# Magmatic history of the East Rift Zone of Kilauea Volcano, Hawaii based on drill core from SOH 1 

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#### Abstract

Deep drilling has allowed for the first time an examination of most of the shield stage of a Hawaiian volcano when it is centered over the hotspot and most of its volume is produced. We determined the lithologies, ages, geochemical characteristics and accumulation rates of rocks from the continuously cored, $\sim 1.7 \mathrm{~km}$ deep Scientific Observation Hole (SOH) 1, which was drilled into Kilauea's East Rift Zone. The uppermost $\sim 750 \mathrm{~m}$ of this hole contain relatively unaltered subaerially quenched lavas; the lower portion is a mixture of weakly to moderately altered hyaloclastites and massive basalts intruded by dikes. Unspiked $\mathrm{K}-\mathrm{Ar}$ dating was attempted on 14 lavas but only four yielded geologically reasonable ages. The oldest age is $351 \pm$ 12 ka for a sample from 1541 m . Eighty XRF major and trace element analyses of lavas from throughout the section demonstrate that these lavas have the same compositional range as historical Kilauea lavas and that there was no systematic geochemical variation for the volcano over the last 350 ka . Kilauea's average yearly magma supply during this period was probably $\sim 0.05 \pm .01 \mathrm{~km}^{3}$, which is identical to its average historical magma supply rate. Thus, Kilauea's overall source and melting conditions have remained remarkably constant for the last 350 ka despite its having drifted $\sim 42 \mathrm{~km}$ over the Hawaiian plume. These results require that the portion of the plume being melted to produce Kilauea magmas be relatively large and well mixed, which is inconsistent with models that place this source on the margin of a radially zoned plume. © 2000 Elsevier Science B.V. All rights reserved.


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## 1. Introduction

Mantle plumes provide an important means of transferring heat and mass from the deep mantle to the crust (e.g. Sleep, 1990). Models for the structure of mantle plumes depend on reconstructing the

[^0]magmatic history of the volcanoes they produce (e.g. DePaolo and Stolper, 1996). The Hawaiian Islands are the classic example of plume-generated volcanoes but their frequent eruptions and rapid subsidence greatly limit attempts to reconstruct the magmatic history of individual volcanoes. Thus, it has been necessary to manufacture a 'composite' Hawaiian volcano using outcrops from many volcanoes (e.g. Macdonald et al., 1983). Such reconstructions are inherently flawed. They have substantial age gaps, especially for the shield stage of volcanism. Also, they assume that the composition of the plume


Fig. 1. Location map for the Scientific Observation Holes along Kilauea's lower east rift zone. The location of SOH 1 is shown by a star; SOH 4 and SOH 2 are shown by $\oplus$. The three holes form a straight line along the crest of the KERZ. Contour interval is 100 m and the gray area is the Pacific Ocean. The inset map shows the location of SOH 1 (star) on the island of Hawaii, the summits (marked by $*$ ) of the island's five shield volcanoes (K: Kilauea; ML: Mauna Loa), and the location of Kilauea's east rift zone (dashed line).
and its melting processes have been relatively constant over at least a few million years. There is ample evidence that even adjacent Hawaiian shield volcanoes have significant differences in their extent of partial melting and source composition (e.g. Kilauea vs. Mauna Loa volcanoes; Kurz et al., 1995; Pietruszka and Garcia, 1999). Thus, it is essential that the long-term history of an individual Hawaiian volcano be examined if we are to realistically evaluate the structure and melting history of mantle plumes. Deep drilling is the only viable method for achieving this goal.

Three Scientific Observation Holes (SOH 1, SOH

2, and SOH 4) were drilled into lower Kilauea's East Rift Zone (KERZ; Fig. 1) between 1989 and 1991 to assess its geothermal resources. The cores from these $1.7-2.0 \mathrm{~km}$ deep holes sample a major segment of Kilauea's history, and were thought to have the potential to penetrate into Kilauea's pre-shield stage and/or the underlying older Mauna Loa basalts. This study examined rock cores from the 1685 m deep SOH 1 because it was continuously cored with a high core recovery ( $\sim 95 \%$ ) and its rocks are relatively unaltered. Although SOH 4 was continuously cored to a greater depth ( 2001 m ), temperatures were 100$150^{\circ} \mathrm{C}$ higher at a given depth and secondary minerals
are more common in its lavas than those from SOH 1 (Bargar et al., 1995). Only the lower portion of SOH 2 was cored, so it was not considered for this time-series analysis.

Kilauea Volcano is one of the most active and intensely studied volcanoes on Earth, yet its prehistoric record is not well documented (Tilling and Dvorak, 1993). Ninety percent of its surface exposures are younger than 1.5 ka (Holcomb, 1987), its thickest exposure is only $\sim 300 \mathrm{~m}$, and no surface lava is older than $\sim 100 \mathrm{ka}$ (caution: this age is based on lava accumulation rate estimates, not on geochronology; Easton, 1987). A geochemical study of these older surface lavas (the Hilina Basalts, $>25 \mathrm{ka}$; Easton, 1987) found that their compositions to be identical to historical Kilauea lavas, which are distinct from those of the adjacent, historically active Mauna Loa Volcano (Chen et al., 1996). The overall age of Kilauea has been estimated using lava accumulation rates at $\sim 150 \mathrm{ka}($ Lipman, 1995) to $200-300 \mathrm{ka}$ (Easton, 1978).

This study presents a summary of the lithology, age, petrography, and geochemistry of the SOH 1 drill core. Fourteen lavas were selected for unspiked $\mathrm{K}-\mathrm{Ar}$ geochronology and 80 rocks were analyzed by XRF for major and trace elements. The new ages demonstrate that the volcano is much older than previous field-based estimates and that it has been in its shield building stage for at least 350 ka . During this period, the compositional range of Kilauea lavas remained unchanged. Likewise, long-term magma supply rates for this period are identical to the average historical rate. These results require long-term stability in melting processes and a large ( $>42 \mathrm{~km}$ across), well-mixed source within the Hawaiian plume for Kilauea magmas. Thus, it is unlikely that the source of Kilauea magmas is located on the margin of a radially zoned plume, as previously suggested.

## 2. Core lithology

The SOH 1 core was subdivided into lithologic units based on a visual examination. Baked, rubbly, and/or glassy contacts and abrupt changes in mineralogy were used to identify unit contacts. The lithologic units include a'a and pahoehoe flows, massive basalts, dikes, hyaloclastites, lithified breccia, loose
rubble, sand, and ash. Flows with rubbly tops and bottoms and sheared vesicles were identified as a'a. Flows with smooth or ropy glassy contacts and abundant round vesicles were identified as pahoehoe. Units with glassy or high-angle glassy contacts, very little to no vesicularity, and less alteration than adjacent rocks were identified as dikes. Fragmental rocks are common in the lower half of the core. Most are hyaloclastites consisting of weakly to moderately indurated, poorly sorted, vesicular, glassy clasts $<1-5 \mathrm{~cm}$ in diameter. Many of the hyaloclastites contain lava fragments up to 1 m long, with the same mineralogy and vesicularity as the smaller clasts. Thicker lava intervals ( $>2 \mathrm{~m}$ ) within the hyaloclastite section are called massive basalts because their flow types could not be identified. Fragmental units with non-vesicular clasts are termed dike breccias. One unit with coral reef and lava flow fragments in a carbonate-bearing matrix was classified as a carbonate breccia.

The $\sim 1.7 \mathrm{~km}$ of core from SOH 1 was subdivided into 368 lithologic units. The most common unit type is lava flows ( $62.2 \%$ ) with $36.4 \%$ a'a, $22.0 \%$ pahoehoe and $3.8 \%$ massive basalt. Dikes comprise $26.7 \%$ of the units, hyaloclastites $7.6 \%$, breccias $2.1 \%$, and subaerial fragmental debris $\sim 1.4 \%$ (Fig. 2). This perspective changes somewhat if the rock type is evaluated relative to its volumetric abundance in the section. Lava flows are less abundant (45.1\%) with a'a dominating other flow types ( $30.7 \%$ vs. $9.8 \%$ for pahoehoe, and $4.6 \%$ massive basalt) and dikes are more abundant ( $36.4 \%$ ). Hyaloclastite represents $13.2 \%$ of the section and breccia (and other debris) account for $5.3 \%$. The upper 740 m of the hole are almost entirely subaerially quenched lava flows and the section below 1200 m is mostly dikes (Fig. 2). The thickness of a'a flows ranges from $<1-12 \mathrm{~m}$; pahoehoe flows are $<1-5 \mathrm{~m}$ thick. There is no systematic change in flow thickness within the subaerial flow section. The dominance of a'a flows within the core is surprising because only $16 \%$ of Kilauea's surface is covered by a'a (Holcomb, 1987) and the hole is located along the axis of the east rift zone (Fig. 1) where pahoehoe lava is more common. However, a'a flows are more prevalent along the lower portion of the east rift zone than elsewhere on Kilauea (Holcomb, 1987), which is thought to be related to the more differentiated composition (and cooler


temperature during eruption) of these lavas ( < 6.8 wt. \% MgO; Moore, 1983).

A carbonate breccia at $\sim 740 \mathrm{~m}$ depth marks the boundary between exclusively subaerial units to interbedded subaerial and submarine units below. The submarine section consists of interfingering hyaloclastites and massive basalts (including pillow basalt fragments). No coherent pillow lava flow units were identified. Dikes are sparse above the submarine-subaerial transition and increase in abundance down-section; they dominate below $\sim 1200 \mathrm{~m}$. The samples in this study are numbered according to their depth of recovery in meters within the hole.

## 3. Geochronology

Dating young ( $<1 \mathrm{Ma}$ ), low potassium ( $<1 \mathrm{wt} . \%$ ) rocks is inherently difficult using either conventional $\mathrm{K}-\mathrm{Ar}$ or $\mathrm{Ar}-\mathrm{Ar}$ techniques. The unspiked $\mathrm{K}-\mathrm{Ar}$ technique (also known as the Cassignol method) was employed for this study because it has proven effective in dating very young ( $<100 \mathrm{ka}$ ) Hawaiian volcanic rocks (Guillou et al., 1997a,b). This method precisely measures minor variations in the ${ }^{40} \mathrm{Ar} /{ }^{36} \mathrm{Ar}$ ratio between a standard and unknown. In order to optimize success in dating SOH 1 lavas, strict guidelines were used to select only unaltered lava flows. These flows have low vesicularity ( $<10 \%$ ), holocrystalline groundmass, and $\mathrm{K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}$ between 1.3 and 2.0 (unaltered Kilauea tholeiitic basalts have ratios of 1.5-2.0; Wright, 1971). Sample preparation procedures are the same as those described by Guillou et al. (1997a). K and Ar measurements were performed on groundmass aliquots. K was analyzed by atomic and flame emission spectrophotometry. Ar was released from the sample by radio frequency induction heating and measured using mass spectrometry. For more details on the unspiked $\mathrm{K}-\mathrm{Ar}$ technique, see

Fig. 2. Stratigraphic section showing rock type variation for the SOH 1 core. Rock type determinations are based on methods described in the text. The subaerial-submarine contact is $\sim 740 \mathrm{~m}$ depth; the zone of greatest dike intensity is from 1200 to 1650 m in the hole. The asterisks show the location of the samples that yielded positive radiometric ages. The age of the uppermost flow in this section has been inferred based on geologic mapping and $\mathrm{C}^{14}$ dating (Holcomb, 1987).

Table 1
$\mathrm{K}-\mathrm{Ar}$ ages of samples from SOH 1. Age calculations are based on the decay and abundance constants from Steiger and Jäger (1977)

| Depth (m) | K (wt.\%) | Weight molten (g) | ${ }^{40} \mathrm{Ar}(\%)$ | ${ }^{40} \mathrm{Ar}\left(10^{-13} \mathrm{~mol} / \mathrm{g}\right)$ | Age ( $\pm 2 \sigma$ ( ka ) | Age mean value |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Samples giving positive ages |  |  |  |  |  |  |
| 1550.5 | $0.168 \pm 0.002$ | 1.63820 | 0.209 | 1.0439 | $358 \pm 11$ |  |
|  |  | 1.34709 | 0.221 | 1.0007 | $343 \pm 22$ | $351 \pm 12$ |
| 940.5 | $0.279 \pm 0.003$ | 1.55585 | 0.435 | 0.9306 | $192 \pm 18$ |  |
|  |  | 1.52914 | 0.394 | 0.7569 | $156 \pm 24$ | $174 \pm 15$ |
| 755.9 | $0.493 \pm 0.005$ | 1.52053 | 0.137 | 0.2486 | $29 \pm 9$ |  |
|  |  | 1.96710 | 0.103 | 0.2262 | $26 \pm 9$ | $28 \pm 6$ |
| 409.0 | $0.394 \pm 0.004$ | 2.03647 | 0.022 | 0.034 | $5 \pm 5$ | $5 \pm 5$ |
|  |  | 2.03647 | -0.008 | <0 |  |  |
| Samples giving zero ages |  |  |  |  |  |  |
| 234.4 | $0.427 \pm 0.010$ | 0.90382 | -0.158 | $<0$ |  |  |
| 537.4 | $0.547 \pm 0.005$ | 1.58540 | -0.034 | $<0$ |  |  |
| 691.9 | $0.502 \pm 0.009$ | 1.56385 | 0.010 | 0.0249 | $3 \pm 30$ |  |

Cassignol and Gillot (1982) and Gillot and Cornette (1986).

Fourteen samples from throughout SOH 1 were selected for unspiked K-Ar geochronology. Eleven of the fourteen samples are from above the submar-ine-subaerial transition. This sampling bias is due to the scarcity of suitable lavas below the submarinesubaerial transition that meet the selection criteria. Despite the stringent sample criteria, only four of the samples gave geologically meaningful ages (Table 1). No radiogenic argon was measured in 10 of the samples (Table 1). The lack of radiogenic argon in sample 624.8 could be explained by its location between two dikes, which may have caused heating of the rock and ${ }^{40} \mathrm{Ar}$ loss. The lack of radiogenic argon in the other samples demonstrates the inherent difficulties in dating young rocks with low amounts of K ( $<0.5 \mathrm{wt} . \%$ ) and ${ }^{40} \mathrm{Ar}(\ll 1 \%)$. Three of the ages were successfully replicated and these samples cover much of the SOH 1 hole (Table 1). The date for shallowest samples was not duplicated and must be considered only an approximate age. The age of the surface flow at the SOH 1 location is estimated to be $250-500$ years old based on geologic mapping and $\mathrm{C}^{14}$ ages (Holcomb, 1987).

## 4. Petrography

In hand specimen, most SOH 1 rocks appear to be aphyric to weakly olivine and plagioclase phyric.

Eighty thin sections were examined to better characterize the petrography and degree of alteration of the SOH 1 core. Fifty-one sections are from subaerial ( $65 \%$ a'a and $35 \%$ pahoehoe) and submarine lava

Table 2
Vesicle-free modal mineralogy of phenocrysts ( $>0.5 \mathrm{~mm}$ ), vesicle percentages and whole-rock MgO content for selected SOH 1 lavas. Modes based on 400 points counted/sample and are in vol\%. Cpxclinopyroxene; Plag-plagioclase

| Depth (m) | Olivine | Cpx | Plag | Matrix | Vesicles | MgO (wt\%) |
| :--- | :--- | :--- | :--- | :--- | ---: | ---: |
| 14.0 | 0.1 | 2.0 | 1.0 | 96.9 | 20 | 6.6 |
| 57.3 | 3.0 | 3.0 | 0.1 | 93.9 | 6 | 10.1 |
| 165.5 | 1.0 | 1.0 | 0.1 | 97.9 | 10 | 6.1 |
| 234.4 | 1.5 | 3.0 | 0.5 | 95.0 | 3 | 10.8 |
| 278.0 | 0.1 | 0.1 | 1.5 | 98.3 | 22 | 5.5 |
| 367.0 | 2.0 | 0.1 | 0.1 | 97.8 | 17 | 8.7 |
| 400.8 | 0.1 | 0.1 | 1.5 | 98.3 | 31 | 6.6 |
| 409.0 | 0.1 | 0.1 | 0.5 | 99.3 | 13 | 7.0 |
| 425.2 | 0.5 | 0.1 | 0.0 | 99.4 | 9 | 7.9 |
| 466.3 | 3.5 | 2.5 | 1.0 | 94.0 | 19 | 7.0 |
| 514.5 | 0.1 | 2.0 | 0.1 | 97.8 | 2 | 6.3 |
| 607.5 | 1.0 | 0.0 | 0.0 | 99.0 | 9 | 9.7 |
| 691.9 | 4.0 | 0.0 | 0.0 | 96.0 | 9 | 10.2 |
| 719.9 | 4.5 | 0.1 | 0.0 | 95.5 | 23 | 15.4 |
| 755.9 | 0.1 | 0.5 | 0.0 | 99.4 | 8 | 6.7 |
| 784.2 | 8.0 | 0.1 | 0.0 | 91.9 | 6 | 12.3 |
| 846.7 | 2.0 | 0.1 | 0.0 | 97.9 | 33 | 8.0 |
| 940.5 | 0.5 | 0.5 | 1.5 | 97.5 | 3 | 7.1 |
| 1036.3 | 0.1 | 0.1 | 0.0 | 99.8 | $<1$ | 7.6 |
| 1247.5 | 1.0 | 0.1 | 0.5 | 98.4 | $<1$ | 6.9 |
| 1466.7 | 0.0 | 0.5 | 0.1 | 99.4 | $<1$ | 6.0 |
| 1550.5 | 0.1 | 0.1 | 0.0 | 99.8 | 12 | 7.5 |
| 1672.3 | 0.0 | 0.0 | 0.1 | 99.9 | 6 | 7.5 |

Table 3
XRF major element analyses of the freshest SOH 1 lavas. $\mathrm{Fe}_{2} \mathrm{O}_{3}^{*}$ is total iron. All values in wt. $\%$. A.I. $=$ Alteration Index: $1 —$ unaltered; $2 —$ thin iddingsite rims on olivine; 3 -extensive iddingsite alteration of olivine, 4 -matrix altered and vesicles have secondary mineral linings; 5 high temperature minerals present (e.g. talc). LOI-loss on ignition. Analysts: T. Hulsebosch

| Depth (m) | $\mathrm{SiO}_{2}$ | $\mathrm{TiO}_{2}$ | $\mathrm{Al}_{2} \mathrm{O}_{3}$ | $\mathrm{Fe}_{2} \mathrm{O}_{3}^{\text {* }}$ | MnO | MgO | CaO | $\mathrm{Na}_{2} \mathrm{O}$ | $\mathrm{K}_{2} \mathrm{O}$ | $\mathrm{P}_{2} \mathrm{O}_{5}$ | LOI | SUM | A.I. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 14.0 | 50.52 | 2.92 | 13.51 | 12.98 | 0.17 | 6.64 | 10.09 | 2.46 | 0.55 | 0.32 | 0.00 | 100.15 | 1.5 |
| 36.3 | 49.47 | 2.30 | 12.69 | 12.91 | 0.17 | 9.78 | 10.36 | 1.93 | 0.45 | 0.26 | 0.00 | 100.32 | 1.5 |
| 57.3 | 49.21 | 2.35 | 12.50 | 12.90 | 0.16 | 10.07 | 10.62 | 1.93 | 0.43 | 0.24 | 0.00 | 100.41 | 1 |
| 67.1 | 50.52 | 3.51 | 13.37 | 13.94 | 0.18 | 5.52 | 9.41 | 2.68 | 0.69 | 0.38 | 0.00 | 100.16 | 1 |
| 109.1 | 51.00 | 3.54 | 13.64 | 13.39 | 0.17 | 5.32 | 9.23 | 2.77 | 0.73 | 0.43 | 0.15 | 100.21 | 1 |
| 125.0 | 49.52 | 2.50 | 13.23 | 11.95 | 0.16 | 7.98 | 11.67 | 2.06 | 0.53 | 0.27 | 0.00 | 99.86 | 1.5 |
| $147.5^{\text {a }}$ | 50.24 | 3.01 | 13.83 | 13.04 | 0.17 | 6.35 | 10.31 | 2.45 | 0.57 | 0.33 | 0.00 | 100.28 | 1.5 |
| 165.5 | 50.27 | 2.85 | 13.85 | 13.39 | 0.17 | 6.12 | 10.28 | 2.25 | 0.56 | 0.30 | 0.00 | 100.04 | 1 |
| 176.8 | 50.64 | 2.83 | 13.72 | 12.78 | 0.17 | 6.10 | 10.02 | 2.41 | 0.61 | 0.34 | 0.00 | 99.61 | 1 |
| 192.0 | 50.24 | 2.97 | 13.22 | 13.40 | 0.17 | 7.06 | 9.82 | 2.33 | 0.54 | 0.31 | 0.19 | 100.04 | 1 |
| 211.5 | 50.21 | 2.46 | 13.54 | 12.72 | 0.17 | 7.28 | 10.74 | 2.01 | 0.40 | 0.23 | 0.00 | 99.77 | 2 |
| $234.4{ }^{\text {a }}$ | 49.58 | 2.04 | 12.62 | 11.87 | 0.16 | 10.80 | 10.58 | 1.91 | 0.36 | 0.20 | 0.00 | 100.11 | 1 |
| 263.6 | 50.81 | 2.64 | 13.95 | 12.71 | 0.17 | 6.45 | 10.41 | 2.22 | 0.52 | 0.27 | 0.00 | 100.14 | 1 |
| 278.0 | 50.49 | 3.41 | 13.32 | 14.01 | 0.18 | 5.50 | 9.52 | 2.59 | 0.64 | 0.37 | 0.00 | 100.02 | 1.5 |
| 300.2 | 50.84 | 3.17 | 13.49 | 13.82 | 0.17 | 5.52 | 9.52 | 2.49 | 0.69 | 0.37 | 0.00 | 100.08 | 1 |
| 319.4 | 50.32 | 2.57 | 13.52 | 12.66 | 0.17 | 7.04 | 10.81 | 2.16 | 0.41 | 0.26 | 0.00 | 99.89 | 1.5 |
| 334.1 | 50.09 | 2.22 | 13.29 | 12.52 | 0.16 | 8.14 | 10.41 | 1.98 | 0.37 | 0.21 | 0.09 | 99.40 | 1.5 |
| $353.0{ }^{\text {a }}$ | 50.39 | 2.41 | 13.52 | 12.67 | 0.17 | 7.14 | 10.82 | 1.98 | 0.42 | 0.24 | 0.09 | 99.73 | 1 |
| 367.0 | 50.14 | 2.59 | 12.82 | 12.48 | 0.17 | 8.69 | 10.04 | 2.27 | 0.45 | 0.27 | 0.97 | 99.91 | 1.5 |
| 383.4 | 50.80 | 2.42 | 13.78 | 12.55 | 0.16 | 6.94 | 10.88 | 2.07 | 0.41 | 0.24 | 0.00 | 100.25 | 1.5 |
| $400.8{ }^{\text {a }}$ | 50.60 | 2.55 | 13.84 | 12.48 | 0.17 | 6.62 | 10.64 | 2.10 | 0.42 | 0.24 | 0.00 | 99.65 | 1.5 |
| $409.0^{\text {a }}$ | 50.51 | 2.47 | 13.62 | 12.58 | 0.17 | 6.98 | 10.71 | 2.29 | 0.37 | 0.25 | 0.33 | 99.93 | 1.5 |
| 425.2 | 50.56 | 2.36 | 13.29 | 12.46 | 0.16 | 7.91 | 10.63 | 1.99 | 0.39 | 0.25 | 0.00 | 100.00 | 1.5 |
| 439.8 | 50.43 | 2.74 | 13.73 | 12.92 | 0.16 | 6.88 | 10.45 | 2.17 | 0.45 | 0.28 | 0.27 | 100.20 | 1 |
| 452.3 | 50.13 | 2.41 | 13.00 | 12.33 | 0.16 | 8.91 | 10.27 | 2.21 | 0.39 | 0.25 | 0.78 | 100.03 | 1 |
| 466.3 | 50.90 | 2.37 | 13.77 | 12.41 | 0.17 | 6.96 | 10.74 | 2.03 | 0.40 | 0.24 | 0.00 | 99.98 | 1 |
| $485.5^{\text {a }}$ | 50.86 | 2.71 | 14.12 | 12.60 | 0.16 | 6.50 | 10.68 | 2.26 | 0.53 | 0.29 | 0.13 | 100.70 | 1 |
| $499.3{ }^{\text {a }}$ | 51.14 | 2.75 | 13.79 | 12.13 | 0.16 | 6.78 | 10.36 | 2.41 | 0.53 | 0.29 | 0.68 | 100.33 | 1 |
| 514.5 | 50.52 | 2.83 | 13.73 | 13.15 | 0.17 | 6.33 | 10.33 | 2.23 | 0.48 | 0.27 | 0.34 | 100.02 | 1 |
| 517.6 | 49.88 | 2.20 | 13.05 | 11.84 | 0.16 | 9.74 | 11.12 | 1.64 | 0.31 | 0.20 | 0.20 | 100.34 | 1 |
| $537.4^{\text {a }}$ | 50.88 | 2.46 | 13.87 | 11.91 | 0.16 | 7.44 | 10.46 | 2.23 | 0.48 | 0.29 | 0.00 | 100.16 | 1 |
| 562.8 | 50.65 | 2.31 | 13.49 | 12.32 | 0.16 | 7.28 | 11.17 | 2.02 | 0.42 | 0.24 | 0.27 | 100.04 | 1.5 |
| 589.8 | 50.50 | 2.19 | 13.45 | 12.57 | 0.16 | 7.86 | 11.01 | 1.99 | 0.40 | 0.24 | 0.03 | 100.36 | 1 |
| $624.8{ }^{\text {a }}$ | 51.14 | 2.55 | 14.00 | 12.25 | 0.16 | 6.68 | 10.49 | 2.00 | 0.45 | 0.25 | 0.25 | 100.24 | 1 |
| $653.8{ }^{\text {b }}$ | 50.75 | 2.84 | 13.78 | 12.86 | 0.16 | 6.69 | 10.29 | 2.19 | 0.57 | 0.31 | 0.00 | 100.41 | 1 |
| 663.9 | 49.69 | 2.50 | 12.81 | 12.28 | 0.17 | 9.34 | 10.40 | 2.16 | 0.47 | 0.25 | 0.02 | 100.04 | 1.5 |
| $670.9{ }^{\text {a }}$ | 50.16 | 2.49 | 12.72 | 12.36 | 0.17 | 9.69 | 10.29 | 1.75 | 0.43 | 0.24 | 0.00 | 100.27 | 1.5 |
| 684.3 | 50.25 | 2.57 | 13.61 | 12.39 | 0.17 | 7.27 | 10.88 | 2.20 | 0.41 | 0.26 | 0.00 | 99.98 | 2 |
| $691.9^{\text {a }}$ | 49.77 | 2.69 | 12.51 | 12.82 | 0.17 | 10.23 | 9.61 | 1.85 | 0.50 | 0.27 | 0.00 | 100.39 | 1.5 |
| 707.7 | 50.23 | 2.72 | 12.66 | 12.20 | 0.16 | 8.80 | 10.92 | 1.85 | 0.54 | 0.28 | 0.00 | 100.35 | 2 |
| 717.5 | 50.66 | 2.83 | 12.97 | 12.86 | 0.18 | 7.30 | 10.09 | 2.38 | 0.45 | 0.29 | 0.43 | 100.00 | 1.5 |
| 719.9 | 48.52 | 1.95 | 10.64 | 13.33 | 0.17 | 15.40 | 8.41 | 1.55 | 0.26 | 0.20 | 2.22 | 100.41 | 1.5 |
| 722.4 | 50.77 | 2.44 | 13.44 | 12.31 | 0.17 | 7.48 | 10.48 | 2.26 | 0.44 | 0.24 | 0.16 | 100.01 | 1.5 |
| 740.4 | 51.05 | 2.36 | 13.50 | 11.99 | 0.16 | 7.28 | 10.65 | 2.26 | 0.39 | 0.23 | 0.00 | 99.83 | - |
| $755.9{ }^{\text {a }}$ | 50.97 | 2.57 | 13.97 | 12.53 | 0.17 | 6.72 | 10.67 | 2.02 | 0.44 | 0.26 | 0.00 | 100.29 | 2 |
| 805.3 | 49.99 | 2.21 | 13.08 | 11.87 | 0.16 | 9.76 | 11.14 | 1.64 | 0.31 | 0.20 | 0.20 | 100.34 | 1 |
| 806.5 | 50.70 | 2.25 | 13.41 | 12.25 | 0.16 | 8.25 | 10.48 | 1.87 | 0.39 | 0.18 | 0.00 | 99.94 | 1 |
| 825.4 | 50.24 | 2.22 | 13.11 | 12.19 | 0.17 | 8.92 | 9.93 | 1.74 | 0.39 | 0.19 | 0.99 | 100.08 | 1.5 |
| $846.7{ }^{\text {b }}$ | 50.11 | 2.40 | 13.55 | 12.11 | 0.17 | 7.96 | 10.94 | 2.27 | 0.35 | 0.26 | 0.92 | 100.10 | 2 |
| 881.2 | 49.27 | 2.27 | 12.93 | 12.06 | 0.17 | 9.10 | 10.25 | 2.11 | 0.40 | 0.23 | 1.10 | 99.88 | - |
| 910.3 | 49.84 | 2.23 | 13.05 | 12.30 | 0.17 | 8.91 | 10.44 | 2.08 | 0.35 | 0.23 | 0.34 | 99.94 | - |

Table 3 (continued)

| Depth (m) | $\mathrm{SiO}_{2}$ | $\mathrm{TiO}_{2}$ | $\mathrm{Al}_{2} \mathrm{O}_{3}$ | $\mathrm{Fe}_{2} \mathrm{O}_{3}^{*}$ | MnO | MgO | CaO | $\mathrm{Na}_{2} \mathrm{O}$ | $\mathrm{K}_{2} \mathrm{O}$ | $\mathrm{P}_{2} \mathrm{O}_{5}$ | LOI | SUM | A.I. |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| $940.5^{\mathrm{a}, \mathrm{c}}$ | 50.51 | 2.51 | 13.79 | 12.50 | 0.16 | 7.14 | 10.63 | 2.08 | 0.34 | 0.23 | - | 99.89 | 2 |
| $983.0^{\text {c }}$ | 50.39 | 2.09 | 12.77 | 11.99 | 0.16 | 9.69 | 11.52 | 1.77 | 0.34 | 0.19 | - | 99.11 | 3 |
| $991.8^{\mathrm{b}}$ | 49.00 | 2.48 | 13.50 | 12.44 | 0.15 | 6.63 | 9.71 | 1.92 | 0.41 | 0.24 | 1.83 | 98.32 | 2.5 |
| $1036.3^{\text {b }}$ | 50.33 | 2.48 | 13.27 | 12.51 | 0.17 | 7.60 | 10.57 | 2.17 | 0.34 | 0.26 | 0.44 | 100.14 | 2.5 |
| $1098.3^{\text {b }}$ | 49.00 | 2.45 | 12.32 | 12.25 | 0.17 | 9.78 | 10.16 | 2.03 | 0.33 | 0.25 | 1.17 | 99.90 | 3 |
| $1217.7^{\text {b }}$ | 49.97 | 2.85 | 13.31 | 12.67 | 0.17 | 7.04 | 10.17 | 2.33 | 0.40 | 0.29 | 0.79 | 100.00 | 3 |
| $1247.5^{\text {b }}$ | 50.09 | 2.68 | 13.40 | 12.65 | 0.16 | 6.90 | 10.21 | 2.33 | 0.40 | 0.29 | 1.17 | 100.27 | 3 |
| $1291.1^{\text {b }}$ | 49.83 | 2.63 | 13.35 | 12.59 | 0.16 | 6.78 | 10.13 | 2.32 | 0.39 | 0.29 | 1.43 | 99.89 | 2.5 |
| $1466.7^{\text {b }}$ | 49.55 | 2.89 | 13.34 | 12.71 | 0.16 | 5.96 | 9.70 | 2.51 | 0.62 | 0.33 | 2.03 | 99.78 | 2.5 |
| $1550.5^{\text {a }}$ | 48.07 | 2.63 | 13.12 | 11.78 | 0.15 | 7.52 | 10.72 | 2.21 | 0.36 | 0.28 | 3.21 | 100.06 | 3.5 |
| $1672.3^{\text {c }}$ | 50.29 | 2.54 | 13.32 | 12.90 | 0.17 | 7.53 | 10.25 | 2.54 | 0.46 | 0.36 | - | 100.34 | 4.5 |

${ }^{\text {a }}$ Sample for geochronology.
${ }^{\mathrm{b}}$ Dike sample.
${ }^{c}$ S. Quane.
flows; 16 are from dikes. Most of the lava flows and all of the dikes are weakly to moderately olivine phyric with no, or only minor amounts of clinopyroxene and plagioclase phenocrysts (Table 2). Plagioclase and clinopyroxene phenocrysts commonly show resorption and/or complex zoning, which are indicative of magma mixing in Hawaiian tholeiitic lavas (e.g. Garcia, 1996). Spinel and glass inclusions are common in olivine from SOH 1 lavas. The groundmass of these lavas consists of plagioclase and clinopyroxene with lesser amounts of olivine, ilmenite, magnetite, and glass. The vesicularity of the flow units ranges widely ( $1-33 \mathrm{vol} . \%$, averaging $\sim 12$ vol. \%; Table 2). The vesicularity of dikes range from 0 to 15 vol. \% but most have $<1$ vol. $\%$.

Twelve hyaloclastite samples and one carbonate breccia were also petrographically examined. The hyaloclastites are aphyric to sparsely olivine, plagioclase, and clinopyroxene phyric. They contain varying proportions of angular to sub-rounded glassy shards ranging in size from $<1$ to 10 mm in diameter. Most shards are thinly coated with alteration minerals (clay, calcite, and aragonite) and some of the hyaloclastites have calcite partially filling matrix voids.

There is no systematic variation in mineralogy with depth in the core except an increase in alteration minerals. The intensity of rock alteration was classified based on thin section analysis. An alteration index, ranging from 1 to 5 , was assigned to each of the samples (see Table 3 for explanation of values). The subaerial rocks show little to no alteration except clay
along fractures; the alteration index for these rocks ranges from 1 to 2.5 (except the more altered samples at 692,720 , and 722 m ). The intensity of alteration increases markedly below the submarine-subaerial boundary with the most altered samples at the base of the hole (Fig. 3). The alteration mineralogy of the submarine rocks grades with depth from smectite, pyrite, and aragonite near the subaerial-submarine boundary to cristobalite, chalcedony, calcite, zeolite, and talc at the bottom of the hole (Bargar et al., 1995). For more information on the hydrothermal mineralogy of SOH 1 rocks, see Bargar et al. (1995).

## 5. Whole rock geochemistry

The same eighty samples examined in thin section were analyzed for major and trace elements by X-ray fluorescence spectrometry (XRF) at the University of Hawaii using a Siemens 303AS XRF spectrometer (Table 3). Samples were ground in a tungsten carbide grinding mill for less than 2 min to minimize contamination by the grinding vessel. Loss-on-ignition (LOI) was determined by weight loss after 8 h of heating at $900^{\circ} \mathrm{C}$. The XRF sample procedures of Norrish and Hutton (1969) and Chappel (1991) were used. For major elements, two fused buttons were made for each sample using a lithium tetraborate flux. A powder pellet was pressed for trace element analysis, which was analyzed in duplicate.

Alteration is always a concern when dealing with

Table 4
XRF determined trace element abundances for the freshest SOH 1 samples (analyst; T. Hulsebosch)

| Depth (m) | Nb | Zr | Y | Sr | Ni | Cr | V | Co | Zn | Cu |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 14.0 | 17.6 | 199 | 32.8 | 343 | 100 | 213 | 338 | 71 | 120 | 130 |
| 36.3 | 13.7 | 167 | 27.8 | 346 | 227 | 490 | 291 | 59 | 105 | 102 |
| 57.3 | 14.0 | 156 | 27.2 | 331 | 256 | 524 | 283 | 58 | 105 | 103 |
| 67.1 | 21.2 | 240 | 38.4 | 377 | 70 | 51 | 356 | 59 | 124 | 132 |
| 109.1 | 24.4 | 275 | 41.8 | 373 | 72 | 51 | 385 | 59 | 153 | 126 |
| 125.0 | 18.7 | 174 | 27.2 | 398 | 150 | 392 | 316 | 50 | 101 | 100 |
| 147.5 | 19.3 | 211 | 33.4 | 390 | 81 | 118 | 318 | 52 | 117 | 126 |
| 165.5 | 17.2 | 194 | 33.1 | 377 | 67 | 76 | 318 | 45 | 113 | 128 |
| 176.8 | 19.0 | 220 | 37.2 | 366 | 83 | 140 | 326 | 45 | 121 | 117 |
| 192.0 | 17.7 | 196 | 31.9 | 340 | 151 | 160 | 364 | 67 | 124 | 99 |
| 211.5 | 11.9 | 152 | 28.7 | 321 | 110 | 278 | 304 | 51 | 109 | 103 |
| 234.4 | 11.7 | 125 | 22.5 | 293 | 300 | 624 | 269 | 77 | 100 | 123 |
| 263.6 | 15.7 | 186 | 32.0 | 343 | 84 | 158 | 336 | 47 | 113 | 99 |
| 278.0 | 20.7 | 236 | 38.1 | 359 | 63 | 38 | 396 | 68 | 136 | 93 |
| 300.2 | 20.8 | 242 | 39.1 | 384 | 64 | 34 | 354 | 43 | 122 | 117 |
| 319.4 | 14.1 | 163 | 28.5 | 322 | 86 | 245 | 314 | 64 | 110 | 105 |
| 334.1 | 13.4 | 142 | 26.4 | 293 | 143 | 383 | 297 | 62 | 103 | 140 |
| 353.0 | 14.9 | 155 | 28.2 | 315 | 80 | 263 | 332 | 61 | 107 | 99 |
| 367.0 | 14.1 | 165 | 27.8 | 315 | 210 | 502 | 298 | 62 | 108 | 112 |
| 383.4 | 14.7 | 157 | 27.7 | 335 | 82 | 201 | 280 | 50 | 101 | 115 |
| 400.8 | 15.1 | 160 | 28.2 | 316 | 87 | 170 | 333 | 68 | 169 | 63 |
| 409.0 | 13.0 | 155 | 27.2 | 320 | 71 | 221 | 289 | 57 | 103 | 108 |
| 425.2 | 15.2 | 160 | 27.0 | 325 | 135 | 347 | 315 | 60 | 109 | 111 |
| 439.8 | 18.5 | 183 | 30.2 | 350 | 112 | 195 | 332 | 55 | 111 | 104 |
| 452.3 | 14.0 | 154 | 26.4 | 307 | 212 | 534 | 302 | 61 | 105 | 86 |
| 466.3 | 14.9 | 155 | 28.0 | 313 | 88 | 212 | 326 | 60 | 108 | 95 |
| 485.5 | 18.6 | 186 | 29.8 | 368 | 88 | 130 | 323 | 54 | 111 | 103 |
| 499.3 | 15.9 | 182 | 30.1 | 337 | 109 | 207 | 318 | 50 | 112 | 98 |
| 514.5 | 16.7 | 174 | 30.4 | 336 | 75 | 99 | 310 | 50 | 110 | 128 |
| 517.6 | 13.0 | 131 | 23.1 | 306 | 212 | 585 | 273 | 60 | 95 | 107 |
| 537.4 | 17.8 | 186 | 31.0 | 338 | 141 | 289 | 314 | 52 | 102 | 73 |
| 562.8 | 15.9 | 160 | 27.3 | 329 | 104 | 284 | 309 | 53 | 107 | 137 |
| 589.8 | 16.1 | 156 | 26.3 | 341 | 131 | 361 | 284 | 58 | 102 | 97 |
| 624.8 | 15.2 | 180 | 31.1 | 347 | 98 | 183 | 320 | 58 | 119 | 106 |
| $653.8{ }^{\text {a }}$ | 17.9 | 195 | 30.7 | 365 | 116 | 193 | 328 | 55 | 114 | 128 |
| 663.9 | 14.2 | 150 | 25.4 | 305 | 230 | 462 | 300 | 115 | 100 | 64 |
| 670.9 | 22.0 | 155 | 26.6 | 327 | 238 | 463 | 273 | 75 | 103 | 56 |
| 684.3 | 15.0 | 158 | 27.2 | 326 | 114 | 258 | 311 | 78 | 106 | 68 |
| 691.9 | 16.8 | 175 | 27.5 | 344 | 325 | 425 | 281 | 58 | 113 | 47 |
| 707.7 | 17.3 | 179 | 27.2 | 370 | 184 | 501 | 291 | 57 | 104 | 93 |
| 717.5 | 16.0 | 185 | 31.8 | 323 | 151 | 289 | 325 | 67 | 109 | 142 |
| 719.9 | 11.0 | 124 | 21.8 | 269 | 538 | 861 | 216 | 74 | 104 | 85 |
| 722.4 | 13.9 | 159 | 27.3 | 304 | 122 | 294 | 318 | 74 | 112 | 131 |
| 740.4 | 12.6 | 152 | 27.0 | 301 | 108 | 334 | 312 | 62 | 113 | 131 |
| 747.1 | 10.3 | 103 | 17.5 | 598 | 119 | 326 | 216 | 41 | 69 | 159 |
| 755.9 | 17.5 | 161 | 28.3 | 340 | 86 | 241 | 275 | 54 | 98 | 131 |
| 805.3 | 12.9 | 139 | 24.4 | 283 | 184 | 501 | 292 | 77 | 107 | 108 |
| 806.5 | 13.2 | 150 | 27.7 | 305 | 143 | 424 | 314 | 56 | 109 | 118 |
| 825.4 | 13.1 | 146 | 27.3 | 292 | 175 | 474 | 313 | 66 | 109 | 121 |
| $846.7{ }^{\text {a }}$ | 13.7 | 147 | 25.4 | 300 | 142 | 436 | 301 | 76 | 147 | 111 |
| 881.2 | 13.5 | 142 | 24.1 | 312 | 182 | 503 | 311 | 64 | 120 | 110 |
| 910.3 | 14.0 | 146 | 24.9 | 293 | 198 | 536 | 306 | 76 | 107 | 115 |
| $991.8^{\text {a }}$ | 14.6 | 173 | 30.7 | 335 | 99 | 161 | 327 | 51 | 117 | 132 |

Table 4 (continued)

| Depth (m) | Nb | Zr | Y | Sr | Ni | Cr | V | Co | Zn | Cu |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| 1036.3 $^{\mathrm{a}}$ | 13.8 | 161 | 26.9 | 314 | 114 | 310 | 319 | 68 | 109 | 124 |
| $1098.3^{\mathrm{a}}$ | 15.5 | 159 | 25.2 | 319 | 276 | 606 | 309 | 72 | 111 | 105 |
| $1217.7^{\mathrm{a}}$ | 16.9 | 187 | 29.5 | 334 | 134 | 262 | 334 | 65 | 116 | 117 |
| $1247.5^{\mathrm{a}}$ | 17.3 | 187 | 29.8 | 336 | 128 | 254 | 335 | 52 | 115 | 115 |
| $1291.1^{\mathrm{a}}$ | 17.3 | 189 | 30.0 | 337 | 120 | 230 | 343 | 55 | 116 | 115 |
| $1466.7^{\mathrm{a}}$ | 19.2 | 205 | 34.0 | 336 | 76 | 57 | 380 | 59 | 113 | 130 |
| 1550.5 | 18.5 | 173 | 27.4 | 382 | 146 | 318 | 339 | 50 | 103 | 135 |

${ }^{\text {a }}$ Sample is a dike.
rocks in a hydrothermally active area like the KERZ. Two geochemical parameters were used to evaluate the extent of alteration in the SOH 1 drill core: LOI and $\mathrm{K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}$ ratio. The LOI results show that the subaerial SOH 1 lavas experienced little or no alteration ( $>90 \%$ have LOI values $<0.35 \mathrm{wt}$. \%; Fig. 3), which is consistent with the thin section petrography results (Table 2). In contrast, there is a marked increase in alteration in the submarine section with LOI values up to $13.5 \mathrm{wt} \%$ just below the subaerial/ submarine transition, although LOI values range widely in the submarine section ( $0.3-20.7 \mathrm{wt} \%$; Fig. 3). The $\mathrm{K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}$ ratio of Hawaiian tholeiitic basalts is commonly used to evaluate their extent of weathering because K is leached during near surface alteration and P is not (e.g. Frey et al., 1994). The SOH 1 subaerial lavas have a restricted range in $\mathrm{K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}$ (1.4-2.0; Fig. 3), which is typical of Kilauea magmatic values (e.g. Wright, 1971). In contrast, rocks from the submarine section (lavas, hyaloclastites, and dikes) have a wide range in $\mathrm{K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}$ with both higher and lower values ( $0.6-9.3$; Fig. 3). The samples with ratios $>2$ have relatively high $\mathrm{K}_{2} \mathrm{O}$ contents and generally higher LOI values (up to $20.7 \mathrm{wt} \%$ ) but there is no correlation of LOI, rock type, or vesicularity with $\mathrm{K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}$. For the evaluation of Kilauea's magmatic history, samples with LOI $>2.25 \%, \mathrm{~K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}<1.3$ or $>2.2$, and alteration indices $>2.5$ were excluded, except for the two freshest samples from the lower part of the hole. These deleted samples represent $\sim 30 \%$ of the 80 analyzed SOH 1 samples and are all from the submarine section.

All of the SOH 1 lavas are shield-building tholeiites based on the Macdonald and Katsura (1964) classifi-
cation. For Hawaiian tholeiites, MgO is the most useful major element for evaluating compositional variation because olivine is the most common phenocryst. MgO contents in SOH 1 lavas range from 5.3 to $15.4 \mathrm{wt} . \%$, with the majority of values between 6 and $10 \mathrm{wt} . \%$ (Table 3). A similar range was reported for Hilina basalts (6-18 wt. \% MgO; Chen et al., 1996). There is no systematic variation of MgO with depth in SOH 1 (Fig. 4) but there is a general correlation between decreasing MgO content and increasing $\mathrm{TiO}_{2}, \mathrm{~K}_{2} \mathrm{O}, \mathrm{Al}_{2} \mathrm{O}_{3}$ (Fig. 4), $\mathrm{SiO}_{2}, \mathrm{Fe}_{2} \mathrm{O}_{3}$, and $\mathrm{P}_{2} \mathrm{O}_{5}$. The wide range in major element concentrations at any given $\mathrm{MgO} \mathrm{wt} . \%$, is indicative of a range of compositionally distinct parent magmas for SOH 1 lavas. $\mathrm{Al}_{2} \mathrm{O}_{3}$ and CaO contents decrease below $6.8 \mathrm{wt} . \% \mathrm{MgO}$, which is the distinction for differentiated Kilauea lavas (Wright 1971). These trends are indicative of plagioclase and clinopyroxene fractionation, which is consistent with both phases being phenocrysts in evolved SOH 1 rocks (Table 2). The increase in $\mathrm{Al}_{2} \mathrm{O}_{3} / \mathrm{CaO}$ ratio with decreasing MgO suggests that clinopyroxene is the dominant fractionating phase in these differentiated rocks (Fig. 4). Among the 20 analyzed a'a lavas, 17 are differentiated. Differentiated lavas are uncommon elsewhere on Kilauea and their abundance among the lavas in the lower east rift zone is an indication of the infrequent supply of magma to this part of the volcano (e.g. Wright, 1971).

Among the trace elements (Table 4), $\mathrm{Nb}, \mathrm{Zr}, \mathrm{Sr}, \mathrm{V}$ and Y behave incompatibly. Plots of these elements vs. K define linear trends with considerable scatter compared to analytical error (Fig. 5), which is probably related to multiple parental magmas. Ni and Cr are strongly compatible elements in SOH 1 lavas.


Fig. 3. Plots of alteration index, loss-on-ignition (LOI), MgO and $\mathrm{K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5} \mathrm{vs}$. depth for SOH 1 lavas. Alteration index is based petrographic examination of 80 thin sections (see Table 3 for details). Loss on ignition is wt. $\%$ loss after 8 h of heating at $900^{\circ} \mathrm{C}$. The alteration index and LOI values increase markedly at the submarine-subaerial transition, as does the range in $\mathrm{K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}$. The MgO plot only includes rocks meeting the alteration criteria ( $\mathrm{LOI}<2.25 \%, \mathrm{~K}_{2} \mathrm{O} / \mathrm{P}_{2} \mathrm{O}_{5}>1.3$ and $<2.2$, and alteration indices $<2.5$ ). There is no systematic variation of MgO with depth and most of the SOH 1 lavas contain $6-10 \mathrm{wt} . \% \mathrm{MgO}$.

When plotted against an incompatible element, these elements define trends with negative slopes and considerable scatter (Fig. 5), which may be the result of both magma mixing and crystal fractionation from multiple parent magmas. None of the trace elements show systematic variations with depth in the section (Fig. 5) nor is there a systematic change in trace element ratios with depth (Fig. 6).

## 6. Glass geochemistry

Glass compositions were measured at the University of Hawaii using a Cameca SX-50, five-spectrometer electron microprobe. Natural glass and mineral standards including VG2, A99, OR-1-5, Apatite and Troilite were used for this study. A PAP-ZAF correction was applied to all analyses and minor corrections


Fig. 4. MgO variation diagrams for SOH 1 lavas meeting alteration criteria (see Fig. 3 caption). Triangles represent samples recovered from $<300 \mathrm{~m}$ depth, squares $300-722 \mathrm{~m}$ depth, and circles $>722 \mathrm{~m}$ depth (open circles are for glasses from $>722 \mathrm{~m}$; Table 5). All oxides are in wt.\%. The kink in the variation trends at $6.5-7.0 \mathrm{wt} . \% \mathrm{MgO}$ marks the transition from olivine-only basalts to differentiated lavas. Error bars are smaller than symbols.
( $<2 \%$ ) were made, based on repeated analysis of Smithsonian glass standards A99 and VG2, to obtain the final values. A 10 nA , defocused beam (spot size $\sim 10 \mu \mathrm{~m}$ ) was used. Peak counting times of 70 s for $\mathrm{Fe}, \mathrm{Mg}, \mathrm{Ti}$, and $\mathrm{Si}, 40 \mathrm{~s}$ for $\mathrm{Ca}, \mathrm{Na}$, and K , and 100 s for $S$, Al, and $P$ were used with background counting times half of the peak times. The reported analyses are averages of five spot analyses per sample.

Glasses from 18 hyaloclastites were microprobed (Table 5) to determine compositional variation in the submarine section of the core where few massive basalts are available for whole rock analyses (7401123 m ). S analyses are useful for qualitatively determining the depth of eruption of the glasses (e.g.

Moore and Clague, 1992). Quenching temperatures for Kilauea glasses can be estimated using Helz and Thornber (1987) geothermometer. The glasses have a very limited range in MgO content (5.9-6.6 wt. \%, with $85 \%$ between 6.4 and $6.6 \mathrm{wt} . \%$ ), CaO content (10.3-11.2 wt.\%) and $\mathrm{K}_{2} \mathrm{O}$ (0.39-0.52) indicating that they were derived from similar parental magmas and were quenched at similar temperatures (1135$1145^{\circ} \mathrm{C}$ ). These glasses have the same compositional range as the submarine lavas (i.e. relatively low $\mathrm{K}_{2} \mathrm{O}$ compared to the younger subaerial lavas; Fig. 4). The S contents of these hyaloclastite glasses are low ( $0.01-0.07 \mathrm{wt} . \%$, with all but three $<0.04 \mathrm{wt} . \%$ ). The lower values are thought to be indicative of


Fig. 5. K variation diagrams for $\mathrm{Nb}, \mathrm{Zr}, \mathrm{Cr}$, and Ni for SOH 1 lavas meeting alteration criteria. All values are in ppm and symbols are the same as in Fig. 4. The incompatible elements form linear trends with scatter well beyond analytical error. The scatter for the compatible elements (Cr and Ni ) may be related to magma mixing; the scatter in the Nb and Zr data require distinct parental magmas. Two sigma error bars are given in lower right corner for each plot.

Table 5
Electron microprobe glass analyses of hyaloclastites from SOH 1 (analyses are in wt\%; total iron as $\mathrm{Fe}_{2} \mathrm{O}_{3}$; analysts; S. Quane and M. Garcia)

| Depth (m) | $\mathrm{SiO}_{2}$ | $\mathrm{TiO}_{2}$ | $\mathrm{Al}_{2} \mathrm{O}_{3}$ | $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | MgO | CaO | $\mathrm{Na}_{2} \mathrm{O}$ | $\mathrm{K}_{2} \mathrm{O}$ | $\mathrm{P}_{2} \mathrm{O}_{5}$ | S | Total |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 740 | 51.48 | 2.47 | 13.43 | 11.79 | 6.35 | 10.73 | 2.34 | 0.40 | 0.18 | 0.02 | 99.16 |
| 746 | 51.58 | 2.47 | 13.52 | 11.80 | 6.39 | 10.80 | 2.32 | 0.38 | 0.17 | 0.01 | 99.43 |
| 793 | 51.16 | 2.53 | 13.52 | 11.81 | 6.47 | 11.02 | 2.27 | 0.42 | 0.20 | 0.02 | 99.40 |
| 808 | 50.97 | 2.47 | 13.55 | 11.76 | 6.53 | 10.99 | 2.32 | 0.42 | 0.19 | 0.02 | 99.19 |
| 814 | 51.04 | 2.51 | 13.56 | 11.82 | 6.52 | 11.04 | 2.29 | 0.42 | 0.20 | 0.00 | 99.40 |
| 824 | 51.16 | 2.53 | 13.56 | 11.86 | 6.51 | 11.00 | 2.30 | 0.41 | 0.18 | 0.02 | 99.51 |
| 841 | 51.07 | 2.54 | 13.54 | 11.83 | 6.52 | 11.01 | 2.25 | 0.42 | 0.21 | 0.02 | 99.38 |
| 870 | 51.09 | 2.55 | 13.53 | 11.77 | 6.50 | 11.03 | 2.27 | 0.42 | 0.20 | 0.02 | 99.35 |
| 880 | 51.16 | 2.59 | 13.46 | 12.06 | 6.52 | 10.85 | 2.31 | 0.43 | 0.20 | 0.03 | 99.57 |
| 883 | 51.08 | 2.49 | 13.55 | 11.87 | 6.33 | 10.96 | 2.30 | 0.43 | 0.21 | 0.02 | 99.21 |
| 888 | 51.03 | 2.60 | 13.48 | 12.05 | 6.32 | 10.93 | 2.31 | 0.43 | 0.22 | 0.02 | 99.37 |
| 893 | 50.89 | 2.81 | 13.15 | 12.84 | 5.88 | 10.19 | 2.34 | 0.44 | 0.22 | 0.07 | 98.76 |
| 911 | 51.08 | 2.59 | 13.57 | 11.98 | 6.40 | 10.95 | 2.28 | 0.43 | 0.21 | 0.02 | 99.49 |
| 919 | 51.00 | 2.56 | 13.53 | 11.97 | 6.39 | 11.05 | 2.30 | 0.42 | 0.20 | 0.02 | 99.42 |
| 927 | 50.23 | 2.73 | 13.67 | 12.14 | 6.36 | 10.73 | 2.24 | 0.44 | 0.21 | 0.06 | 98.73 |
| 934 | 50.52 | 2.77 | 13.57 | 12.34 | 5.95 | 10.53 | 2.38 | 0.52 | 0.22 | 0.04 | 98.80 |
| 998 | 51.05 | 2.81 | 13.21 | 12.42 | 6.18 | 10.64 | 2.34 | 0.45 | 0.22 | 0.01 | 99.33 |
| 1123 | 51.03 | 2.62 | 13.42 | 12.01 | 6.49 | 10.98 | 2.27 | 0.42 | 0.22 | 0.02 | 99.45 |



Fig. 6. $\mathrm{Zr} / \mathrm{Nb}$ vs. depth and $\mathrm{Sr} / \mathrm{Y}$ for SOH 1 lavas meeting alteration criteria. Only lavas with $\mathrm{MgO}>7 \mathrm{wt} . \%$ are used for $\mathrm{Sr} / \mathrm{Y}$ plot to avoid the affects of plagioclase fractionation on Sr . There is no systematic variation for these immobile trace element ratios with depth. There is a wide range in composition for each depth range, as observed for historical Kilauea lavas (Pietruszka and Garcia, 1999). Two sigma error bars are shown in lower right (depth plot) and upper right (Sr/Y plot) corners.
subaerial eruption; the higher values are indicative of shallow submarine ( $<500 \mathrm{~m}$ ) eruption (see Moore and Clague, 1992; Garcia, 1993).

## 7. Temporal geochemical variation of Kilauea

When we began this study of rock cores from a 1.7 km deep drill hole, we anticipated finding lavas from either the preshield stage of Kilauea and/or from neighboring Mauna Loa Volcano. To our surprise,
neither was encountered. The SOH 1 core contains no rocks with preshield stage compositions (i.e. lower $\mathrm{SiO}_{2}$, higher $\mathrm{Na}_{2} \mathrm{O}+\mathrm{K}_{2} \mathrm{O}$ tholeiites, transitional, or alkalic lavas, which typify Loihi Volcano, the type example of preshield stage Hawaiian volcanism; e.g. Moore et al., 1982). Also, there is no indication of a change in geochemistry with depth towards pre-shield stage compositions. Furthermore, all of the SOH 1 lavas have Kilauea-like compositions, which are distinct from those of neighboring Mauna Loa Volcano. For example, SOH 1 lavas have relatively high concentrations of the immobile incompatible elements Ti and Nb , a characteristic difference between the lavas from the two volcanoes (Fig. 7). It should be noted that there are a few unpublished analyses of Mauna Loa submarine southwest rift lavas (J.M. Rhodes, 1998, pers. comm.) with Kilauea-like major and trace element compositions but they are interbedded with and subordinate ( $\sim 10 \%$ ) to lavas having typical Mauna Loa compositions (Fig. 8). Although some Mauna Loa rocks with Kilauea-like compositions could be present in the SOH 1 core, it is unlikely due to the absence of any lavas with Mauna Loa-like compositions within the core.

Overall, the SOH 1 rocks are similar in major element chemistry to both the oldest exposed Kilauea surface lavas, the prehistoric Hilina basalts, and to the youngest Kilauea lavas from the ongoing Puu Oo eruption (Chen et al., 1996; Garcia et al., 1996). The only trend noted for major element variation with depth for the SOH 1 lavas is that the submarine rocks have lower $\mathrm{K}_{2} \mathrm{O}$ than the upper subaerial rocks ( $<300 \mathrm{~m}$; Fig. 4), although the intermediate depth rocks overlap both groups with no depth trend. Hilina Basalts have lost K due to weathering (Chen et al., 1996), so we cannot compare their $K$ variation with the SOH 1 results. No other long-term systematic geochemical variation is apparent for the SOH 1 section (Figs. 4-5), which is consistent with the lack of a temporal compositional variation for the Hilina Basalts (Chen et al., 1996). Compared to the Hilina Basalts and historical Kilauea summit lavas ( $<210$ years; Pietruszka and Garcia, 1999), SOH 1 lavas display only slightly more trace element and less major element variation (Fig. 8), despite the much greater age range represented by the SOH 1 samples.

Recent studies of Hawaiian volcanoes (e.g. Frey


Fig. 7. MgO variation diagrams for Nb and $\mathrm{TiO}_{2}$ in SOH 1 lavas meeting alteration criteria. Nb and $\mathrm{TiO}_{2}$ define distinct fields for historical Mauna Loa ( $<160$ years old, field bounded by solid line; Rhodes and Hart, 1995) and Kilauea summit lavas (field bounded by dashed line; Pietruszka and Garcia, 1999; Garcia unpub. data). The SOH 1 lavas plot well above the Mauna Loa field and mostly within the narrow field for Kilauea historical lavas $(<210$ years old). Symbols as in Fig. 4; errors are about the size of the symbols.
and Rhodes, 1993; Frey et al., 1994; Roden et al., 1994; Kurz et al., 1995; Rhodes and Hart, 1995; Chen et al., 1996) have shown that Sr and Pb isotopes are well correlated with incompatible trace element ratios involving Nb . Thus, these trace element ratios reflect different source components for Hawaiian lavas (Rhodes, 1996). Rhodes et al. (1989) also found that $\mathrm{TiO}_{2} / \mathrm{Al}_{2} \mathrm{O}_{3}$, which is sensitive to partial melting variations, distinguishes Kilauea from Mauna Loa lavas. SOH 1 lavas have Kilauea-like $\mathrm{Zr} / \mathrm{Nb}$ and $\mathrm{TiO}_{2} / \mathrm{Al}_{2} \mathrm{O}_{3}$ (Fig. 8). Hence, Kilauea's source and melting conditions may have remained essentially the same over the last 350 ka .

## 8. Age of Kilauea Volcano

Kilauea Volcano is one of the world's most active volcanoes and is presently in its shield-building stage. This stage is thought to be comprised entirely of tholeiitic rocks and to last 500 ka (Moore and Clague, 1992) to 900 ka (Guillou et al., 1997a). Little was known prior to the SOH drilling program study about how long Kilauea had been in the shield building stage because of the rapid resurfacing rate of Kilauea (Holcomb, 1987) and the absence of geochronological ages $>23 \mathrm{ka}$ (Easton, 1987). The unspiked $\mathrm{K}-\mathrm{Ar}$ dates and rock chemistry from SOH 1 (Tables 1 and 3) show that Kilauea has erupted shield stage


Fig. 8. $\mathrm{TiO}_{2} / \mathrm{Al}_{2} \mathrm{O}_{3}$ vs. $\mathrm{Zr} / \mathrm{Nb}$ plot for SOH 1 lavas (symbols as in Fig. 4), Kilauea summit historical lavas (field bounded by solid line; Garcia unpublished), Hilina Basalts (long dashed lines; Chen et al., 1996), Mauna Loa submarine lavas (field bounded by dashed line; Rhodes unpub. data), and Mauna Loa historical lavas (field bounded by solid line; Rhodes and Hart, 1995). Rhodes et al. (1989) demonstrated that this plot was effected in discriminating historical lavas of Mauna Loa from those of Kilauea. The SOH 1 data plot mostly within the field for historical Kilauea lavas but partially overlap with the rare older submarine Mauna Loa lavas with Kilauea-like compositions.
tholeiites for at least $\sim 350 \mathrm{ka}$. This relatively old age is consistent with the $436 \pm 47 \mathrm{ka}$ age for a tholeiitic lava from SOH 4 (Guillou et al., 1997b) from a somewhat greater depth. There is geochemical evidence that the old SOH 4 lava is from a section of rocks with Mauna Loa compositions (i.e. higher ${ }^{87} \mathrm{Sr} /{ }^{86} \mathrm{Sr}$ and $\mathrm{Zr} / \mathrm{Nb}$ ratios and lower ${ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}$ ratios than typical Kilauea rocks; H. West, 1998 unpublished data). However, the presence of rocks with Kilauealike geochemistry $\sim 210 \mathrm{~m}$ above the dated sample suggests that Kilauea lavas in SOH 4 are probably $\sim 400 \mathrm{ka}$.

If Kilauea had a pre-shield stage, which is believed to occur in all Hawaiian volcanoes (Walker, 1990), and the duration of this stage was similar to Loihi's ( $\sim 250 \mathrm{ka}$; Guillou et al., 1997a), then Kilauea's age may be $\sim 600-650 \mathrm{ka}$. This age is markedly greater than previous estimates for the age of Kilauea based on magma supply rates and surface exposures (e.g. 150 ka, Lipman, 1995; 200-300 ka, Easton, 1978). It is, however, in agreement with the $\sim 600 \mathrm{ka}$ age predicted for Kilauea based on lava accumulation rates from the Hawaiian Scientific Drilling Project pilot hole (DePaolo and Stolper, 1996). Their volcano growth model also predicts that Kilauea is $\sim 60 \%$ of the way through its shield building stage (DePaolo
and Stolper, 1996). If this estimate is correct, and Kilauea follows the general evolutionary model for Hawaiian shield volcanoes with a pre-shield ( $\sim 250 \mathrm{ka}$ for Loihi; Guillou et al., 1997a) and a post-shield ( $\sim 240 \mathrm{ka}$ for Mauna Kea; Frey et al., 1990), the entire life span for Kilauea would be $\sim 1150 \mathrm{ka}$. This estimate is consistent with recent estimates for the life span of Hawaiian volcanoes ( $\sim 1000 \mathrm{ka}$, Lipman, 1995; $\sim 1400 \mathrm{ka}$, Guillou et al., 1997a).

## 9. Lava accumulation rates for Kilauea's east rift zone

Lava accumulation rates for Kilauea's east rift zone can be estimated using the new $\mathrm{K}-\mathrm{Ar}$ ages for SOH 1 . The oldest sample gives an overall growth rate of $\sim 4.4 \mathrm{~mm} /$ year. If the other ages are considered individually, the rate increases upsection (5.4, 27 and $81 \mathrm{~mm} / \mathrm{year}$ ). Similar but less dramatic increases in growth rate are evident for the nearby SOH 4 drill hole ( $3.8-11 \mathrm{~mm} /$ year; Guillou et al., 1997b). Combining the ages from these two holes, which are only $\sim 8 \mathrm{~km}$ apart, a trend of progressively increasing lava accumulation emerges (curved trend in Fig. 9).


Fig. 9. Lava accumulation rates for Kilauea's lower east rift zone based on ages from SOH 1 (triangles) and SOH 4 (squares) lavas. The $4.4 \mathrm{~mm} /$ year line is a linear accumulation rate based only on the oldest dated SOH lava and ignores the presence of dikes (most abundant below 1200 m in SOH 1 and below 1000 m in SOH 4 ). If the ages from two drill holes are combined, a pattern of increasing lava accumulation rate emerges even if the older lavas are ignored (because of the presence of dikes of unknown but younger age). The lavas from above the intense dike zones in these two holes to 500 m may have accumulated at $\sim 4 \mathrm{~mm} /$ year but the younger lavas seem to have accumulated more rapidly ( $\sim 10 \mathrm{~mm} /$ year). The two younger SOH 1 lavas have ages that seem too young for their position in the section.

This trend is somewhat muted by the contribution to the deeper section of these holes ( $>1200 \mathrm{~m}$ for SOH 1 and $>1000 \mathrm{~m}$ for SOH 4 ) from dikes. If the dikes in these two holes are removed, the lava accumulation rate trend upsection increases more dramatically from rates of $1-2 \mathrm{~mm} /$ year near the base to $\sim 10 \mathrm{~mm} /$ year for the upper section of the holes.

A lava accumulation rate of $10 \mathrm{~mm} /$ year for upper SOH lavas is somewhat higher than any previously calculated rate for the flanks of Hawaiian volcanoes (e.g. $7.8 \pm 3.2 \mathrm{~mm} /$ year for Mauna Kea; Sharp et al., 1996; $4.3 \mathrm{~mm} /$ year Mauna Loa; Lipman and Moore, 1996) but much lower than the rates estimated for the summits of Hawaiian volcanoes ( $25-35 \mathrm{~mm} /$ year; DePaolo and Stolper, 1996). The relatively high lava accumulation rate determined here for the east rift zone may be indicative of the dominance of this rift zone over Kilauea's other rift zone ( $>90 \%$ of historically erupted rift zone lavas; Macdonald et al., 1983). Likewise, the cause for the increase in lava accumula-
tion rate in the SOH sections is probably related to the development of east rift zone as a major locus of Kilauea's volcanism rather than to an increase in its magma supply rate (see below).

## 10. Long-term magma supply rates for Kilauea

Determining the magma supply rate for any volcano is inherently problematic. For example, estimates for Kilauea's magma supply rate based on summit tiltmeter and leveling surveys utilize simplistic point source models for inflation of its summit reservoir and ignore passive movement of magma into rift zones (e.g. Dzurisin et al., 1984). For periods prior to tiltmeters and leveling surveys, eruption volumes are used to infer magma supply rates. Periods of sustained eruptions, such as the current eruption of Kilauea, are thought to give more reliable estimates (e.g. Dvorak and Dzurisin 1993), although they must also be considered minimum rates because of
endogenous growth of the volcano (e.g. Francis et al., 1993). These methods indicate wide variations in the magma supply rate ( $0.01-0.12 \mathrm{~km}^{3} /$ year; Pietruszka and Garcia, 1999) for Kilauea's historical period (1790-1998).

To avoid many of the problems in estimating Kilauea's historical magma supply rate, we have taken published volume estimates for the volcano (15-22.5 $\times 10^{3} \mathrm{~km}^{3}$; Bargar and Jackson, 1974; Lipman 1995; DePaolo and Stolper, 1996) and our estimated age for Kilauea shield volcanism ( $\sim 400 \mathrm{ka}$ ) to calculate a 'long-term' magma supply rate. A correction for Kilauea's preshield stage is not needed because it would be minor compared to the difference in the volume estimates for the volcano ( $10 \%$, assuming a volume similar to Loihi's preshield stage, $0.75 \times 10^{3} \mathrm{~km}^{3}$; Garcia et al., 1995). Dividing the volume estimates by the age yields a long-term magma supply rate of $0.05 \pm 0.01 \mathrm{~km}^{3} /$ year for Kilauea. This value is remarkably similar to the estimated average historical rate of $0.06 \mathrm{~km}^{3} /$ year (Dvorak and Dzurisin, 1993). If magma compositions are primarily controlled by the extent of partial melting (e.g. Watson and McKenzie, 1991), then the narrow range of Kilauea lava compositions for the last 350 ka supports the interpretation that the volcano's overall magma supply rate probably has remained relatively constant during this period.

## 11. Implications for the structure of the Hawaiian plume

Hawaii has played a central role in understanding mantle plumes because so much is known about its volcanoes and underlying plate compared to other plumes. For example, Hawaiian volcanoes show a systematic compositional variation during the later stages of shield (e.g. Mauna Loa; Kurz et al., 1995) and post-shield volcanism (e.g. Haleakala; Chen et al., 1991), which is thought to be related to compositional heterogeneity in the Hawaiian plume (e.g. Lassiter et al., 1996). The age of the Pacific plate under Hawaii is $\sim 105 \mathrm{Ma}$ (Waggoner, 1993) and recent estimates have it moving towards the northwest at $\sim 12 \mathrm{~cm} /$ year (Wessel and Kroneke, 1997). It is assumed that the Hawaiian plume is stationary compared to the motion of the Pacific plate. These observations and
assumptions have been used to infer a radially zoned structure for the Hawaiian plume with a $\sim 100 \mathrm{~km}$ wide melt extraction zone (Watson and McKenzie, 1991; DePaolo and Stolper, 1996) and a $\sim 50 \mathrm{~km}$ wide central core that feeds shield-stage volcanism (e.g. Lassiter et al., 1996). Previous studies have suggested that Kilauea is located on the margin of this central core (e.g. Lassiter et al., 1996; Hauri et al., 1996). The geochronological and geochemical results for SOH 1 lavas allow us to evaluate these models for the structure and size of the Hawaiian plume.

If the plume remains stationary relative to the Pacific plate, the oldest reliable age for SOH 1 lavas ( 351 ka ) and the velocity of Pacific plate motion for Hawaii ( $12 \mathrm{~cm} / \mathrm{year}$; Wessel and Kroneke, 1997) indicate that Kilauea Volcano may have sampled a 42 km long section of the Hawaiian plume (or 48 km , if Kilauea's shield stage volcanism has been active for 400 ka ). During this period, the composition of Kilauea magmas remained remarkably consistent based on SOH 1 lava compositions (Tables 3 and 4). Therefore, the portion of the plume supplying Kilauea must be well mixed (i.e. shows no sign of the largescale heterogeneity that characterizes later stages of Hawaiian volcano growth; Chen et al., 1991) and continually changing to allow a "fresh" portion of the plume to be melted. Neighboring Mauna Loa Volcano, which erupts geochemically distinct lavas (Fig. 7), is only $\sim 35 \mathrm{~km}$ from Kilauea. Thus, the area of the plume that has been melted to form Kilauea is $<35 \mathrm{~km}$ wide (perhaps $\sim 30 \mathrm{~km}$ ) and $\sim 48 \mathrm{~km}$ long, or $\sim 1130 \mathrm{~km}^{2}$. If the volume of Kilauea is $\sim 20,000 \mathrm{~km}^{3}$ and its lavas are produced by $5-10 \%$ melting (Watson and McKenzie, 1991; Pietruszka and Garcia, 1999), then $\sim 150,000 \mathrm{~km}^{3}$ of the plume was melted to form Kilauea. Therefore, a $\sim 130 \mathrm{~km}$ thick region of the plume would have been melted to form Kilauea, which is considerably thicker than previous estimates of the melting thickness in the Hawaiian plume ( $\sim 55 \mathrm{~km}$; Watson and McKenzie, 1991; ~70 km; Ribe and Christensen, 1999). Either these models underestimated the thickness of the zone of melting within the plume, or the core of the plume is not stationary.

Kilauea is estimated to be $\sim 60 \%$ of the way through its shield-building stage (DePaolo and Stolper, 1996). Therefore, the plume core diameter may
be $70-80 \mathrm{~km}$ where Kilauea passes over it. This interpretation assumes that the mantle magma conduit is not deformed during the life of the volcano. Unfortunately, there are too few earthquakes under Mauna Loa, which is near the end of its shield stage (Moore and Clague, 1992), to allow the orientation of its deep mantle conduit to be determined (Klein et al., 1987). If Kilauea has preshield and postshield stages of similar duration to those of adjacent volcanoes ( $\sim 250$ ka each), then a $\sim 140 \mathrm{~km}$-wide zone of melt extraction might be needed for Kilauea. This estimate is larger than previous estimates of $\sim 100 \mathrm{~km}$ for the width of the primary melt extraction region of the Hawaiian plume but is smaller than a new estimate of $\sim 180 \mathrm{~km}$ based on a dynamic model for plume melting (Ribe and Christensen, 1999).

## 12. Conclusions

The lavas from the $\sim 1.7 \mathrm{~km}$ deep SOH 1 drill hole are virtually identical in mineralogy and composition to historical Kilauea lavas. The new geochronological data for SOH 1 confirm previous dating of drill core from another Kilauea east rift zone drill hole (SOH 4; Guillou et al. 1997b) demonstrating that Kilauea has been producing shield-building stage tholeiitic lavas for at least 350 ka , probably 400 ka . If Kilauea had a preshield stage of similar duration to Loihi Volcano ( $\sim 250 \mathrm{ka}$; Guillou et al. 1997a), its overall age would be at least 600 ka , which is much older than previous field-based studies but consistent with a theoretical model for its age (DePaolo and Stolper, 1996). Our estimate of the average long-term ( $\sim 350 \mathrm{ka}$ ) magma supply rate for Kilauea's shield building is nearly identical to a recent estimate for Kilauea's historical magma supply rate ( $0.05 \pm 0.01$ vs. $0.06 \mathrm{~km}^{3} / \mathrm{year}$; Dvorak and Dzurisin, 1993). The overall consistency in the geochemistry of SOH 1 lavas and of Kilauea's magma supply rate suggests neither the melting conditions nor the source for Kilauea lavas has changed significantly during the past 350 ka .

The age and geochemistry of the SOH lavas can be used to infer the size of the Hawaiian plume based on a few observations and assumptions. If the Pacific plate has been moving $12 \mathrm{~cm} /$ year for the last few million years (Wessel and Kroneke, 1997) and the plume is stationary relative to the motion of the Paci-
fic plate, SOH 1 lavas sampled $\sim 42 \mathrm{~km}$ of the core of the Hawaiian plume ( 48 km , if the 400 ka age is used). During its shield stage, the composition of Kilauea magmas remained remarkably consistent based on SOH 1 lava compositions (Tables 3 and 4). Therefore, the portion of the plume supplying Kilauea must be well mixed (i.e. shows no sign of the large-scale heterogeneity that characterizes later stages of Hawaiian volcano growth; Chen et al., 1991) and continually changing to allow a "fresh" portion of the plume to be melted. Thus, the simplistic cartoons for the structure of the Hawaiian plume which place Kilauea near the edge of a radially zoned plume (e.g. Lassiter et al., 1996) need modification to accommodate a wellmixed source for at least the last 350 ka .
If the plume remains stationary relative to the Pacific plate, then the portion of the Hawaiian plume that has been melted to form Kilauea is about 48 km long, 30 km wide and 130 km deep, which is much greater than previous estimates. Although the location of the plume may be stationary, the mass within the plume may move faster than the Pacific plate and replenish the region that is producing Kilauea magmas.

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