \pm 13$  in MORB and OIB. In contrast, P/Nd is less than 70 in all Mauna Kea lavas; e.g., P/Nd is 48.8 to 58.7 in the eight shield tholeiites. Kilauea sample BHVO-1 also has a low P/Nd ratio (47.3); therefore, the relatively low P/Nd in Mauna Kea lavas is not an analytical bias.

**Ba/Rb and Ba/Nb.** The Ba/Rb range, 12.0 to 14.6 in seven shield tholeiites lies between the mean MORB ratio 11.4 [Hofmann and White, 1983] and the average value  $14.7 \pm 1.4$  given by Hofmann [1986] for Hawaiian tholeiitic basalts. These values are close to the primitive mantle estimate of 11.3 [Hofmann and White, 1983]. Also, Ba/Nb ratios in the postshield lavas are very similar to primitive mantle estimates.

In summary, Sr/Nd and Ti/Eu ratios show that some high MgO lavas are probably mixtures containing an evolved

component. However, the range of Sr/Nd, Ti/Eu, Ba/Nb, Sm/Hf and La/Ce in Mauna Kea basalts encompasses the estimates for primitive mantle, and most basalts have ratios that are within 20% of primitive mantle values. In contrast, the range of other ratios, such as Nb/La, Sm/Zr, Zr/Hf, P/Nd, Nb/Th and Pb/Ce, in Mauna Kea lavas does not encompass estimates for primitive mantle. These differences require a non-primitive mantle source or changes in abundance ratios caused by partial melting.

#### 4.4. Shield Tholeiites: Implication for Primary Magma Composition and Relationship to Postshield Tholeiites

The submarine lavas dredged from the Mauna Kea shield are MgO-rich tholeiites (9.4 to 16.8% MgO, Table 1) containing abundant olivine phenocrysts and microphenocrysts (9.8 to 27.8%, Table 3) and relatively low MgO glasses (5.3 to 6.7% [Garcia et al., 1989]). Several of these samples represent mixed magmas. Evidence for a mixing component with a relatively evolved composition is the wide range of olivine-phenocryst compositions (cores range from 78.5 to 90.4% Fo) and the reversely zoned olivine phenocrysts in samples 1-3 and 5-13 (Table 4a). Also sample 1-8 contains sodic (An<sub>58</sub>), resorbed plagioclase (Table 4c), and the whole rock has the most extreme composition; compared to the other shield lavas it is anomalously enriched in incompatible elements (Figure 4), has anomalously low Ti/Eu and Sr/Nd (Figure 11) but relatively high  $Al_2O_3/CaO$  (e.g., Figure 12). These are characteristics of an evolved basalt which has fractionated plagioclase, clinopyroxene and Fe-Ti oxides. Also the glass in sample 1-8 has relatively low MgO (5.3%) and high  $K_2O$ ,  $P_2O_5$ , and  $H_2O$  contents (Garcia et al., 1989).

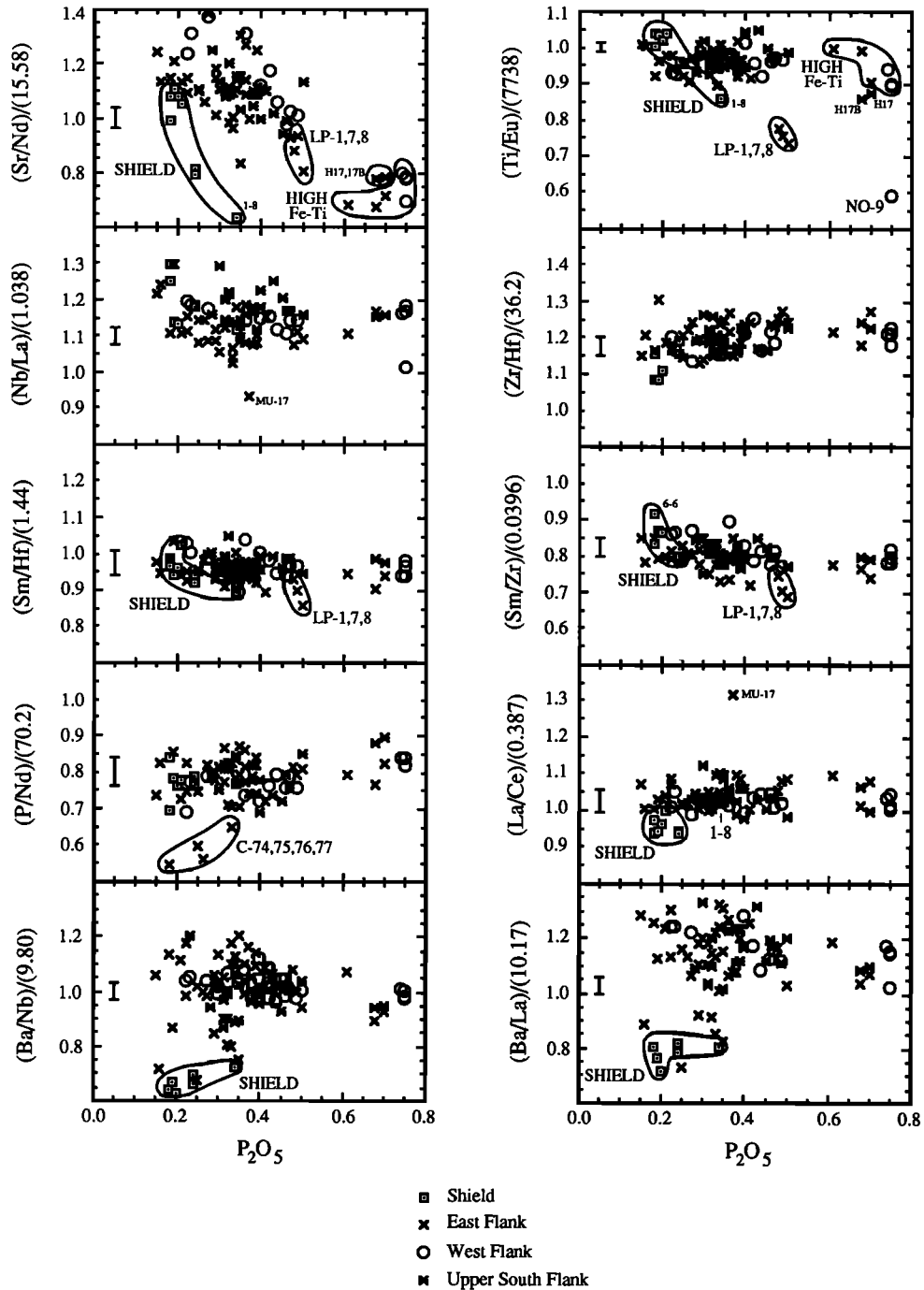


Fig. 11. Normalized abundance ratios of elements with very similar incompatibility (in a peridotite-melt system) versus  $P_2O_5$  content (wt. %) in shield lavas from the submarine east rift and basaltic postshield lavas from the east, west and upper south flanks. Normalization ratios (see vertical axis) are estimates of primitive mantle ratios from Sun and McDonough [1989]. Important features are that many Mauna Kea lavas have Sr/Nd, Ti/Eu, Nb/La, Sm/Hf, La/Ce and Ba/Nb ratios similar to primitive mantle ratios. However, the shield lavas have relatively low Ba/Nb, Ba/La and La/Ce (except for sample 1-8, which also has the lowest Sr/Nd and Ti/Eu). In general, alkalic postshield basalts have  $>0.35\%$   $P_2O_5$ . Alkalic samples LP-1, -7 and -8, have relatively low Sr/Nd and Ti/Eu. Sample MU-17 has an anomalously high La content (La/Ce, not plotted, is 0.51), and it may be contaminated. The low P/Nd for C74, C75, C76 and C77 reflects anomalously low  $P_2O_5$  data from Macdonald and Katsura [1964]. Data for shield and east flank Hamakua Volcanics from this paper; data for Hamakua Volcanics from the west flank and Waikahalulu Gulch from Frey et al. [1990] except for the following data (ppm) which were incorrectly reported in Table 1b of Frey et al. [1990] [sample 65-3: Ba = 226; sample H-14: Sr = 557, Ba = 302, Th = 1.9, sample 65-4: Sr = 587, sample H-15: Ba = 324, Nb = 31.4]. Error ranges indicate  $\pm 1$  standard deviation based on replicate analyses of BHVO-1.

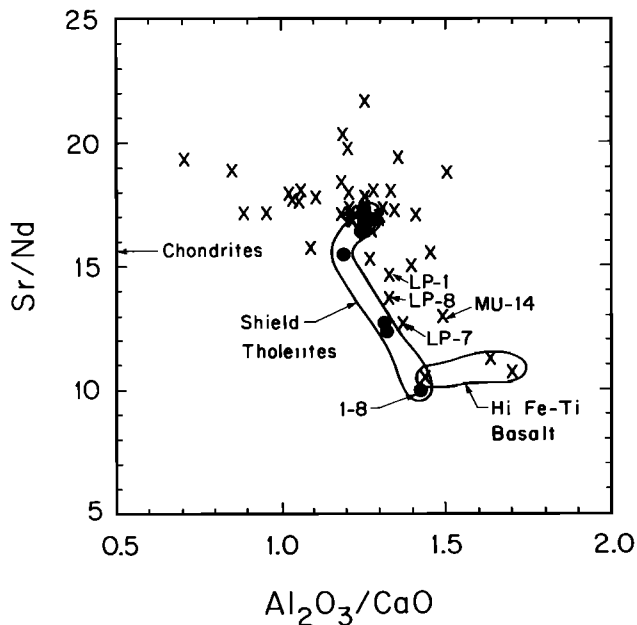


Fig. 12. Sr/Nd versus  $\text{Al}_2\text{O}_3/\text{CaO}$  for dredged shield samples (solid circles) and all Hamakua Volcanics from the east flank (crosses). The most evolved lavas have low Sr/Nd and high  $\text{Al}_2\text{O}_3/\text{CaO}$ , and this trend reflects plagioclase and clinopyroxene control.

We conclude that the bulk rock is a mixture of an evolved magma and olivine.

The MgO content of the most primitive melt can be inferred from the most forsterite-rich olivine (90.5 in sample 6-18, Table 4a) by using an olivine/melt  $K_D$  for Fe/Mg (e.g., 0.3) and assuming an FeO content for the primitive melt equal to that in sample 6-18. The latter assumption is reasonable because FeO contents of basaltic lavas are not sensitive to accumulation or fractionation of olivine [e.g., Langmuir and Hanson, 1980, p. 392]. As an example, the total Fe content in the eight shield lavas varies only by 9% (relative) as MgO varies from 9.4 to 16.9% (Table 1). The inferred MgO content of the equilibrium lava for  $F_{0.5}$  olivine ranges from 14.4% to 17.0% for a range in  $\text{FeO}/(\text{FeO}+0.9\text{Fe}_2\text{O}_3)$  from the commonly assumed value of 0.85 to the maximum value of 1. This MgO range is similar to estimates for Hawaiian primary tholeiitic magmas (12–15% for Mauna Loa, Kilauea, and Kauai [Wilkinson and Hensel, 1988; Nicholls and Stout, 1988; Maaloe et al., 1989 and 16.6% for Haleakala Volcano [Chen, 1990]).

We have  $^{87}\text{Sr}/^{86}\text{Sr}$  data for only two shield samples, 0.70365 for 1-8 and 0.70350 for 6-6. These values are significantly different, but they overlap the range found in Hamakua lavas (0.70336 to 0.70361 [Kennedy et al., 1991]). Some abundance ratios involving elements of very similar incompatibility in a peridotite-melt system are also similar in the shield (except for 1-8) and postshield basalts; e.g., Ti/Eu, Nb/La, Sm/Zr and Sm/Hf (Figure 11). However, the shield lavas have lower Ba/Nb, Ba/La and La/Ce (except for 1-8, Figure 11). These ratios in the shield lavas are lower than chondritic ratios whereas the postshield lavas have chondritic (Ba/Nb) or greater than chondritic ratios (Ba/La and La/Ce). Thus, the shield lavas have a more depleted character. These differences in incompatible element ratios must reflect differences in source composition or a range in degree of melting. The Sr isotopic similarity of shield and postshield lavas provides no evidence for different source components, but this inference must be evaluated with more isotopic data for the shield samples. If the same source components are

involved, the different incompatible element abundance ratios require that the shield lavas were derived by higher degrees of melting.

Possibly the sequence shield tholeiites-postshield tholeiites-postshield alkalic basalts represents a systematic decrease in degree of melting of a compositionally homogeneous source. For example, relative to postshield tholeiitic basalts the alkalic basalts have substantially higher La/Yb and Nb/Zr, (Figure 6) and similar La/Ce ratios (Figure 5). These differences are consistent with differences in peridotite/melt partition coefficients, and the hypothesis that the alkalic basalts formed by lower degrees of melting. Surprisingly, the differences between the shield and postshield tholeiitic lavas in Ba/Nb, Ba/La and La/Ce are much greater than the differences between postshield tholeiitic and postshield alkalic lavas (Figure 11). Because the largest change in abundance ratios of incompatible elements must occur at the lowest degrees of melting (e.g., Figure 13 of Chen et al. [1991]), the large differences in Ba/Nb, Ba/La and La/Ce between tholeiitic shield and postshield basalts require that the postshield tholeiitic basalts formed by significantly lower degrees of melting than the shield tholeiites. Alternatively, the postshield and shield basalts were derived from compositionally different sources.

#### 4.5. Relationship Between Tholeiitic and Alkalic Mauna Kea Lavas

The relationship between tholeiitic and alkalic basalts is a classical problem in igneous petrology. Recent studies of Hawaiian lavas bear directly on this problem. For example, at Haleakala Volcano, Chen and Frey [1985] and Chen et al. [1991] found Sr and Nd isotopic differences between tholeiitic basalts (Honomanu Basalts) and the overlying postshield alkalic basalts (Kula Volcanics). However, at Mauna Kea Volcano the oldest subaerial postshield lavas are intercalated tholeiitic and alkalic basalt which have overlapping radiogenic isotopic ratios. This was first demonstrated for Pb isotopes by Tatsumoto [1978], who found that Pb isotopic ratios were the same within analytical error for two postshield lavas, an alkalic basalt (C-78) and tholeiitic basalt (C-74) from Laupahoehoe Gulch. More recently, Kennedy et al. [1991] showed that intercalated tholeiitic and alkalic postshield Mauna Kea lavas have very similar Sr, Nd and Pb isotopic ratios. Isotopically similar alkalic and tholeiitic basalts have also been found at Kohala Volcano (Pololu Basalt [Feigenson et al., 1983; Lanphere and Frey, 1987; Hofmann et al., 1987]) and Haleakala Volcano (Honomanu Basalt [Chen et al., 1991]).

Comparisons of incompatible element abundances in phenocryst-poor tholeiitic and alkalic Mauna Kea basalts with very similar isotopic ratios shows that the alkalic basalts have higher abundances of incompatible elements and MgO (Figure 13). Therefore, the incompatible element enrichments of the alkalic basalts were not caused by post-melting, crystal-melt fractionation. The simplest interpretation of these data are that these postshield tholeiitic and alkalic lavas formed from a compositionally homogeneous source by different degrees of melting. If this is a valid hypothesis, then the compositions of the lavas may be used to constrain the composition and mineralogy of the sources of the postshield lavas [e.g., Feigenson et al., 1983].

However, these Hamakua lavas are not primary basalts (i.e., postmelting processes have caused significant compositional changes). Thus the effects of post-melting processes, such as fractional crystallization, must be evaluated before the lava compositions can be inverted to constrain source characteristics. In order to minimize the uncertainties resulting from the effects of fractional crystallization, we focus on abundance ratios of highly incompatible elements in lavas with  $\text{Al}_2\text{O}_3/\text{CaO}$  ratios ranging from 1.15 to 1.40 and assume that these samples have primarily been affected by olivine fractionation. (Most Hawaiian tholeiites have  $\text{Al}_2\text{O}_3/\text{CaO}$  in

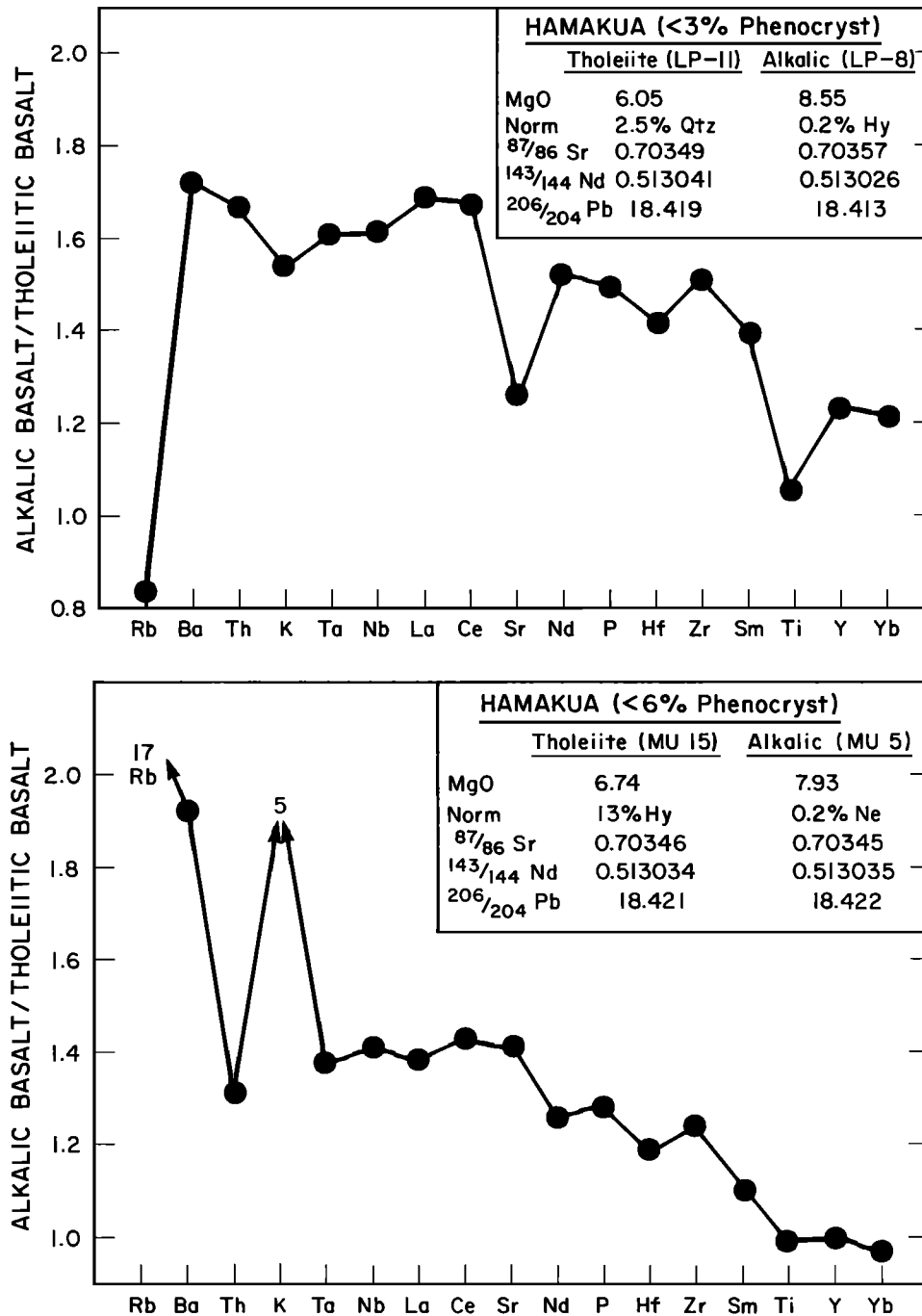


Fig. 13. Incompatible element abundances in pairs of postshield tholeiitic and alkalic basalts (Hamakua Volcanics) with similar isotopic ratios (isotopic data from Kennedy et al. [1991]). In each case, the alkalic basalt has higher abundances of incompatible elements and a higher MgO content. The anomalous points for K and Rb reflect postmagmatic alteration.

this range; e.g., in a summary of average tholeiitic basalt compositions for 14 Hawaiian volcanoes, Kirkpatrick et al. [1980] found a range of 1.22 to 1.59. Also the chondritic ratio, perhaps characteristic of the primitive mantle source component is 1.23.) Within this Al<sub>2</sub>O<sub>3</sub>/CaO range there are 26 Hamakua lavas varying in MgO from ~6% to 10%. Mauna Kea lavas with obvious complex or extensive crustal history, such as the ankaramites and high Fe-Ti basalts, have lower or higher Al<sub>2</sub>O<sub>3</sub>/CaO, respectively (Figure 3b). We could not use the more stringent criteria of selecting only samples with more than 7% MgO because most of the postshield tholeiitic

basalts have 6 to 7% MgO (Table 2). Although it is likely that Hawaiian lavas with 6 to 7% MgO have experienced olivine, plagioclase and clinopyroxene fractionation and many of these lavas have phenocrysts of all three phases (Table 3), it is difficult to properly correct for plagioclase and clinopyroxene fractionation. Rather than make poorly constrained fractionation corrections, we used Sr/Nd ratios to identify samples that may have fractionated plagioclase. Specifically, alkalic basalts LP-1, -7 and -8 have Al<sub>2</sub>O<sub>3</sub>/CaO at the high end (1.33 to 1.37) of the selected range, and they have atypically low Sr/Nd ratios (Figures 11 and 12). Therefore these

samples were not used in constraining source characteristics. Because olivine-melt equilibria are well understood, we corrected for olivine fractionation. Olivine with a calculated equilibrium composition ( $Fe/Mg K_D = 0.3$ ) was added in 1% increments until a composition with 16% MgO was attained. These calculated major oxide compositions (Table 7) are similar for alkalic and tholeiitic lavas (except for  $Na_2O$ ,  $K_2O$  and  $TiO_2$ ).

Minster and Allegre [1978] showed how abundances of incompatible elements in a suite of isotopically homogeneous primary magmas can be inverted to obtain source characteristics. The relevant equation for nonmodal batch melting [Shaw, 1970] is  $C_1/C_0 = 1/(F+D_0-PF)$ , where  $C_1$  is concentration in liquid,  $C_0$  is source concentration,  $F$  is degree of melting,  $D_0$  is initial bulk solid/melt partition coefficient, and  $P = \sum p_i D_i$ , ( $p_i$  are weight fractions of phases entering the melt with mineral/melt partition coefficients,  $D_i$ ). This inversion process in a simplified form was used for Hawaiian posterosional alkalic lavas [Clague and Frey, 1982], alkalic and tholeiitic lavas from Kohala Volcano [Feigenson et al., 1983], and tholeiitic lavas from Kilauea Volcano [Hofmann et al., 1984]. The approach involves plots of  $C^H/C^i$  versus  $C^H$ , where  $C^H$  is the concentration of a highly incompatible element and  $C^i$  is the concentration of a different incompatible element. In choosing the appropriate highly incompatible element for Mauna Kea lavas, we avoided alkali metals because their abundances were affected by postmagmatic mobility and Th because our data are not precise at concentrations less than 1 ppm Th. The elements La and Nb yield good correlations in  $C^H/C^i$  vs  $C^H$  plots (e.g., Figure 14). A very simplistic but reasonable approach is to assume that when  $C^H/C^i$  is nearly constant relative to analytical error over a wide concentration range of  $C^H$  (e.g., Nb/Ba, Nb/Ta, Nb/La, Nb/Ce in Figure 14), the  $C^H/C^i$  ratio in the lavas is characteristic of the source. With this assumption the source of the postshield lavas has chondritic ratios of Nb/Ba and Nb/Ta but slightly greater than chondritic ratios of Nb/La and Nb/Ce (Table 8). For a nonmodal batch melting process the slope of a  $C^H/C^i$  versus  $C^H$  plot is  $D_0/C_0^i$  and the intercept is  $C_0^H/C_0^i (1-P^i)$ ; therefore the nearly uniform Nb/Ba, Nb/Ta, Nb/La and Nb/Ce in the postshield lavas indicates that  $D_0$  and  $P^i$  are  $\sim 0$ . However, for relatively more compatible elements, such as Zr, Hf, Sr, Y and Nd to Lu, estimates of  $C_0^H/C_0^i$  require a knowledge of  $P^i$ . In an inversion of geochemical data for basalts from Kohala Volcano, Feigenson et al. [1983] assumed  $P^i$  was constant

because they concluded that the tholeiitic and alkalic primary magmas were similar in major element composition (see Figure 3 of Feigenson et al. [1983]). However, in detail melt compositions must vary with degree of melting and Albarede [1983, see appendix] showed that the assumption of constant  $P^i$  can lead to erroneous implications even when  $C^H/C^i$  trends are linear.

Albarede [1983] and Albarede and Tamagnan [1988] presented an alternative inversion model. Like the inversion process proposed by Minster and Allegre [1978], the principal assumptions are that the primary lavas formed by batch equilibrium melting of a compositionally homogeneous source. The inputs are trace element abundances in primary melts, the residual mineralogy and mineral/melt partition coefficients for residual phases. No assumptions are required about the phases entering the melt. Moreover, if only incompatible elements are used the approach is not very sensitive to post-melting mineral fractionation. This inversion process finds the source composition which lies at the closest possible distance to the locus of combinations between the melt compositions and their residual phase compositions (sample subspace). The outputs are element concentrations in the source and the proportions of residual phases. However, if a mineral phase does not contain the elements measured (e.g., REE in residual olivine), only relative values of source concentrations and residual phases can be inferred. Albarede [1983] used an early version of this approach to infer that the alkalic lavas forming the Honolulu Volcanics [Clague and Frey, 1982] were derived from a source with relative light REE enrichment, but that the tholeiitic and alkalic basalts erupted at Kohala Volcano [Feigenson et al., 1983] were derived from a source with near-chondritic proportions of REE.

As discussed by Albarede and Tamagnan [1988] the best source composition is found from a cumulated projection matrix as the eigenvector associated with the smallest eigenvalue. The eigenvector-eigenvalue pair is referred to as the N-th eigenvector, where N is the number of elements. The N-th eigenvalue has the significance of a sum of squared residuals, and it is a direct measure of the misfit. Although this solution is statistically the best, it may require unacceptable results such as negative source concentrations, or negative amounts of a residual phase. If the N-th eigenvalue leads to physically unacceptable results, a small "proportion" of the eigenvector associated with the second smallest eigenvalue (N-1) is added until a physically acceptable solution is obtained. Because this addition increases the misfit, it is kept to a minimum and monitored with the angle  $\phi$  between the solution and the N-th eigenvector.

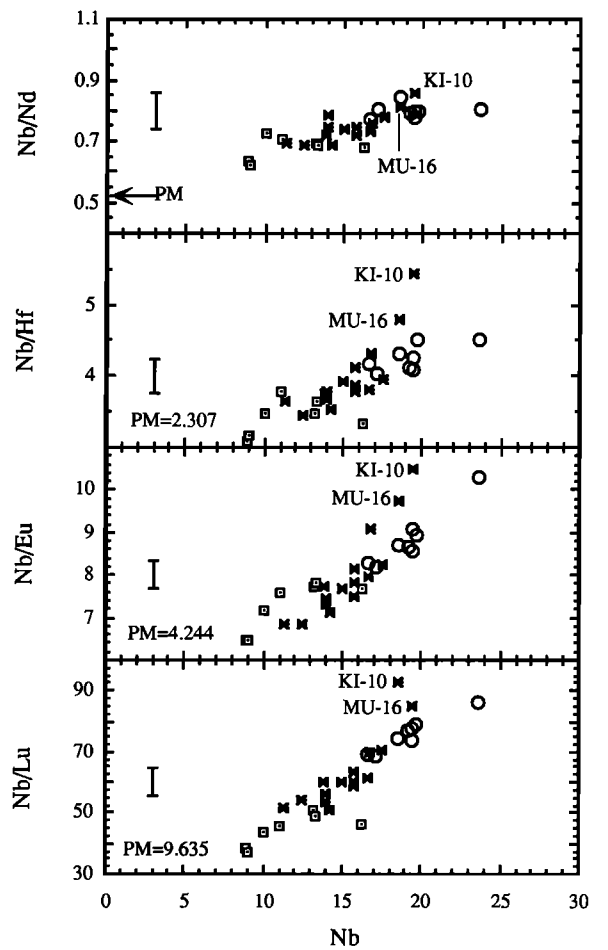
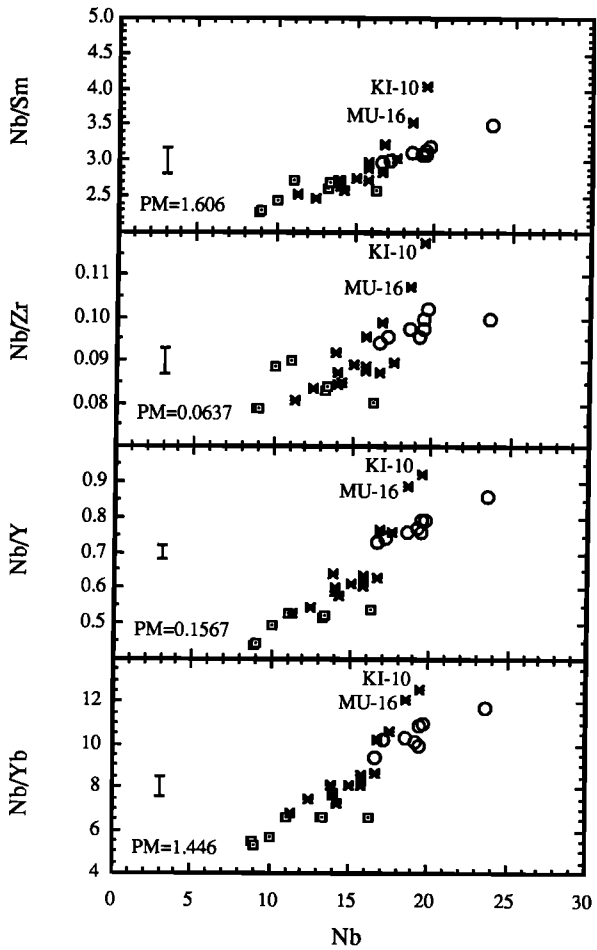
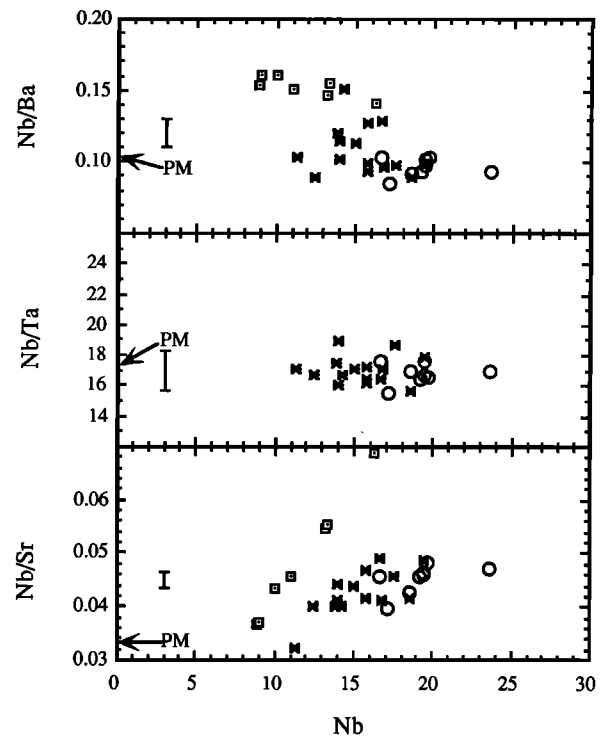
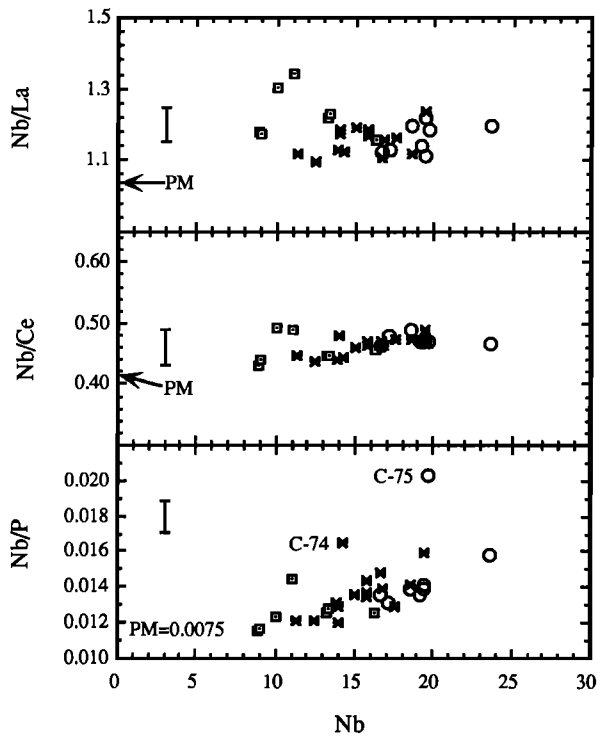
Using this approach several subsets of the Mauna Kea basalt data were used to infer source composition and the proportions of residual clinopyroxene and garnet. The inputs to the inversion program are incompatible element abundances (observed abundances corrected to 16% MgO by incremental addition of equilibrium olivine) and mineral/melt partition coefficients for clinopyroxene, garnet and olivine (Table 9). Data for 7 shield samples and 21 postshield samples (see Table 7 for listing) were used. Samples KI-10 and MU-16 have atypically high Nb/REE and Nb/Zr-Hf ratios (Figure 14). Because these discrepancies may reflect a more complex petrogenesis than partial melting and olivine fractionation, data for these two samples were not used in the inversion. Possible relative errors on the observed data were set at 5%. The relative uncertainty associated with the choice of partition coefficients was set at 50% with an 85% positive error correlation. Calculations were made for the postshield tholeiitic and alkalic basalts as individual groups and as a combined group. Also, the postshield and shield lavas were considered together as a single group.

To first order, the results are similar regardless of the Mauna Kea data set chosen for inversion. For example, the

TABLE 7. Range of Inferred Primary Magma Compositions for the Postshield Stage of Mauna Kea Volcano

	Alkalic Basalts	Tholeiitic Basalts
SiO <sub>2</sub>	45.7±0.6	46.4±0.8
TiO <sub>2</sub>	2.67±0.11	2.41±0.14
Al <sub>2</sub> O <sub>3</sub>	11.2±0.3	11.1±0.4
Fe <sub>2</sub> O <sub>3</sub>	1.7±0.2	1.7±0.2
FeO	11.2±0.2	11.6±0.6
MgO	16.0	16.0
CaO	9.0±0.4	8.7±0.5
Na <sub>2</sub> O	2.0±0.1	1.8±0.1
K <sub>2</sub> O	0.60±0.09	0.28±0.16

Lava compositions adjusted to 16% MgO by adding equilibrium olivine ( $K_D = 0.3$ ) in 1% increments to the 23 lavas in Table 2 with  $Al_2O_3/CaO$  between 1.15 and 1.40. Mean and one standard deviation is for 8 alkalic basalts (KI-1, KI-2, KI-6, KI-7, KI-13, MU-5, LP-9 and C-75) and 15 tholeiitic basalts (KI-9, KI-10, MU-13, MU-15, MU-2, MU-3, MU-4, MU-7, MU-10, MU-16, LP-11, LP-2, LP-3, LP-6, and C-74). Data normalized to 100% with  $FeO/(FeO + 0.9 Fe_2O_3) = 0.85$ .



- Shield Tholeiite
- Postshield Alkalic Basalt
- ✕ Postshield Tholeiitic Basalt



TABLE 8. Incompatible Element Ratios in Mauna Kea Basalts

	Primitive Mantle*	Postshield†	Postshield‡	Shield§
La/Nb	0.96	0.86±0.03	0.91±0.02	0.81±0.05
Ce/Nb	2.49	2.13±0.07	2.91±0.04	2.19±0.11
Ba/Nb	9.80	9.6±1.4	5.2±0.2	6.6±0.3
Nb/Ta	17.4	17.2±1	-	-

\*Values from Table 1 of Sun and McDonough [1989].

†Average (±1 standard deviation) of 26 tholeiitic and alkalic basalts with Al<sub>2</sub>O<sub>3</sub>/CaO ranging from 1.15 to 1.40. Samples are those listed in Table 7 footnote plus alkalic basalts LP-1, -7 and -8.

‡Ratios inferred from inversion of postshield data; see caption to Figure 15.

§Average (±1 standard deviation) of seven shield basalts (Table 1). Ta is not reported because these lavas were ground in WC.

data set for postshield tholeiitic and alkalic basalts yield a statistically good fit ( $\chi^2 = 0.002$ ) for  $\phi = 15^\circ$ . The required source is relatively depleted in the most incompatible elements and a similarly depleted source is required when only the postshield tholeiites are considered or if the shield and postshield lavas are considered as a group (Figure 15). Compared to the postshield lavas the inferred source has similar La/Nb and Ce/Nb but significantly lower Ba/Nb (Table 8). As shown in Figure 16, characteristics of the  $\phi = 15^\circ$  results for the postshield basalts are ~6-8% residual garnet, ~2-5% residual clinopyroxene, and a very low degree of melting (0.15 to 0.32%). By increasing  $\phi$  to 70° larger degrees of melting are required (0.6 to 1.5%), the garnet/clinopyroxene ratio decreases in the residue, and the source is not as depleted in highly incompatible elements; however, the misfit ( $\chi^2 = 0.021$ ) is larger (Figures 15 and 16).

There are several aspects of the inversion results which are puzzling. For example, the inferred degrees of melting (Figure 16) are significantly lower than commonly proposed for tholeiitic magmas (typically 5 to 40%). Because the

TABLE 9. Trace Element Mineral/Melt Partition Coefficients Used in the Inversion Model

	Cpx	Garnet	Olivine
Ba	0.00	0.00	0.00
Nb	0.07	0.00	0.00
La	0.067	0.004	0.00
Ce	0.098	0.01	0.00
Sr	0.18	0.00	0.00
Nd	0.21	0.047	0.00
Zr	0.25	0.14	0.00
Sm	0.26	0.14	0.001
Yb	0.29	4.3	0.027

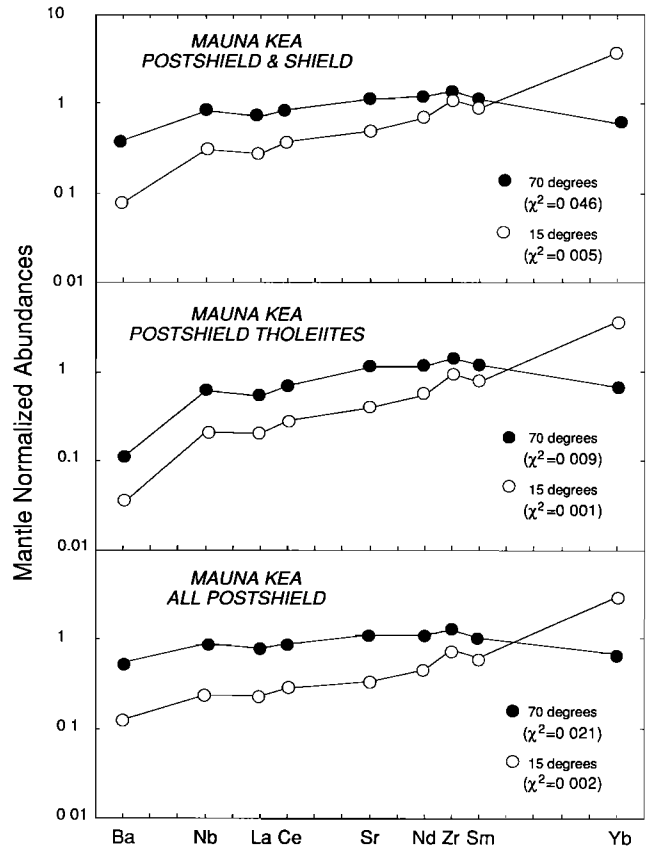


Fig. 15. Mantle normalized incompatible element abundances inferred for mantle sources using the approach of Albaredo and Tamagnan [1988] to invert the data in Figure 14 except that data for samples MU-16, KI-10, LP-1, LP-7 and LP-8 were not used; see text. The inversion results only constrain the relative abundances of incompatible elements in the source. In order to have inferred sources with REE abundances similar to those observed in upper mantle rocks, abundances of intermediate and high atomic number REE were fixed at approximately chondritic abundances. Lower panel shows inversion results for postshield tholeiitic and alkalic basalts; middle panel is based only on inversion of postshield tholeiite data; upper panel is based on inversion of postshield and shield basalt data. The best fit solutions ( $\chi^2 = \lambda_n \cos^2\phi + \lambda_{n-1} \sin^2\phi$ ) in each case (open circles) are for sources which are relatively depleted in the most incompatible elements. Normalized results and estimated errors for  $\phi = 15^\circ$  in the lower panel are Ba = 0.128±0.006; Nb = 0.242±0.003; La = 0.229±0.003; Ce = 0.281±0.002; Sr = 0.338±0.002; Nd = 0.468±0.003; Zr = 0.779±0.004; Sm = 0.625±0.003; Yb = 2.990±0.001; for  $\phi = 70^\circ$  degrees in the lower panel, Ba = 0.549±0.015; Nb = 0.891±0.007; La = 0.788±0.007; Ce = 0.882±0.004; Sr = 1.139±0.004; Nd = 1.131±0.004; Zr = 1.305±0.008; Sm = 1.060±0.007; Yb = 0.694±0.003. Normalizing values for REE are from Boynton [1984], and values for other elements were chosen to be consistent with chondritic ratios (e.g., Table 1 of Sun and McDonough [1989]).

Fig. 14. Nb/X versus Nb (ppm), where X is another incompatible element. If eutectic melting and a batch equilibrium melting model are assumed, the slopes and intercepts in these diagrams can be used to constrain source composition and mineralogy [Minster and Allegre, 1978; Hofmann and Feigenson 1983]. Plotted data are for the seven shield samples (open squares) and postshield Hamakua Volcanics with Al<sub>2</sub>O<sub>3</sub>/CaO between 1.15 and 1.40 with all samples corrected to 16% MgO by addition or subtraction of olivine. The alkalic basalts have >17 ppm Nb whereas the tholeiitic basalts have <17 ppm Nb, except for anomalous samples KI-10 and MU-16. We interpret the high Nb/P ratios in C-74 and C-75 as reflecting inaccurate P<sub>2</sub>O<sub>5</sub> data in Macdonald and Katsura [1964]. Primitive mantle ratios (PM) from Sun and McDonough [1989]. Vertical error bar indicates ±1 standard deviation based on replicate analyses of BHVO-1.

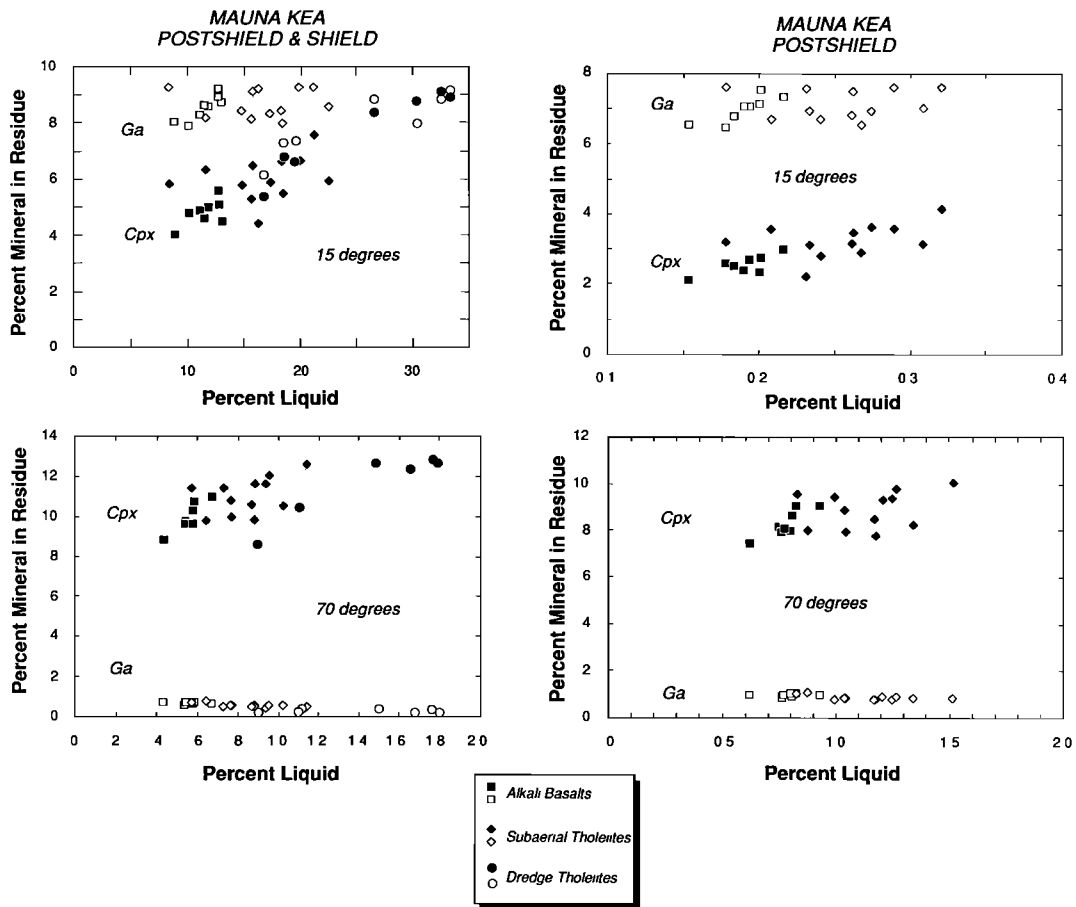


Fig. 16. The proportion of garnet and clinopyroxene in residues (wt. %) as a function of percent melting for various inversion models. Left panels indicate results for shield and postshield data set; right panels indicate results for only postshield basalts. Upper panels indicate results for best fit models ( $\phi = 15^\circ$ ). Both data sets yield residues with ~7 to 10% garnet and lesser amounts of clinopyroxene with the unexpected result of residual clinopyroxene increasing with increasing melting. The lower panels provide statistically poorer fits ( $\phi = 70^\circ$ ), and they yield higher degrees of melting and small amounts of residual garnet.

inversion constrains only the relative incompatible element abundances in the source and relative proportions of residual phases (i.e., olivine/(clinopyroxene + garnet)), the inferred degree of melting (Figure 16) depends upon our choice of source abundances. In Figure 15, we chose sources with near-chondritic abundances of intermediate and high atomic number REE. This choice is consistent with REE abundances in upper mantle rocks [e.g., McDonough and Frey, 1989]. However, sources with higher abundances of incompatible elements would lead to higher degrees of melting. Another problem with the inversion results is the increase in the proportion of residual clinopyroxene with increasing degrees of melting (Figure 16). This trend is contrary to the established experimental result that the abundance of residual clinopyroxene decreases with increasing extents of melting. A likely explanation is that some of the assumptions made for the inversion calculation are invalid. For example, a low percent of batch melting is required in order to create the observed variations in abundance ratios of incompatible elements (Figures 11 and 14). If the melt segregation process is more efficient in fractionating trace element abundances than batch melting [e.g., Ribe, 1988], then higher degrees of melting are plausible. Another possibility is that the sources were isotopically homogeneous but modally heterogeneous; i.e., if the sources contain variable amounts of clinopyroxene, the larger amounts of residual clinopyroxene at higher degrees of

melting (Figure 16) may reflect derivation from a more clinopyroxene-rich source.

##### 5. Evolution Of Mauna Kea Volcano

Pillow lavas dredged from >1.9 km water depth on the east rift are interpreted to be late-shield lavas. Their whole rock and glass compositions overlap with shield lava compositions from Kilauea and Mauna Loa (e.g., Figure 3a and Garcia et al. [1989]). Although these whole rocks are MgO-rich (9.4 to 16.8%), they are not primary melts; however, they contain magnesian olivine phenocrysts (up to Fo<sub>90.5</sub>) which formed in parental melts with ~14 to 17% MgO.

An important result is that unlike lavas from Haleakala [Chen and Frey, 1985] and Kauai [Feigenson, 1984; Clague and Dalrymple, 1988], lavas from Mauna Kea do not define systematic temporal trends in  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  [Kennedy et al., 1991]. This absence of systematic temporal isotopic variations is not understood [Kennedy et al., 1991]; however, more isotopic data for Mauna Kea shield lavas are required to fully assess temporal trends. Mauna Kea lavas do define systematic temporal changes in composition, ranging from tholeiitic basalt shield lavas to relatively young evolved alkalic postshield lavas with an intervening postshield interval of intercalated tholeiitic and alkalic basalt. In addition, with decreasing age the tholeiitic lavas are more enriched in highly

incompatible elements. If a common source composition is assumed for the late shield and postshield tholeiites, the lower incompatible element content of the shield tholeiites and their lower Ba/Nb, Ba/La and La/Ce ratios require that the shield tholeiites formed by significantly higher degrees of melting.

During the basaltic substage of postshield volcanism, tholeiitic and alkalic basalts with very similar isotopic ratios (Sr, Nd and Pb) were erupted. During this growth stage eruption rates were decreasing precipitously [Frey et al., 1990], presumably reflecting a decreasing magma supply as Mauna Kea moved away from the mantle plume. As a result, it is plausible that basalts of similar age formed by significantly different degrees of melting thereby forming stratigraphic sections of intercalated tholeiitic and alkalic basalt. The basaltic substage also includes high Fe-Ti basalts which formed principally by crystal-melt segregation at shallow crustal levels. Eventually, the decrease in eruption and magma production rates led to crystallization of shallow magma chambers, cessation of basaltic eruptions and stagnation of basaltic magma deep within the crust. Cooling and extensive mineral-melt segregation at lower crustal pressures formed the evolved, low density, alkalic lavas which comprise the hawaiitic substage of postshield volcanism on Mauna Kea Volcano [Frey et al., 1990].

#### 6. Implications For The Generation Of Hawaiian Basalts

Important questions are (1) Was the source or mixture of source components enriched in incompatible elements relative to a primitive mantle composition? (2) What was the range in degree of melting? (3) Were residual garnet and clinopyroxene important in controlling melt composition?

From a geochemical perspective the first two questions are inter-related. The principal facts are that Hawaiian shield lavas are enriched in incompatible elements relative to MORB (e.g., Figure 10), and they have Sr and Nd isotopic ratios; consistent with long-term depletion in highly incompatible elements (i.e., Rb/Sr and Nd/Sm less than bulk Earth ratios). Therefore the observed enrichment of incompatible elements resulted from relatively small degrees of melting of a long-term depleted source or larger degrees of melting of a recently enriched source. To date, all inversion models based on the incompatible element contents of Hawaiian shield lavas require a source that is relatively depleted in highly incompatible elements [e.g., Albarede, 1983; Feigenson et al., 1983; Hofmann et al., 1984; Budahn and Schmitt, 1985; this paper]. Of course, this is a rigorous constraint only if the assumptions used in the inversion models are valid; e.g., batch melting of a source with uniform composition, isotopic ratios and modal mineralogy. Although the batch melting model is unrealistically simple, it is not evident that more complex melting models would require an enriched source [e.g., Richter, 1986; Ribe, 1988].

Finally, the third question, particularly, the role of residual garnet is clearly answered by trace element abundances. Specifically, the relative uniformity of the fractionation corrected Y and HREE contents ( $Y = 24.1 \pm 1.6$  ppm and  $Yb = 1.80 \pm 0.11$  ppm in postshield Mauna Kea lavas, which range in Nb from 11.3 to 23.7 ppm) requires a residual phase that preferentially incorporates Y and HREE, i.e.,  $D_{min/melt} > 1$ . The only likely phase is garnet. Also, Hofmann et al. [1984] noted the uniformity of Yb in tholeiitic basalts from Kilauea, and they argued strongly that source and residues of Kilauea basalts contain garnet.

**Acknowledgments.** Several individuals provided important contributions to our geochemical analysis program. We thank Elizabeth Neilsen for helping with sample collection along the Hamakua coast, C.-Y. Chen and Amy Moser for assistance in sample preparation, T. Furman, D. Gerlach, R. Hickey-Vargas, and D. Tormey for their efforts in obtaining XRF

data, R. Multhaup for assistance with microprobe analyses, and P. Ila in obtaining NA data. We thank J.M. Rhodes for access to the University of Massachusetts XRF lab, S. Hart for access to the MIT Mass Spectrometry Lab, and R. Price (LaTrobe University) for  $Fe^{+3}/Fe^{+2}$  determinations, H.-J. Yang for assistance with graphics, and D. Frank for word processing. In addition, this paper benefitted from review comments by D. Clague, M. Feigenson, A. Hofmann, and M. Rutherford and discussions with T. Grove and T. Sisson. Neutron irradiations were done at MIT Nuclear Reactor. This research was supported by NSF grants EAR 8218982, EAR8419723, and EAR875809 to (F.A.F.), and EAR826787, OCE87167042 (M.O.G.) and EAR8018798 (W.S.W.). SOEST Contribution 2500.

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(Received June 12, 1990;  
revised March 29, 1991;  
accepted March 29, 1991.)